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PART B

SOLAR - GEOPHYSICAL DATA

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U. S. DEPARTMENT OF COMMERCE NATIONAL BUREAU OF STANDARDS CENTRAL RADIO PROPAGATION LABORATORY BOULDER, COLORADO

SOLAR - GEOPHYSICAL DATA

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SOLAR - GEOPHYSICAL DATA

INTRODUCTION

This monthly report series is intended to keep research workers abreast of the major particulars of solar activity and the associated ionospheric, radio propagation and other geophysical effects. It is made possible through the cooperation of many observatories, laboratories and agencies as recorded in the detailed description of the tables and graphs which follows. The report is prepared in the Radio Warning Services Section, edited by Miss J.V. Lincoln.

I DAILY SOLAR INDICES

<u>Relative Sunspot Numbers</u> -- The table includes (1) the daily American relative sunspot numbers, R_A ', as compiled by the Solar Division of the American Association of Variable Star Observers, and (2) the provisional daily Zurich relative sunspot numbers, R_Z , as communicated by the Swiss Federal Observatory. Because of the time required to collect and reduce the observations, R_A ' will normally appear one month later than R_Z .

The relative sunspot number is an index of the activity of the entire visible disk. It is determined each day without reference to preceding days. Each isolated cluster of sunspots is termed a sunspot group and it may consist of one or a large number of distinct spots whose size can range from 10 or more square degrees of the solar surface down to the limit of resolution (e.g. 1/8 square degrees). The relative sunspot number is defined as R=K(10g+s), where g is the number of sunspot groups and s is the total number of distinct spots. The scale factor K (usually less than unity) depends on the observer and is intended to effect the conversion to the scale originated by Wolf. The observations for sunspot numbers are made by a rather small group of extraordinarily faithful observers, many of them amateurs, each with many years of experience. The counts are made visually with small, suitably protected telescopes.

Final values of R_Z appear in the IAU <u>Quarterly Bulletin on Solar</u> <u>Activity</u>, the <u>Journal of Geophysical Research</u>, these reports, and elsewhere. They usually differ slightly from the provisional values. The American numbers, R_A , are not revised.

<u>Solar Flux Values, 2800 Mc</u> -- The table also lists the daily values of solar flux at 2800 Mc recorded in watts/ M^2 /cycle/second bandwidth (x 10⁻²²) in two polarizations by the National Research Council at Ottawa, Canada. These solar radio noise indices are being published in accordance with CCIR Report 25 that a basic solar index for ionospheric propagation should be measured objectively and "preferably refer to a property of the sun such as radiation flux which has direct physical relationship to the ionosphere." <u>Graph of Sunspot Cycle</u> -- The graph illustrates the recent trend of Cycle 19 of the 11-year sunspot cycle and some predictions of the future level of activity. The customary "12-month" smoothed index, R, is used throughout, the data being final RZ numbers except for the current year. Predictions shown are those made for one year after the latest available datum by the method of A. G. McNish and J. V. Lincoln (Trans. Am. Geophys. Union, <u>30</u>, 673-685, 1949) modified by the use of regression coefficients and mean cycle values recomputed for Cycles 8 through 18. Cycle 19 began April 1954, when the minimum \overline{R} of 3.4 was reached.

II SOLAR CENTERS ()F ACTIVITY

<u>Calcium Plage and Sunspot Regions</u> -- The table gives particulars of the centers of activity visible on the solar disk during the preceding month. These are based on estimates made and reported on the day of observation and are therefore of limited reliability.

The table gives the heliographic coordinates of each center (taken as the calcium plage unless two or more significantly and individually active sunspot groups are included in an extended plage) in terms of the Greenwich date of passage of the sun's central meridian (CMP) and the latitude; the serial number of the plage as assigned by McMath-Hulbert Observatory; the serial number of the center in the previous solar rotation, if it is a persisting region; particulars of the plage at CMP: area, central intensity; a summary of the development of the plage during the current transit of the disk, where b = born on disk, ℓ = passed to or from invisible hemisphere, d = died on disk, and / = increasing, - = stable, $\setminus =$ decreasing; and age in solar rotations; particulars of the associated sunspot group, if any, at CMP: area and spot count and the summary of development during the current disk transit, similar to the above. The unit of area is a millionth of the area of a solar hemisphere; the central intensity of calcium plages is roughly estimated on a scale of 1 = faint to 5 = very bright.

Calcium plage data are available through the cooperation of the McMath-Hulbert Observatory of the University of Michigan. The sunspot data are compiled from reports from the U. S. Naval Observatory, and from reports from Europe and Japan received through the daily Ursigram messages.

<u>Coronal Line Emission Indices</u> -- In the table are summarized solar coronal emission intensity indices for the green (Fe XIV at λ 5303) and red (Fe X at λ 6374) coronal lines. The indices are based on measurements made at 5° intervals around the periphery of the solar disk by the High Altitude Observatory at Climax, Colorado, and by Harvard University observers at Sacramento Peak (The USAF Upper Air Research Observatory at Sunspot, New Mexico, under contract AF 19(604)-146). The measurements are expressed as the number of millionths of an Angstrom of the continuum of the center of the solar disk (at the same wavelength as the line) that would contain the same energy as the observed coronal line. The indices have the following meanings:

- $R_{\rm A}$ = same for $\lambda 6374$.
- G_1 = highest value of intensity in quadrant, for λ 5303.
- $R_1 = \text{same for } \lambda 6374.$

The dates given in the table correspond to the approximate time of CMP of the longitude zone represented by the indices. The actual observations were made for the North East and South East quadrants 7 days before; for the South West and North West quadrants 7 days after the CMP date given.

To obtain rough measures of the integrated emission of the entire solar disk in either of the lines, assuming the coronal changes to be small in a half solar rotation, it is satisfactory to perform the following type of summation given in example for 15 October:

$$(\overset{\text{MEAN DISK EMISSION}}{\underset{\text{IN }\lambda \text{ 5303}}{\text{ 3503}}})_{\text{IS OCT}} = \frac{1}{N} \left[\sum_{\text{IS OCT}}^{22 \text{ OCT}} \left\{ \left(G_6 \right)_{\text{NE}} + \left(G_6 \right)_{\text{SE}} \right\} + \sum_{\text{S OCT}}^{14 \text{ OCT}} \left\{ \left(G_6 \right)_{\text{SW}} + \left(G_6 \right)_{\text{NW}} \right\} \right]$$

where N is the number of indices entering the summation.

Such integrated disk indices as well as integrated whole-sun indices are computed for each day and are published quarterly in the "Solar Activity Summary" issued by the High Altitude Observatory at Boulder, Colorado. In the same reports are given maps of the intensity distribution of coronal emission derived from all available Climax and Sacramento Peak observations, as well as other information on solar activity, such as maps made from daily limb prominence surveys in H α and notes regarding the history of active regions on the solar disk.

Preliminary summaries of solar activity, prepared on a fast schedule, are issued Friday of each week from High Altitude Observatory in conjunction with CRPL and include solar activity through the preceding day. These are useful to groups needing information on the current status of activity on the visible solar disk, but are not recommended for research uses unless such a prompt schedule of reporting is essential. The same information is included in the subsequent quarterly reports, with extensive additions, corrections and evaluations.

 G_6 = mean of six highest line intensities in guadrant for $\lambda 5303$.

III SOLAR FLARES

Optical Observations -- The table presents the preliminary record of solar flares as reported to the CRPL on a rapid schedule at the sacrifice of detailed accuracy. Definitive and complete data are published later in the Quarterly Bulletin on Solar Activity, I.A.U., in various observatory publications and elsewhere. The present listing serves to identify and roughly describe the phenomena observed.

Reporting directly to the CRPL are the following observatories: McMath-Hulbert, Wendelstein, Sacramento Peak, Mitaka and Swedish Astrophysical Station on Capri. The remainder report through the URSIgram centers or are available through the IGY World Data Center for Solar Activity in Boulder. Observations are in the light of the center of the II-alpha line unless noted otherwise. The reports from Sacramento Peak, New Mexico (communicated to CRPL by the High Altitude Observatory at Boulder) are from observations at the USAF Upper Air Research Observatory at Sunspot, New Mexico, by Harvard University observers, under contract AF 19(604)-4961.

For each flare are listed the reporting observatory, the date, beginning and ending times, time of maximum phase, the heliographic coordinates in degrees, McMath serial number of the region, duration, the flare importance on the IAU scale of 1- to 3+, observing conditions where 1 means poor, 2 fair and 3 good, time of measurement for tabulated width of H α or tabulated area, measured (i.e. projected) maximum area in square degrees, corrected maximum area in square degrees which equals measured area times secant h where h is the heliocentric angle, maximum effective line-width in H α expressed in Angstroms, and maximum intensity of H α expressed in per cent of the continuous spectrum. The following symbols are used in the table:

D	=	Greater than	F	=	Approximately
E	=	Less than	3	=	Plus

A final column lists provisionally the occurrence of simultaneous ionospheric effects as observed on selected field-strength recordings of distant high-frequency radio transmissions; a more nearly definitive list of these ionospheric effects, including particulars, appears in these reports after the lapse of a month (see below). All times are Universal Time (UT or GCT). Subflares (importance 1-) are listed by date, time of beginning and their heliographic coordinates. A graph presents intervals for which there were no patrols for flare observations from the observatories whose complete data are published in the table.

Ionospheric Effects -- SID, sudden ionospheric disturbances (and GID--gradual ionospheric disturbances) may be detected in a number of ways: short wave fadeouts (SWF), enhancement of low frequency atmospherics (SEA), increases in cosmic absorption (SCNA), and so forth.

A table lists SWF events that have been recognized on fieldstrength recordings of distant high-frequency radio transmissions. Under a coordinated program, the staffs at the following ionospheric sounding stations contribute reports that are screened and synthesized at CRPL-Boulder: Puerto Rico, Ft. Belvoir, Va., and Anchorage, Alaska (CRPL Stations: PR, BE, AN); Huancayo, Peru (CRPL-Associated Laboratory: HU); and Ft. Monmouth, N.J., White Sands, N. Mex., Adak, Alaska, and Okinawa (U.S. Signal Corps Stations: FM, WS, AD, OK). McMath-Hulbert Observatory (MC) also contributes such reports. In addition, reports are volunteered by RCA Communications Inc., Marconi Wireless, Netherlands Postal and Telecommunications Services, Swedish Telecommunications, and others; these usually specify times of SWF and the radio paths involved. Through the URSIgrams, reports are available from still other stations as given monthly in the footnotes.

In the coordinated program, the abnormal fades of field strength not obviously ascribable to other causes, are described as short wave fadeouts with the following further classification:

S-SWF: sudden drop-out and gradual recovery Slow S-SWF: drop-out taking 5 to 15 minutes and gradual recovery G-SWF: gradual disturbance; fade irregular in either drop-out or recovery or both.

When there is agreement among the various reporting stations on the time (UT) of an event, it is accepted as a widespread phenomenon and listed in the table.

The degree of confidence in identifying the event, a subjective estimate, is reported by the stations and this is summarized in an index of certainty that the event is widespread, ranging from 1 (possible) to 5 (definite). The times given in the table for the event are from the report of a station (underlined in table) that identified it with high confidence. The criteria for the subjective importance rating assigned by each station on a scale of 1- to 3+ include amplitude of the fade, duration and confidence; greater consideration is given to reports on paths near the subsolar point in arriving at the summary importance rating given in the table. Note: The tables of SID observed at Washington included in CRPL Freports prior to F-135 were restricted to events classed here as S-SWF.

A second table lists sudden ionospheric disturbances which have been recognized on recorders for detecting cosmic absorption at about 18 Mc (SCNA) or on recorders for detecting enhancements of low frequency atmospherics at about 27 kc (SEA) together with solar radio bursts at 18 Mc as identified on the SCNA records.

Reports are received either directly or through the IGY World Data Center for Solar Activity at the High Altitude Observatory, Boulder, Colc. The following observatories report SCNA: Rensselaer Polytechnic Institute Observatory, Grafton, N.Y. (RE); McMath-Hulbert

Observatory (MC); Sacramento Peak, N.Mex. (SP); High Altitude Observatory, Boulder, Colo. (BO); University of Hawaii, Makapuu Pt., Hawaii (HA); and the Royal Observatory Edinburgh (ED). All of these except the Royal Observatory Edinburgh also report solar noise bursts observed at 18 Mc. The SEA reports come from the following: Department of Terrestrial Magnetism, Carnegie Institution of Washington, Station at Derwood, Md. (DE); Dunsink Observatory, Ireland (DU); Royal Observatory Edinburgh (ED); three stations operated by the Netherlands PTT at Hollandia, Dutch West Indies (HO), Nederhorst den Berg, Netherland (NE), and Paramaribo, New Guinea (PA); Panska Ves Observatory near Prague, Czech. (PU); High Altitude Observatory, Boulder, Colo. (BO); Sacramento Peak, N. Mex. (SP); McMath-Hulbert Observatory (MC); University of Hawaii (HA): Neustrelitz (NU): Kuhlungsborn (KU); and a group of American Association of Variable Star Observers located at Brooklyn. N.Y. (A1), Pittsburgh, Pa. (A2), Paterson, N.J. (A3), Powell, Ohio (A4), Ramsey, N.J. (A5), Oshkosh, Wis. (A6), China Lake, Calif. (A7), and Manhattan, Kansas (A8).

These reports are coordinated at CRPL-Boulder. When there is agreement among the various reporting stations on the time (UT) of an event, it is accepted as a widespread phenomenon and listed in the table. Some phenomena are listed, if noted at only one location, if there has been a flare or another type of flare-associated effect reported for that time.

In the table under the type of event the importance of the event is given on a scale of 1 minus to 3 plus. Next there is the index of widespread certainty ranging from 1 (possible) to 5 (definite). The time of beginning, maximum and end of the event in UT is given as reported by the station underlined in the group of observing stations. If the event is an SCNA, a percent absorption figure is given. This absorption is calculated by

SCNA % =
$$\frac{I_n - I_f}{I_n}$$
 x 100

- where I_n = noise diode current required to give a recorder deflection equal to that which would have occurred in the absence of a flare, i.e. a value extrapolated from cosmic noise level trend before and after a flare. The previous day's record may be considered if necessary.
- and I_f = noise diode current required to give a recorder deflection equal to the level at the time of maximum absorption.

6

2800 Mc Observations

The data on solar radio wave events made in Ottawa. Canada by the Radio and Electrical Engineering Division of the National Research Council (A. E. Covington) at 2800 Mc (10-cm emission) are presented, Near local noon (about 1700 UT) the sensitivity of the radiometer is determined and a mean flux for the whole day calculated. These values are given in a tabular form (see table I-1) in units of 10-22 watts/ $M^2/c/s$. Burst phenomena are measured above this level and are given in terms especially suitable for the variations observed on this frequency. The basis for the classifications is described by Covington -J.R. Astro. Soc. Can. 45, 49, 1951 and Dodson. Hedeman and Covington. Ap. J. 119, 541, 1954. A modification in terminology with a view to simplification has been introduced and consists essentially of the omission of the descriptive word "Single" from the "Single-Simple" and "Single-Complex" classes; in designating the "Single", "Single-Simple" and "Rise and Fall" bursts into a single classification designated as "Simple Bursts" with an appropriate type number; in the addition of the letter "f" to indicate that the burst deviates from the basic pattern by the presence of one or more small fluctuations in intensity; and by the addition of the letter "A" to indicate that the event has another smaller duration event superimposed upon it.

Simple Burst

Any single burst which rises to one maximum and then decreases to the pre-burst level.

1 - Simple 1 - Simple burst, type 1 (formerly "single"). Bursts of intensity less than 7 1/2 flux units and duration less than 7 1/2 minutes.

2 - Simple 2 -- Simple burst, type 2 (formerly "singlesimple"). Bursts of impulsive nature with intensity greater than 7 1/2 flux units.

3 - Simple 3 - Simple burst, type 3 (formerly "rise and fall"). Bursts of moderate intensity with duration greater than 7 1/2 minutes.

4 - <u>Post-burst increase</u> -- Postburst level is greater than the preburst level. The gradual return to normal flux may require as long as several hours.

5 - Absorption following burst (negative post).

6 - <u>Complex</u> -- (formerly "single-complex"). A single burst which shows two or more comparable maxima before the activity has declined to zero.

7 - <u>Period of irregular activity or fluctuations</u> -- Series of overlapping bursts of moderate intensity and duration.

8 - Group -- Series of single isolated bursts occurring in succession with intensity between the events equal to the level before and after the group.

9 - <u>Precursor</u> -- A small increase of intensity occurring before a larger increase.

Great Burst

Infrequently occurring bursts of great intensity, often of complicated structure.

Letter "A"

Indicates that this event has another event superimposed upon it.

Letter "f"

Indicates that the basic form of the event is modified by secondary fluctuations.



200 Mc Observations

Data on solar radio emission on 200 Mc recorded by the University of Hawaii (I. Miyake) at Makapuu Pt., Hawaii, are presented. The outstanding occurrences are reported as described under <u>170 Mc Observations</u> with the exception that no intensity measurements are given.

170 Mc Observations

Data on solar radio emission at the nominal frequency of 170 Mc recorded at the Gunbarrel Hill (Boulder) station of the National Bureau of Standards (C.G. Little) are presented. The half width of the antenna lobe is appreciably greater than the solar disk. Polarization is not determined, but the dipole is oriented E-W. All times are in Universal Time (UT or GCT). Observations are interrupted during the period from 26 to 29 minutes after each hour for calibrations.

Beginning January 1, 1959 the method of reducing the records has been changed. The 3-hourly and daily flux density and variability are no longer determined. The outstanding occurrences are reported. However, instead of giving the intensity to the nearest unit of 10^{-22} watts meter $^{-2}(c/s)^{-1}$, a scale of 1 to 3 is now used where for the estimate of smoothed maximum flux:

1 signifies $<100 \times 10^{-22} \text{ wm}^{-2}(\text{c/s})^{-1}$

- 2 signifies >100 <1000 x 10^{-22} wm⁻²(c/s)⁻¹
- 3 signifies >1000 x 10^{-22} wm⁻²(c/s)⁻¹.

Starting and maximum times are read to the nearest 1/10 minute if they are very definite and otherwise to the nearest minute. If the duration is less than five minutes, it is given to the nearest 1/10 minute; otherwise to the nearest minute. The following qualifying symbols are used:

- E = Event in progress before observations began.
- D = Event continues after observations cease.
- I = Event apparently continued during an interruption of the observations. The period of the interruption may be given in the remarks.
- S = Measurement may be influenced by interference or atmospherics.

The types of the outstanding occurrences follow the classification described by Dodson, Hedeman and Owren (Ap J. <u>118</u>, 169, 1953), in which the types are identified by numbers which describe the character of the trace, but not the magnitude of the event, as follows:

0 - <u>Rise in base level</u> -- A temporary increase in the continuum with duration of the order of tens of minutes to an hour.

1 - Series of bursts -- Bursts or groups of bursts, occurring intermittently over an interval of time of the order of minutes or hours. Such series of bursts are assigned as distinctive events only when they occur on a smooth record or show as a distinct change in the activity. 2 - Groups of bursts -- A cluster of bursts occurring in an interval of time of the order of minutes.

3 - <u>Minor burst</u> -- A burst of moderate or small amplitude, and duration of the order of one or two minutes.

4 - Minor burst and second part -- A double rise in flux in which the early rise is a minor burst.

6 - <u>Noise storm</u> -- A temporary increase in radiation characterized by numerous closely spaced bursts, by an increase in the continuum, or by both. Duration is of the order of hours or days.

7 - <u>Noise storm begins</u> -- The onset of a noise storm occurs at some time during the observing period.

8 - <u>Major hurst</u> -- An outburst, or other burst of large amplitude and more than average duration. A major burst is usually complex, with a duration of the order of one to ten minutes.

9A, 9B, or 9 -- <u>Major burst and second part or large event</u> without distinct first and second parts -- If there is a double rise in flux, the first part, a major burst, is listed as 9A and the second part as 9B. The second part may consist of a rise in base level, a group or series of bursts, a noise storm. A major increase in flux with duration greater than ten minutes but without distinct first and second parts, is listed simply as 9.





Note: In the present table, the type classifications 0 and 1 are not used; they have been included above only for information.

169 Mc Interferometric Observations

The 169 Mc interferometric observations are recorded around local noon at Nançay (Cher), France, (N47°23', E8^m47^s) the field station of the Meudon Observatory.

The main lobes are parallel to the meridian plane: the half-power width is 3.8 minutes in the East-West direction and much larger than the solar diameter in the North-South direction. The main lobes are about 2° apart (Ann. Astrophys. 20, 155, 1957). The records give the strip intensity distribution from the center of the disk to 30° to the West and East.

These daily distributions are plotted on the same chart giving diagrams of evolution (C.R. 244, 1460, 1957). Points of intensity 0.5 - 0.75 - 1.0 - 1.5 and 2.0 times 10^{-22} watts/m²/c/s are joined day after day in the form of isophotes. Black dots give the position of the center of the radio spots for each day; a line indicates the width of the recorded lobe pattern when it can be measured with certainty. For each radio spot the smoothed intensity around noon is given in 10^{-22} watts/m²/c/s.

Note that the isophotes cannot be measured when a radio spot of large intensity is on the disk.

Spectrum Observations

Data on solar radio emission in the spectral range 25-580 Mc/s recorded at the Radio Astronomy Station of Harvard College Observatory, Fort Davis, Texas, are presented. The research program is supported by financial assistance from the Air Force Cambridge Research Center, through the offices of Sacramento Peak Observatory.

The receiving equipment consists of five separate sweep frequency receivers covering the bands 25-50, 50-100, 100-180, 170-320, 300-580 Mc/s. The 25-50 and 50-100 Mc/s receivers are each connected to broad band dipoles which are cross polarised and mounted over a reflecting screen. The other three receivers are attached to separate broad band feeds mounted coaxially at the primary focus of an 8.55 meter diameter paraboloid, the 170-320 Mc/s feed being cross polarised with the other two feeds. The effective collecting area of the antenna is 40 sq. meters at 100 Mc/s and 45 sq. meters at 500 Mc/s.



The four types of recognized spectral activity are idealized below:

The large scale examples of continuum, sometimes called Type IV, are listed as "Cont. IV" in the tables. Photographic examples of the bursts have been published by Maxwell, Swarup, and Thompson (Proc. IRE <u>46</u>, 142, 1958), and Maxwell (Sky and Telescope <u>17</u>, 388, 1958; <u>18</u>, 544, and 556, 1959). A few remaining solar radio bursts are tabulated as unclassified. The symbols used in the tables are:

b = single burst
g = small group (<10) of bursts
G = large group (≥10) of bursts
→ = Arrows indicate continuity of solar activity
between two Greenwich days.</pre>

The minimum detectable level of solar activity is a function of frequency: approximately 5×10^{-22} watts meter -2 (c/s)⁻¹ at 500 Mc/s. The equipment records signals over an intensity range of approximately 10,000:1. There are three classes of intensity given in the tables. For 100 Mc/s they are:

1 = Faint, 5 to 40 x 10^{-22} watts meter⁻² (c/s)⁻¹. 2 = Moderate, 30 to 200 x 10^{-22} . 3 = Strong, >200 x 10^{-22} .

The times are Universal Time (U. T.). The accuracy is to the nearest half minute, except in the case of major outbursts which are specified to the nearest O.l minute.

Details of the frequency ranges of activity may be obtained on request to the Radio Astronomy Station, Ft. Davis, Texas.

C. Kp. Ap. and Selected Quiet and Disturbed Days -- The data in the table are: (1) preliminary international character figures, C; (2) geomagnetic planetary three-hour range indices, Kp; (3) daily "equivalent amplitude," Ap; (4) magnetically selected quiet and disturbed days.

This table is made available by the Committee on Characterization of Magnetic Disturbances of IAGA, IUGG. The Meteorological Office, De Bilt, Holland collects the data from magnetic observatories distributed throughout the world, and compiles C and selected days. The Chairman of the Committee computes the planetary and equivalent amplitude indices. The same data are also published quarterly in the <u>Journal of Geophysical Research</u> along with data on sudden commencements (sc) and solar flare effects (sfe).

The C-figure is the arithmetic mean of the subjective classification by all observatories of each day's magnetic activity on a scale of O (quiet) to 2 (storm). Kp is the mean standardized K-index from 12 observatories between geomagnetic latitudes 47 and 63 degrees. The scale is 0 (very quiet) to 9 (extremely disturbed), expressed in thirds of a unit, e.g. 5- is $4 \ 2/3$, 50 is 5 0/3, and 5+ is 5 1/3. This planetary index is designed to measure solar particle-radiation by its magnetic effects, specifically to meet the needs of research workers in the ionospheric field. A complete description of Kp has appeared in Bulletin 12b, "Geomagnetic Indices C and K, 1948" of the Association of Terrestrial Magnetism and Electricity (IATME), International Union of Geodesy and Geophysics.

Ap is a daily index of magnetic activity on a linear scale rather than on the quasi-logarithmic scale of the K-indices. It is the average of the eight values of an intermediate 3-hourly index "ap," defined as one-half the average gamma range of the most disturbed of the three force components, in the three-hour interval at standard stations; in practice, ap is computed from the Kp for the 3-hour interval. The extreme range of the scale of Ap is 0 to 400. The method is described in IATME Bulletin No. 12h (for 1953) p. viii f. Values of Ap (like Kp and Cp) have been published for the Polar Year 1932/33 and for the years 1937 onwards.

The magnetically quiet and disturbed days are selected in accordance with the general outline in <u>Terr. Mag.</u> (predecessor to <u>J.</u> <u>Geophys. Res.</u>) <u>48</u>, pp 219-227, December 1943. The method in current use calls for ranking the days of a month by their geomagnetic activity as determined from the following three criteria with equal weight: (1) the sum of the eight Kp's; (2) the sum of the squares of the eight Kp's; and (3) the greatest Kp.

<u>Chart of Kp by Solar Rotations</u> -- The graph of Kp by solar rotations is furnished through the courtesy of Dr. J. Bartels, Geophysikalisches Institute, Gottingen.

VI RADIO PROPAGATION QUALITY INDICES

One can take as the definition of a radio propagation quality index: the measure of the efficiency of a medium-powered radio circuit operated under ideal conditions in all respects, except for the variable effect of the ionosphere on the propagation of the transmittal signal. The indices given here are derived from monitoring and circuit performance reports, and are the nearest practical approximation to the ideal index of propagation quality.

Quality indices are usually expressed on a scale that ranges from one to nine. Indices of four or less are generally taken to represent significant disturbance. (Note that for geomagnetic K-indices, disturbance is represented by higher numbers.) The adjectival equivalents of the integral quality indices are as follows:

1	Ξ	use less	4	Ξ	poor-to-fair	7	Ξ	good
2	=	very poor	5	=	fair	8	Ξ	very good
3	Ξ	poor	6	=	fair-to-good	9	Ξ	excellent

CRPL forecasts are expressed on the same scale. The tables summarizing the outcome of forecasts include categories P-Perfect; S-Satisfactory; U-Unsatisfactory; F-Failure. The following conventions apply:

P	forecast quality equal to observed	U - forecast quality two or mo grades different from obse when <u>both</u> forecast and obs were > 5, or both < 5	re rved erved
s -	forecast quality one grade different from ob- served	F - other times when forecast ity two or more grades dif from observed	qual- ferent

Full discussion of the reliability of forecasts requires consideration of many factors besides the over-simplified summary given.

The quality figures represent a consensus of experience with radio propagation conditions. Since they are based entirely on monitoring or traffic reports, the reasons for low quality are not necessarily known and may not be limited to ionospheric storminess. For instance, low quality may result from improper frequency usage for the path and time of day. Although, wherever it is reported, frequency usage is included in the rating of reports, it must often be an assumption that the reports refer to optimum working frequencies. It is more difficult to eliminate from the indices conditions of low quality for reasons such as multipath or interference. These considerations should be taken into account in interpreting research correlations between the Qfigures and solar, auroral, geomagnetic or similar indices.

North Atlantic Radio Path -- The CRPL quality figures, Qa, are compiled by the North Atlantic Radio Warning Service (NARWS), the CRPL forecasting center at Ft. Belvoir, Virginia, from radio traffic data for North Atlantic transmission paths closely approximating New Yorkto-London. These are reported to CRPL by the Canadian Defense Research Board, Canadian Broadcasting Corporation, and the following agencies of the U.S. Government:--Coast Guard, Navy, Army Signal Corps, U.S. Information Agency. Supplementing these data are CRPL monitoring, direction-finding observations and field-strength measurements of North Atlantic transmissions made at Belvoir.

The original reports are submitted on various scales and for various time intervals. The observations for each 6-hour interval are averaged on the original scale. These 6-hour indices are then adjusted to the 1 to 9 quality-figure scale by a conversion table prepared by comparing the distribution of these indices for at least four months, usually a year, with a master distribution determined from analysis of the reports originally made on the 1 to 9 quality-figure scale. A report whose distribution is the same as the master is thereby converted linearly to the Q-figure scale. The 6-hourly quality figure is the mean of the reports available for that period. The 6-hourly quality figures are given in this table to the nearest one-third of a unit, e.g. 50 is 5 and 0/3; 5- is 4 and 2/3; 5+ is 5 and 1/3. Other data included are:

(a) Whole-day radio quality indices, which are weighted averages of the four 6-hourly indices, with half weight given to quality grades 5 and 6. This procedure tends to give whole-day indices suitable for comparison with whole-day advance forecasts which seek to designate the days of significant disturbance or unusually quiet conditions.

(b) Short-term forecasts, issued every six hours by the North Atlantic Radio Warning Service. These are issued one hour before 00^{h} , 06^{h} , 12^{h} , 18^{h} , UT and are applicable to the period 1 to 7 hours ahead.

(c) Advance forecasts (CRPL-J) are issued once a week and are applicable 1 to 7 days ahead. They are modified as necessary by the Special Disturbance Warning (CRPL-SDW) applicable 1 to 3 days ahead, which may be followed by a supplementary forecast (CRPL-Js) applicable to days remaining until next CRPL-J forecast. The forecast entitled "final" consists of the most recent of the above forms and is scored against the whole-day quality index.

(d) Half-day averages of the geomagnetic K indices measured by the Fredericksburg Magnetic Observatory of the U.S. Coast and Geodetic Survey.

A chart compares the short-term forecasts with Qa-figures. A second chart compares the outcome of the final advance forecasts with a type of "blind" forecast. For the latter, the frequency for each quality grade, as determined from the distribution of quality grades in the four most recent months of the current season, is partitioned among the grades observed in the current month in proportion to the frequencies observed in the current month.

Ranges of useful frequencies on the North Atlantic radio path are shown in a series of diagrams, one for each day. The shaded area indicates the range of frequencies for which transmissions of quality 5 or greater were observed. The blacker the diagram, the quieter the day has been; a narrow strip indicates either high LUHF, low MUF, or both. These diagrams are based on data reported to CRPL by the German Post Office through the Fermeldetechnischen Zentralamtes, Darmstadt, Germany, being observations every one and a half hours of selected transmitters located in the eastern portion of North America. Since January 6, 1958 the transmitters monitored are restricted to those located north of 39° latitude. The magnetic activity index, A_{Fr} , from Fredericksburg, Va., is also given for each day. <u>Note</u>: Beginning with data for September 1955, Qa has been determined from reports that are available within a few hours or at most within a few days, including for the first time, the CRPL observations. Therefore these are the indices by which the forecasters assess every day the conditions in the recent past. Over a period of several years, they have closely paralleled the former Qa indices which excluded CRPL observations and included three additional reports received after a considerable lag. Qa was first published to the nearest one-third of a unit at the same time.

North Pacific Radio Path -- The CRPL quality figures, Qp, are compiled by the North Pacific Radio Warning Service (NPRWS), the CRPL forecasting center at Anchorage, Alaska, from radio traffic data for moderately long transmission paths in the North Pacific equivalent to Seattle-to-Anchorage or Anchorage-to-Tokyo. These include reports to CRPL by the Army Command and Administrative Network, U. S. Air Force and Federal Aeronautical Administration. In addition, there are CRPL monitoring, direction finder observations and field strength measurements of suitable transmissions.

The original reports are on various scales and for various time intervals. The observations for each 8 hours or 24 hour period are averaged on the original scale. This average is compared with reports for the same period in the preceding two months and expressed as a deviation from the 3-month mean. The deviations are put on the 1 to 9 scale of quality which is assumed to have a standard deviation of 1.25 and a mean for the various periods as follows:

07-18	hours	UT	5,33	00-24	hours	UT	5.67
19-06			6.00				

The 12-hour and 24-hour indices Qp are determined separately. Each index is a weighted mean where the CRPL observations have unit weight and the others are weighted by the correlation coefficient with the CRPL observations.

The table, analagous to that for Qa, includes the 12-hourly quality figures; whole day quality figures; short-term forecasts issued by NPRWS twice daily at O6^h and 18^h UT, applicable to the stated 12-hour periods; advance forecasts issued weekly by NPRWS (CRPL-Jp report) modified as necessary by Special Disturbance Warnings (CRPL-SDW) and supplementary forecasts (CRPL-Jps); and half-day averages of geomagnetic K indices from Sitka.

The chart compares the outcome of the final advance forecasts, on the same basis as the similar chart for the North Atlantic Radio Path.

Note: Beginning with March 15, 1959 the short-term forecast schedule was changed from three times daily to twice daily. The North Pacific quality figures used for evaluation are now 12-hourly rather than 8-hourly.

VII ALERT PERIODS AND SPECIAL WORLD INTERVALS

This table gives the Advance Geophysical Alerts as initiated by the Western Hemisphere Regional Warning Center at Ft. Belvoir, Va., and also the Worldwide Geophysical Alerts and Special World Intervals as designated by the World Warning Agency, Ft. Belvoir, Va.

Advance Alerts are of four types, defined as follows:

1 - <u>Solar Flare Alert</u> -- this warning is issued whenever a solar flare of median importance 2 plus or greater has been reported. There will be only one alert issued per flare and only one a day at most.

2 - <u>Magnetic Storm Alert</u> -- this warning is issued whenever a significant magnetic storm, K figure 5 or greater at a middle latitude station has begun.

3 - <u>Cosmic Ray Alert</u> -- this warning is issued whenever a very outstanding change in cosmic ray flux has been observed -- increase or decrease.

4 - <u>Aurora Alert</u> -- this warning is issued whenever a magnetic storm in middle latitudes has reached K figure 7 intensity or whenever selected auroral stations report the presence of outstanding aurora.

Worldwide Alerts are of the same types as the Advance Alerts, except that the Solar Flare Alert and Cosmic Ray Decrease Alert are omitted. Alert announcements include the event and time of event upon which the alert is based, and, in the case of the Advance Alerts, the station reporting the event.

The World Alerts and Special World Intervals are issued by the World Warning Agency on decisions based on Advance Alerts, advice received from Regional Warning Centers and overall policy.



Dec. 1959	American Relative Sunspot Numbers R _A '	Jan. 1960	Zürich Provisional Relative Sunspot Numbers R _Z	Daily Values Solar Flux at 2800 Mc, Ottawa, Canada Flux
1	164	1	136	
2	192	2	110	175
3	194	3	133	182
Ĩ4	174	- 4	156	193
5	155	5	158	213
6	152	6	174	215
7	131	7	167	224
- 8	138	8	153	219
9	97	9	150	201
10	76	10	127	194
11	72	11	143	200
12	74	12	108	184
13	79	13	108	178
14	82	14	118	176
15	120	15	112	183
16	108	16	119	183
17	105	17	117	179
18	131	18	89	176
19	163	19	80	164
20	174	20	94	157
21	154	21	103	162
22	116	22	134	172
23	110	23	138	188
24	121	24	130	210
25	143	25	152	230
26	134	26	209	242
27	150	27	186	248
28	153	28	159	252
29	120	29	193	237
30	113	30	178	230
31	131	31	178	224
Mean:	129.9	Mean;	139.1	199.7

COMMERCE - STANDARDS - BOULDER

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CALCIUM PLAGE AND SUNSPOT REGIONS

JANUARY 1960

CMP		McMath	Return		Calcium P	lage Data	Su	inspot	Data
Jan.	Lat	Plage	of	CMP	Values	114 - 4	CMP Va	lues	
1900		Number	Region	Are	a int.	History, Ag	e Area (ount	History
01.1	N09	5512	5478	2000	2.5	$\ell - \ell = 2$	80	7	Ъ∕ℓ
02.1	N22	5513	5478	3000	2.5	l l 2	50	2	ℓ \ d
03.2	NU8	5519	New 5482	400		bAd 1			
04.5	1 315	5516	New	1400	2.5	l = d = 5	80	4	$\ell - \ell$
		5510	new	1400	2.5		20	T	l V l
05.6	N09	5517	5484	700	2.5	$\ell = \ell = 5$	320	6	$\ell - \ell$
06.6	S18	5515	5486	3 5 0 0	3	$\ell - \ell = 4$	370	3	l – l
07.6	N07	5521	5495	900	2	$\ell - \ell = 2$	50	2	b∧d
07.7	N25	5520	New	2700	2.5	$\ell = \ell = 1$	20	1	l / l
09.0	S23	5522	New	3200	3	$\ell = \ell = 1$	440	2	l — l
09.1	N17	5524	New	600	2.5	l = l = 1	(20)	(1)	171
09.3	S01	5523	New	2100	3	$\tilde{\ell} - \tilde{\ell}$ 1	560	14	a - a
10.2	S16	5525	New	6300	3	$\ell = \ell = 1$	1020	6	l \ l
10.2	S03	5526	5490	600	2.5	<i>lll</i> 4	70	3	ЪΛ̃ď
11.6	N20	5527	5491	3500	2.5	$\ell = \ell = 2$	730	5	$\ell - \ell$
12.5	S21	5530	New	500	1	0,001	50	1	2
13.4	N03	5528	*	3000	2		1 30	L	07 £
14.4	N25	5531	New	500	1.5		1		
15.1	S16	5529	New	300	1.5	lend 1			
15.9	N10	5532	5497	600	2	l\d 6			
16.3	506	5533	5/00	800	25		50	,	
16.7	N22	5535	New	500	2.5		50	T	l \ d
17.7	N09	5534	5501	2000	2.5				
17.8	S18	5536	5500	800	1	l l l l 4	60	1	e v d
19.3	NO 7	5538	New	2200	3.5	$\ell - \ell = 1$	100	5	$\ell - \ell$
19.5	N20	5537	5502	1700	2				
21.1	N09	5540	**	3700	2		150	<i>I</i> .	0.1
21.5	N26	5539	**	2600	3		540	4	
22.2	N13	5541	**	2900	3	l = l 3	(20)	$\frac{1}{1}$	
23.0	S13	5543	New	700	2	$\tilde{\ell} - \tilde{\ell} = 1$	100	3	δΛd
	1101								
23.3	N26	5542	**	3200	3	$\ell - \ell = 3$			
23.4	N11	5545	**	2900	3	$\ell - \ell = 3$	390	1	l – l
25.0	N12	5540	5507	(5600)	(2.5)	$\ell = \ell = 10$	100		
26.5	N24	55/.0	5500	(2000)	(2)	k = k = 2	190	1	$\ell = \ell$
20.0	1124	5540	505	4900	د	x - x - 4	610	S	l _ l
27.2	N09	5549	***	2000	3	$\ell - \ell = 3$	400	10	l – l
29.5	N12	5550	***	11000	3	$\ell - \ell = 3$	580	26	l – l
31.8	S18	5551	5514	5000	3	$\ell - \ell = 6$	510	12	l — l
							COMMERCE -	STANDA	RDS - BOULDER

* 5493 and 5494. ** 5504,5505, 5506. *** 5511, 5512, 5513.

CORONAL LINE EMISSION INDICES

JANUARY 1960

								C
ant ater)	ч Г	24 74 74	к к Г †	\$6, × 6, 7				
t Quadr davs 1	R6	30 X X X X	х х ³ х х	725 × 41 ×	× × × × ×	x x 19a x	* * * * *	X
th Wes rved 7	G1	199 109 x x	x 167 x	137 137 107 99	x 174 x 208 x	* * * * *	* * * * *	X MENCE -
Nor (obse	6 ₆	161 92 × ×	x 131 x x	113 x 84 82	107 107 137 x	* * * * *	* * * * *	CON X
ant ater)	RI	48 8 8 8	x x 9 65 x x	20 3 x 36 x	х х х х I O	x x 2la x	* * * * *	×
, Quadra davs la	R ₆	19 7 x x x	x x 41	12 23 x 22 x	ххх Ц	x x x 16a	* * * * *	×
th West rved 7	G ₁	124 85 x x x	x 182 x x	x 135 x 98	x 114 x 47 x x	* * * * *	* * * * *	×
Sou Cobse	9 ⁰	75 56 x x x	х 127 х	87 87 76 76	х 3 х 5 х	****	* * * * *	×
drant earlier)	RI	x x 51 51	75 x x x x	24 × × 20	lo × × × ×	16 × × × ×	24 39 20 x	×
с Quadre lavs ear	^в 6	x x 24 27	34 × × × × ×	19 × × × 15	∞ × × × ×	15 X X X X	13 × 23 × 13	×
th East ved 7 d	Ъ,	x x 115 370	190 x x 126	x x 78 104	78 × × × × × × × ×	x 47 x x x x	135 x 104 117 x	261
Sou (obser	9 ⁰	x x 91 148	146 × × ×	х х 74	57 × × ×	× × × × × × × ×	93 61 51	157
nt lier)	R,	x x 36 24	36 x x x x x	33 x 40	ж ж ж ж ж ж	x x x x	36 54 x 4	×
Quadre avs ear	R, R	20 x x x 20 15	29 X X X X	24 x x 25 25	17 * * * *	52 x x x	21 x 40 30 x	×
th East ved 7 d	5	x x 120 167	194 x x 102	132 x 114 128	107 × × ×	156 x x	138 x 152 178 x	267
Nor (obser	^д	× × × × 803	162 x x 88	711 × 001 011	78 × × ×	x 124* x x	123 x 132 * 143 x	219
CMP	1960	10045	96836	12525	16 17 19 20 20	22 22 25 25 25	26 28 30 30 30 30 30 30 30 30 30 30 30 30 30	31

Hb

* = yellow line observed.

a = index computed from low weight data.

x = no observations.

FLARES	1960
SOLAR	JANUARY

PROVISIONAL	IONOSPHERIC	EFFECT						S-SWF				Slow S-SWF	G-SWF		Slow S-SWF	
	MAX.	INT.				30			19000 1900	13		30	17	10	20	
	MAX.	WIDTH Ha													2.50	
ASUREMENTS	CORR.	AREA Sq. Deg.		900 900 900		6 • 70 4 • 95	2•50 3•00	13.00			5 • 00			3.40	12.00 11.50	3.00
ME	MEAS.	AREA Sq. Deg.	1.40		2.10 1.10	1.90 1.65 1.10 1.10 2.20	2.50	3.70	2.50 2.50 6.40 6.85 3.10	2.76		22.00 10.30 1.90	2•60	1.70 2.10	4.80 2.00	
	TIME	1 1 1	0158 1618		1947 1948	0852 0848 2004 2029 2029 2149	0920	1514	1825 1825 1840 1850			2128 2124 2250 2302		1217 2305	1401 1404 1658	
OBS.	, COND.		е г		ωm	100101	ωω	ო ო ო	N N N N N	2	ი ო	 0000	ŝ		001-	
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DURA-	NOI I	MINUTES	32 D 43	. 35 D 11 D 13 D	26 D 4	39 7 33 25 25 25	4 D 12	10 D 38 D	120 120 170 164 U 148 D	24	4 18 D 31 D	195 D 148 D 98 14	24	16 D 63	26 D 79 D 45 D	16 D
NO	McMATH	PLAGE	5507 5520	5511 5520 5515	5513 5525	5525 5525 5512 5512 5512 5512	5517 5525 5525	5525 5512 5512	5515 5515 5520 5520 5520	5527	5515 5516 5516	5527 5527 5527 5515	5525	5525 5530	5525 5525 5533 5533	5539
LOCATIC	BOX.	MER. DIST.	w 54 E 8 0	W20 E64 E55	W 28 E 85	Е 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	E 04 E 65 E 58	E40 W71 W78	W33 W12 W14 W14	E21	W5 0 W9 0 W85	E 03 E 02 W 08 W 90	W37	W44 W08	W 66 W 71 W 69 E 08 F 90	E73
	APP	LAT.	N12 N24	N17 N13 S21	N17 523	S19 S14 N08 N08 N09 N15	N10 511 512	S12 N08 N07	S23 S23 N28 N26 N26 N28	N 19	S17 N08 N08	N23 N22 N18 S12 S12	S10	S18 S03	520 520 520 520 506	N26
		MAX. PHASE	0158		1948	2029 2030 2149			1810 1825 1840 1842	2112		2128 E 2124 2250 2302	1650	2305	1658 1732	
OBSERVED	INIVERSAL TIME	END	0200 D 1650	1025 D 1040 D 1333 D	2010 1952	0906 0854 2014 2058 2048 22048	0959 D 1007	0819 D 1542 D	1950 1950 2050 2050 U 2048	2130	0815 0925 D 1210 D	2355 D 2326 2422 2312	1710	1232 D 2350	1400 D 1455 D 1412 D 1738 U	1021 D
	P	START	0128 1607	0950 E 1029 1320	1944 E 1948	0327 0847 1956 E 2025 2142 2142	0920 E 0955 0955	0809 E 0819 E 1504 E	1750 1750 1800 1806 1820 E	2106	0811 0907 E 1139 E	2040 U 2058 E 2244 2258	1646	1216 E 2247	1334 1336 1400 E 1645 1730 D	1005
DATE	NAU	1960	01 01	000	03 03	000000 444444	0050	07 07 07	888888	60	10 10	====	12	13 13	1155	16
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PROVISIONAL	IONOSPHERIC	EFFECT	Slow S+SWF	Slow S-SWF		G-SWF										
	MAX.	INT.	40	30 30		17 30							16	16	20 16	э о
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CASUREMENTS	CORR.	AREA Sq. Døg.					3 • 00		00.4	14.00			3.10	5 • 00	2.20	3 • 00 3 • 00
IM	MEAS.	AREA Sq. Deg.	1.40 5.00 1.50	1.10 4.20 2.20	1.00 1.40	3.97	1•30	1, 90 1, 20		4 • 00	1.40		2•60 3•00	5 • 00 4 • 78	2.00 2.40 2.43	2•00 1•40
	TIME	1 5	2112 2252 2248	0036 1720 1720 1810	2308 2328	1950	2302	0002 1041 1058 1122 2038	0858 0938 1245	1313 1321	1243 2008		1530	0832	1040 1746	00150110
OBS.	COND.		222	8118	<i></i>	1	9	~~~~~	<i>ოოო</i>	ωN	53	2	22	1	7 1	0 0 0
IM.	POR-	TANCE	- 2 -	666 110 111		1 1	пп	°		16	11	1	11	$\frac{1}{16}$	ннн	папал
DURA-		MINUTES	4 D 56 18 D	14 D 112 D 112 D 56 D	9 D 68 88	60 88 D	68 D 16	14 44 D 13 D 8	9 13 D 27 D	52 D 26 D 50 D	22 28	6 D	26 24 D	86 D 42	7 D 32 42	28 8 D 27 D 23 D 5 D
NO	McMATH	PLAGE	5539 5540 5540	5545 5545 5545	5541 5545 5541	5545	5528 5541	5538 5538 5538 5538 5549	5550 5538 5550	5549	5550 5550	5550	5549 5549	5550 5550	5551 5550 5554	5550 5550 5551 5551 5551
LOCATI	IOX.	MER. DIST.	E62 E76 E76	E67 E62 E62 E72	E41 E58 E49	E46 E45	W57 E25	W16 W59 W59 F490	E57 W69 E55 F55	E35 E35	E 55 E 50	E31	E12 E13	E12 W05	E27 W17 E58	w22 w14 E21 E17 E49
	APPE	LAT.	N26 N11 N13	N11 N17 N17 N17 N23	N08 N17 N16	N14 N18	N07 N12	N27 N09 N09 N12 N12	N06 N08 N08 N06	000 800 800 800 800	N06 N12	50N	10N N05	N07 N00	S15 N06 S19	N03 N10 S15 S17 S17 S18
		MAX. PHASE	2112 2252 2248	0036 1720 1750	2308	1944 1950 U	2302	0002 1122 2038	0858		1243 2008		1528	2044	1746 1846	0015
OBSERVED	UNIVERSAL TIME	END	2112 D 2335 2258 D	0046 D 1900 U 1900 U 1856	0943 D 2358 2442	2028 . 2115 U	1146 D 2316	0014 1044 D 1104 1130 2046	0904 0941 1308 D	1358 D 1335 D 1402 D	1303 2032	1028 D	1544 1544 D	0954 D 2118	1045 D 1810 1902	0033 0118 0922 D 0926 D 0932 D
		START	2108 2239 2240	0032 1708 E 1708 E 1800 E	0934 E 2250 2314	1928 1947 E	1038 E 2300	0000 1040 E 1056 E 1117 2038	0855 0928 E 1241 1242 F	1306 1309 E 1312 E	1241 2004	1022 E	1518 1520 E	0828 E 2036	1038 E 1738 1820	0005 0110 E 0855 E 0903 E 0927 E
DATE	NAI	1960	16 16 16	17 17 17	18 18 18	19	20	23333 23333	5 4 4 5 4 4 5 4 4	2 t	25	26	27 27	28 28	29 29 29	00000
	OBSERVATORY		HAWAII LOCKHEED HAWAII	HAWAII LOCKHEED LOCKHEED HAWAII	ARCETRI HAWAII HAWAII	<pre>{ sac peak { LockHeed</pre>	WENDEL HAWAII.	HAWAII ONDREJOV ONDREJOV ONDREJOV HAWAII	ONDREJOV ONDREJOV ONDREJOV	WENDEL ONDREJOV CAPRI S	ONDRE JOV HAWAII	ARCETRI	{ SAC PEAK CAPRI S	CAPRI S SAC PEAK	CAPRI S LOCKHEED SAC PEAK	LOCKHEED HAWAII WENDEL ARCETRI WENDEL

SOLAR FLARES JANUARY 1960 COMMENCE - STANDARDS - BOULDER

IIIb

TART END	MAX. LAT. PHASE NO 9	PROX. McMATI MER. PLAGE	H TION		-coo						PROVISIONAL
stAar Erto 035 1057 138 E 1150 D 1445 1150 D 1442 D 159 E 1229 D 159 E 1524 D 355 E 1524 D 336 E 1320 D 336 E 1320 D 336 E 1320 D 336 E 1524 D 336 E 1524 D 336 E 1522 D	MAX. LAT. PHASE NO 9	MER. PLAGE		POR-	COND.	TIME	MEAS.	CORR.	MAX.	MAX.	IONOSPHERIC
035 1057 045 1101 0 140 E 1150 0 159 E 1152 0 159 E 1252 0 159 E 1252 0 336 E 1322 0 336 E 1322 0 334 1350 0 432 E 1502 0	815 S15		NINUTES	TANCE		۲ <mark>۲</mark>	AREA Sq. Deg.	AREA Sq. Deg.	WIDTH Ha	INT. %	EFFECT
1138 E 1150 D 1140 E 1152 D 1159 E 1229 D 1355 E 1524 D 1366 E 1320 D 1366 E 1320 D 1334 1332 D 1334 1330 D 1432 E 1502 D	101	E36 555	2 22					3.00			
1140 E 1159 E 1159 E 1355 E 1355 E 1366 E 1326 E 1326 E 1336 E 1336 E 1336 E 1336 E 1336 E 1320 D 1334 1320 1335 E		N21 5550									
1159 E 1229 D 1159 E 1229 D 1355 E 1524 D 1306 E 1320 D 1306 E 1320 D 1334 1350 D 1432 E 1502 D	N16	E19 555	2 12 D					4 • 00			
1159 E 1442 D 1355 E 1524 D 1306 E 1320 D 1306 E 1320 D 1334 1350 D 1432 E 1502 D	S16	E20 555	1 30 D	1				4 e 00	-		
1355 E 1524 D 1306 E 1320 D 1306 1321 1334 1350 D 1432 E 1502 D	S17	E19 555	1 163 D	16				8.00			
1306 E 1320 D 1306 1321 1334 1350 D 1432 E 1502 D	S17	E21 555	1 89 D	1	ŝ	1420	2.00	2.20			
1306 1321 1334 1350 D 1432 E 1502 D	N03	W27 5550	0 14 D		ŝ	1310	3 • 00	3.30			
1334 1350 D	NO3	W26 5550	41 0	<u>ا</u> د		-		5 ° 00			
1432 E 1502 D	N21	E39 555	5 16 D	1				4 • 00			
_	N12	E35 5552	2 30 D	-	m	1436	3¢00	3.70			
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SOLAR FLARES JANUARY 1960

INTERVALS OF NO FLARE PATROL OBSERVATIONS JANUARY 1960



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SUBFLARES Noted as follows: Date-Universal Time-Coordinates DECEMBER 1959

SAC PEAK LOCKHEED LOCKHEEO	01 1504 N15 W75 01 1649 N20 E16 01 1806 N24 W85	* HAWAII 05 1946 E N11 W04 LDCKHEEO 05 2002 N07 E50 HAWAII 05 2008 N13 W10	SAC PEAK 13 1654 N15	E 20
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SUBFLARES Noted as follows: Date-Universal Time-Coordinates DECEMBER 1959

HAVAIT 26	1826 E	N28 W47	SAC PEAK	2.8	1752	N17 W18	CAPRI S	30	0824 E	509	E05
LOCKHEED 26	1914	N28 E44	SAC PEAK	2.8	1804	N08 W36	LOCKHEEO	30	1546	N12	w57
LOCKHEED 26	2011	S17 W75	HAWATT	2.8	2012	N12 N4I	LOCKHEEO	30	1552	N10	W90
LOCKHEED 24	2041	NZB FAB	HAWALI	2.8	2050	512 F28	LOCKHEEO	30	1820 E	N13	W50
LOCKHEED 20	21041	N22 UAB	LOCKHEED	2.8	2216	N09 W12	LOCKHEEO	30	1820 E	N13	₩50
LUCKHEED 26	21104	C17 U76	LOCKHEEO	2.8	2220	508 F25	LOCKHEEO	30	1830	NIL	W68
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SAC PEAK 27	1460	N11 E27	LOCKREED	27	1702	N30 C47	ED CHITEED		L.A		
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HAWAII 27	2026	N08 W63	LOCKHEEO	29	1938	N16 W88	SAC PEAK	31	1742	N12	¥61
LOCKHEED 27	2211	N11 W27	LOCKHEEO	29	1938	N16 W88	LOCKHEEO	31	1742	N12	w59
HAWAII 27	2217 E	N10 W28	LOCKHEED	29	2020	N10 W53	1 I AWAH	31	1746	N11	W65
HAWA1I 27	2230	N19 E67	HAWAII	29	2022	N20 W54	HAWAII	31	1846	N11	W90
LOCKHEED 27	2233 U	N21 E70	LOCKHEED	29	2045	S07 E13	LOCKHEED	31	1915	512	E80
HAWA1I 27	2254 E	N30 E29	LOCKHEED	29	2050	NI0 W53	LOCKHEED	31	2128	N06	E50
			HAWAII	29	2222	N10 W56	LOCKHEED	31	2133	N13	E07
ARCETR1 28	1003 E	N08 W26	HAWAII	29	2356	510 E09	HAWAII	31	2134	506	W14
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SOLAR FLARES OCTOBER 1959

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OBSERVED	UNIVERSAL TIME	END	0807 1137 1114 1455 1455 1455 1450 1441 1441	1104 1107 1304 0546 D	0455 0624 0546 D 0518 0518 0523 D	1034 1034 1050 1050 11247 1247	0454 0834 1032 1245 D 1257 1257 1414 1444 D 1444 D 1444 D	0653 0815 0915 0948 0948 1258 1258 1249 1249 1249	0645 0755
ŝ		START	0748 1051 E 1057 1425 1428 E 1428 E 1430 1431 E 1432 E	1054 1058 E 1249 0326 E	0329 0437 0456 0457 0836 0836	1029 E 1029 E 1029 E 1030 E 1102 1242 E 1440 E	0433 0828 1022 1123 E 1147 E 1148 1148 1148 1442 1442 2258	0636 08636 0843 C 0917 0917 1232 1234 1243 C 1343 C 1343 C	0635 0713 E
DATE		0CT 1959	000000000000000000000000000000000000000	09 09 09 10		00000000		××××××××××××××××××××××××××××××××××××××	13 13
	OBSERVATORY		CAPRI S CAPRI S LOCARNO CAPRI S MEUDON ONDREJOV	<pre>{ PIRCULI CAPRI G LOCARNO J ALMA-ATA</pre>	C SYDNEY TASHKENT ALMA-ATA SYDNEY SYDNEY	<pre>{ OUDREJOV ZURICH ZURICH COLARNO GOOD HOPE ONDREJOV CAPRI G</pre>	SYDNEY CAPRI G CAPRI S CAPRI S CAPRI S GOOD HOPE GOOD HOPE GOOD HOPE SYDNEY	GOOD HOPE PIRCULI FIRCULI COCARNO MEUDON PIRCULI MEUDON ZURICH ZURICH ZURICH ZURICH	LOCARNO { CAPRI G

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	OBSERVED		D.L	CATION		DURA-	ż	OBS.	8	IM	LAS UREMENTS			TWNOIBLAOBA
INIVE	TIME TIME		APPROX	C. Mch	MATH	NOIL	POR-	COND.	TIME	MEAS.	CORR.	MAX.	MAX.	IONOSPHERIC
	END	MAX. PHASE	LAT. D	IER. PL	AGE GION	INUTES	TANCE		15	AREA Sq. Deg.	AREA Sq. Deg.	WIDTH Hα	.ENI %	EFFECT
000	805 925	0821	S15 W S19 W	188	401 401	22 D 70	101		0743 0821	1•90 1•20				
00000	8357 D 9907 0 900	0852	N188 N196 N188 N196 N188 N196 N196 N196 N196 N196 N196 N196 N196	14 18 18 18 18 18 18 18 18 18 18 18 18 18	401 426 426 424	30 D 28 D 15 D 15 D	н н н го	n n ⊣ n n	0852 0838 0858	1.10	11•00 3•30	1•40 2•50	185	
10	101		M OON	2 021	401	3 D		0 0	1358		1 • 00			
1000	649 D 52 D	0533 0600	S09 % S08 E N06 W	139 5 125 5 185 5	423 424 405	31 D 4 7 D		9 1 6	0533 0600 1246	1.29 .73	1.85 .82 4.00		4 5 0 0	
	200 D 245 D	1149 1149	N28 W S10 W N29 W	152 5 157 5	408 423 408	21 D 21 D 15 D	ппп	~~- ~	1149	1•19 •92	2 •1 3 1•88		45	
онана	726 D 101 D 120 120 232		N 2005 W N 2	63 5 5 5 5	418 418 418 418 408	55 D 20 D 20 D 20 D		<i>ო</i> ₽ ₽ <i>ო</i>	0722 1036 1100 1227	2•60	2 00 2 00			
00000	104 735 D 801 D 920	0057 0705 0900 11	S08 N27 N31 N31 N31 N31 N31 N31 N31 N31 N31 N31	140 5 176 5 174 5 53 55	424 408 408 427	112 2000 2000		~~~	0057 0705 0754 0754 0900	1-50 1-47	2 00 5 4 1 2 00	2•60	2 2 2 2 2 2	
60	115 21 00 D	0852	S05 E S05 E S05 E	56 56 56 56	427 427	244 c 30 D 8 D		i m	0852	1,20	3.00		>	
09 14 23	16 18 D 08 20	0858 0856	S04 E S05 E N04 E S07 E	50 01 50 50 50 50 50 50 50 50 50 50 50 50 50	427 420 420	23 D 23 D 22 D 64 D	16 1 16	~~~~~	0858 0908 1408 2218	1.10 2.72 3.34	3.18 5.06 3.00	2 • 50	64 70 73	
00	01 32	2343 0121	N07 E N29 W	.46 5. 86 54	430 408	24 20	16	6 N	2343 0121	3.07 .81	4.28 4.77		80 78	
000000000000000000000000000000000000000	117 30 D 58	0607 0819 U 2256	0021022	4 4 4 2 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4	427 427 427 427 427	14 15 D 12 D	16 1 1 1 1 1 6	N N N N N	0608 0611 0819 1117 2256	1•01 1•82 1•44	2.00 3.20 2.13	3 80 2 90	110 61 91	
90	32 D 15 20	20707	N02 S03 E N09 E E N09 E	32 5	427 427	8 0 12 0 12 0 2 2		n 00 €	0627 0707 0710	06•	1 • 00 4 • 00	2.20	74	
0.046	22 56 41	1025 1329	хоохо 2000 2000 2000 2000 2000 2000 2000	14 55	418 425 425 425	1900 1900 1900		กุลลุ	1025 1322 1322 1329	3.11	00444 0444 0400	2.10	66	S – SWF S – SWF

COMMERCEY - STANDARDS - BOULDERS

SOLAR FLARES OCTOBER 1959

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SOLAR FLARES OCTOBER 1959

PROVISIONAL	IONOBPHERIC	EFFECT							MONITOR - MOLIN OF
	MAX.		120	\$ \$ \$	63 76 52 106 146	140	5 0 0 0 0 0	5 2 2 2 2 2 2	
	MAX.	на	1.90 2.30 1.16		1•60 2•30 2•90	1 • 36 3 • 90			COMMENSION
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ME	MEAS.	Sq. Deg.	2•50 1•45	1 • 00 • • 46 • 20 • 20	1.08 1.82 1.682 1.602 1.602 1.682 645	2.08 1.46 1.20	3 • 00 1 • 13 • 55	1.80 .81 .60 1.60 .01 .46	6 • 00
	TIME	υŢ	1351 1400 1405 1407 2335	0630 0853 0844 1215 1239 1333 1333 1417	0800 0800 0800 0800 0803 0804 0804 11100 0956 1114 0956 11440 0956	0658 0710 0801 1021 1241 1241 1249 1426	0313 0734 0750 1123 1515 1525 1525	0802 0752 0755 0758 0804 0828 0940	0150
OBS.	COND.		~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~	NNWWW WW	N N M M M M M M M M	<u>നന നന</u>	<u> </u>	<u> </u>	9
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NC	McMATH	REGION	5427 5430 5430 5433 5433 54133	5433 5430 5430 5433 5433 5433 5423 5427 5427 5427	5418 5418 5418 5427 5427 5423 54218 54218 5423 54233 54233	5438 5433 5433 5423 5423 5423 5433	5431 5427 5421 5431 5433 5433 5433 5433 5433	5431 5431 5431 5427 5423 5433	5428
LOCATI	tox.	DIST.	E26 E21 E21 E22 E90 W55	E85 E816 E816 E816 E816 E816 E816 E815 E815 E815 E815 E815 E815 E815 E815	К К К К С С С С С С С С С С С С С С С С	E90 W17 E89 E77 W66 W17 W17 E48	¥ 19 19 19 19 19 19 19 19 19 19 19 19 19	ы 198 198 198 198 198 198 198 198 198 198	W50
	APP LAT.		S06 N08 N09 N17 N17 S07 S07	N05 N06 N068 N066 N066 N066 N066 N066 N066	S10 S10 S055 S055 S055 S055 S055 S055 S0	S11 S10 S10 S10 S10 S113 N13	S11 S13 N04 N05 N05 S08 S08 S09	S05 S08 S12 S03 S03 S03 S02 S02 S12 S12	N40
		PHASE		0853 U 0844 1333	0758 0800 0803 0814 0814 0848 0956 1003 1144	0801	0313 0734 0750 0750	0 755 0 758 0 804 0 828	0150
OBSERVED	UNIVERSAL TIME		1354 1414 D 1430 1420 1530 D 2344 D	0656 0910 0849 0849 0849 1245 1412 1402 1412 1402 1435 0	0810 0805 0810 D 0821 0845 0905 11006 1145 D 1145 D	0712 D 0715 D 0821 D 1035 D 1259 D 1257 D 1455 D	0327 D 0740 0755 0800 1132 D 1520 1532 D	0814 D 0756 D 0801 0815 0815 0815 0810 0837 0958	0234 D
	START		1340 1356 1357 E 1358 E 1507 E 2335 E	0630 E 0843 E 0843 E 0843 E 1210 E 1326 E 12	0756 0758 0802 E 0802 E 0836 E 0836 E 1000 1139 E 1142 E	0658 E 0709 E 0758 E 1220 E 1236 1248 E 1424	0302 0729 0738 0740 1120 E 1517 1517	0747 0750 E 0750 E 0754 0759 0822 0822	0128
DATE	1.00	1959	19 19 19 19 19	000000000000000000000000000000000000000	222222222	222222222222222222222222222222222222222	00000000000000000000000000000000000000	* * * * * * * * * * * * * * * * * * *	25
	OBSERVATORY		t ONDREJOV { ONDREJOV CAPRI G CAPRI G CAPRI G KYOTO	6000 HOPE SIMEIZ SIMEIZ CAPRI 6 CAPRI 6 CAPRI 6 CAPRI 6 CAPRI 6 CAPRI 6 CAPRI 6	<pre>{ SIMEIZ</pre>	KYOTO CAPRI G TASHKENT CAPRI G GOOD HOPE CAPRI G CAPRI G CAPRI G	SYDNEY PIRCULI PIRCULI CAPRI G LOCARNO LOCARNO CAPRI G	<pre>CAPRI S CAPRI G CAPRI G CAPRI G F CAPRI G F CAPRI G CAPRI G CAPRI G</pre>	SYDNEY

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PROVISIONAL	IONOSPHERIC	EFFECT																	DARDS - BOULDER
	MAX.	.TNI %	48											100		100	100	115	5 - 3TAN
	MAX.	WIDTH Ha							1•90										COMMUNIC
ASUREMENTS	CORR.	AREA Sq. Deg.	1 . 31		4 • 00		1.00	4 • 00		2.80	4 • 00		2.50	1•40	2.00			• 80	
ME	MEAS.	AREA Sq. Deg.	1-19							1.20		• 90	2.40	2.27	2.00	1.25	3.74	1.40	
	TIME	I D	0704		1202		1452	0748	1046	0742	0948	1039	1051	, 1049	0158	0159	0158	0914	
OBS.	COND.			ŝ	e	ŝ	ю	ŝ	ę		ŝ			1	2			2	
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DURA-	NOIL	MINUTES	10	30	12 D	7	13	13 D	16 D	11	5 D	18	12	21	21	7 D	0 6	ŝ	
N	McMATH	PLAGE	5433	5438	5433	5438	5437	5433	5433	5431	5433	5431	5433	5440	5438	5438	5438	5437	ĺ
LOCATIC	ROX.	MER. DIST.	Ell	E 4 4	E08	E42	E 22	E02	EOI	W63	W13	M77	W15	E30	W34	W28	W33	W3I	
	APP	LAT.	N13	S09	40N	S12	6 0 N	N05	50N	516	20N	S15	6 O N	N29	S10	S10	508	NII	
		MAX. PHASE	0704				1500			0742		1039	1051	1049	0158			0914	
OBSERVED	UNIVERSAL TIME	END	0710	1055	1212	1440	1505	0758 D	1101	0450	0.952 D	1052	1102	1106	0216	0205 D	0207 D	0919	
		START	0 7 0 0	1025	1200 E	1433	1452	0745	1045 E	0739	0947 E	1034	1050	1045	0155	0158 E	0158 E	0914	
DATE		0CT 1959	25	25	25	25	25	26	26	27	27	27	27	28	31	31	31	31	
	VOCTENENSES	OBSERVAION	PIRCULI	LOCARNO	CAPRI G	LOCARNO	LOCARNO	CAPRI G	ONDRE JOV	GOOD HOPE	CAPRI G	GOOD HOPE	GOOD HOPE	KRASNYA	SYDNEY	KYOTO	ίκγοτο	KRASNYA	

SAC PEAK: ALL VALUES IN MAX, INT, COLUMM ARE ARBITRARY UNITS (O-40), NOT PERCENT OF CONTINUOUS SPECTRUM, E - LESS THAN & - PLUS D - GREATER THAN MINUS U - APPROXIMATE □ - NOT REPORTED	LOCKHEED OBSERVATIONS: ALL VALUES IN THE MAXI- MUM INTENSITY COLUDAN ARE ARBITRARY UNITS ON A SCALE OF 1 TO 4 - NOT PERCENT OF THE CONTINUOUS SPECTRUM.
MOSCOW - GAISH ROYAL OBSERVATORY, EDINBURGH GREENWICH ROYAL OBSERVATORY, HERSTMONCEUX SACRAMENTO PEAK SCHAUINSLAND SCHAUINSLAND UNITED STATES NAVAL RESEARCH LABORATORY	
MOSCOM-G R O EDIN R O HERST SAC PEAK SCHAUINS USNRL	
AMACAFRI - GERMAN ANACAPRI - GERMAN ANACAPRI - SWEDISH ROYAL OBSERVATORY, CAPE OF GOOD HOPE KIEV UNIVERSITY KODALKAMAL KODALKAMAL KRASNAYA PAKHRA LOS ANGELES	
CAPRI G CAPRI S GOOD HOPE KIEV* KODAIKNAL KRASNYA LOCKHEED	

INTERVALS OF NO FLARE PATROL OBSERVATIONS OCTOBER 1959



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IONOSPHERIC EFFECTS OF SOLAR FLARES

(SHORT-WAVE RADIO FADEOUTS)

DECEMBER 1959

Dec. 1959	Start UT	End UT	Туре	Wide Spread Index	Impor- tance	Observation Stations	Known Flare, UT CRPL-F 185
1 1 1 2 2	1247 1512 1705 0330 0500	1327 1613 1900 0450 0550	S-SWF Slow S-SWF S-SWF S-SWF Slow S-SWF	1 5 5 1 1	3 2+ 3 3- 2	<u>JU</u> BE, FM, HU, LA, MC, NE, <u>PR</u> , WS, CW* * AN, BE, FM, HU, LA, <u>MC</u> , NE, PR, WS, CW*** <u>OK</u> <u>OK</u>	1208E 1456E 1638 * *
2 3 3 3 4	1246 1017 1414 1757 0028	1402 1047 1500 1903 0208	S-SWF S-SWF S-SWF S-SWF S-SWF	5 1 3 5 5	2+ 2 2- 2+ 2+	BE, DA, HU, MC, NE, <u>PR</u> , PU, SW, CW [*] _* <u>NE</u> HU, <u>PR</u> .AN, BE, <u>FM</u> , HU, LA, MC, NE, PR, WS, CW* AD, <u>OK</u>	1219E 1010 1408E 1756 0032
4 4 5 5	0210 0800 1820 0620 1003	0240 0840 1920 0633 1022	G-SWF - S-SWF S-SWF S-SWF	1 5 1 3	1+ - 2+ 3 2	<u>OK</u> <u>C₩</u> ** BE, FM, HU, LA, <u>MC</u> , PR, WS <u>KO</u> NE, <u>PU</u>	0208 * 1814 * 1004E
5 5 7 7 8	1220 1615 0440 1042 0118	1232 1630 0520 1107 0200	S-SWF S-SWF S-SWF S-SWF S-SWF	5 5 5 5 5	2 1 2- 1 1+	BE, NE, <u>PR</u> , PU <u>BE</u> , MC, PR, WS CA, KE, <u>OK</u> PR, <u>PU</u> AD, CA, <u>OK</u>	1230E 0434 * 0120E
8 10 10 11 12	0755 0518 2355 0407 0800	0820 0545 0037 0430 0810	S-SWF S-SWF Slow S-SWF S-SWF -	5 5 1 1	1+ 1+ 2 1+ -	CA, KO, OK, NE CA, KO, OK AD, CA, OK OK CW++	* * * 0812
17 18 19 24	0400 0637 0345 0343	0422 0657 0405 0450	Slow S-SWF S-SWF S-SWF Slow S-SWF	4 4 1 1	1+ 1+ 1 2	$ \begin{array}{l} \underline{OK}, & TO \\ \overline{KO}, & \underline{OK} \\ \underline{OK} \\ \underline{OK} \end{array} $	* * *
CA =	Canbern	ra, Aus	tralia		PU = Pr	ague, Czechoslovakia commercz -	STANDARDS - BOULDER

DA = Darmstadt, G.F.R. JU = Juhlesruh, G.D.R.

TO = Hiraiso Radio Wave Observatory, Japan

DA = Darnstadt, G.F.R.TU = Hiraiso Kadio wave Observatory, JapanJU = Juhlesruh, G.D.R.SW = Enkoping, SwedenKE = KerguelenCW* = Cable and Wireless, BarbadosK0 = Kodaikanal, IndiaCW** = Cable and Wireless, Somerton, EnglandLA = Los Angeles, Calif.CW*** = Cable and Wireless, Brentwood, EnglandNE = Nederhorst den Berg, NetherlandsCW++ = Cable and Wireless, Singapore

IONOSPHERIC EFFECTS OF SOLAR FLARES

(Sudden Cosmic Noise Absorption Sudden Enhancements Of Atmospherics) Solar Noise Bursts At 18 Mc. SEPTEMBER 1959

SEPT. 1959	SCN A	CLASS SEA	Burst	WIDESPREAD INDEX	(UN BEGIN	TIME IVERSAL T MAX.	ME) END	PERCENT ABSORPTION SCNA	OBSERVATION STATIONS
1 1 1 1	1 3	2	1+ 2	1 3 5 5 5	0206 1420 1655 1656 1730	0211 1425 1708 1701	0225 1428 1735 1800U 2355	15 80	<u>HA</u> <u>MC, RE</u> <u>BO</u> , DU, HA, NE, PA, SP <u>BO</u> , HA, MC, RE, SP <u>BO</u> , HA, MC, SP, (Noise storm, peaks at 1810-1830 (RE) 1950- 2007 (RE) 2228-2307)
1 2 2 2 2	2	2+ 2 1 1+	1	3 3 5 5 5	1928 2007 1311 1605 1606 1607	1940 2022 1615 1610	2007U 2100 1329 1607 1640 1635	55	A1, A3, <u>A5</u> A1, A3, <u>A5</u> <u>JU</u> , KU <u>B0</u> , MC, RE, SP A3, A5, <u>B0</u> , DU, NE, PA, SP <u>B0</u> , MC, RE, SP
2 2 3 3	1	1	1 2 2	4 1 5 1 4	1730 1735 1800 1425 1739	1747 1752	1810 1810 0026 1605 1900	15	BO, MC BO BO, <u>HA</u> , MC, (Noise Storm) <u>MC</u> , (Group of bursts) BO, MC, SP (Group of bursts)
3 4 4 5		1+	1 1 2 2	1 3 5 1 5	2115 1605 1700 2320 1554	1632	0030 1710 2345 2326 1600		<u>HA</u> (Noise Storm) A2, <u>A5</u> <u>BO</u> , HA, (Noise Storm) <u>HA</u> BO, <u>MC</u> , RE, SP
6 6 7 9		2	1 2 1	1 1 1	2045 2158 1615 0703		2048 2205 1623 0741		HA HA MC NE
9 9 9 9	1	2	1 1 1	5 5 5 4	1557 1558 1620 1655 1711	1605 1611	1625 - 1625 1658 1713	20	<u>BO</u> , MC, RE, SP <u>BO</u> , MC, NE, SP <u>BO</u> , MC, RE, SP <u>BO</u> , MC, SP <u>BO</u> , MC
9 9 11 12			1 1 1 1	4 5 1 5 5	1824 1833 2235 2000 2038		1827 1839 2240 2001 2050		BO, <u>MC</u> BO, <u>MC</u> , SP <u>MC</u> <u>BO</u> , HA, MC <u>BO</u> , HA, MC, SP
13 14 14 14 15	1	1	1	5 1 5 5 5	1725 0745 1833 2156 2107	1740 2200	1800 0821 18 35 2210 2109	10 25	<u>BO</u> , HA, MC <u>NE</u> <u>BO</u> , MC, SP BO, <u>HA</u> , SP <u>BO</u> , HA
16 16 19 19 19	1	1	1 1 1	4 5 5 1	1846 1846 1939 2027 2203	1850	1902 - 1948 2030 2208	15	<u>BO</u> , MC <u>BO</u> , HA, MC <u>BO</u> , HA, SP,(Group of bursts) <u>BO</u> , HA, MC, SP <u>HA</u>
19 20 20 21 22 23	1		1. 1 1 1 1	1 5 1 4 1	2211 2147 2255 0121 1652 1909	2300 1915	2218 2148 2312 0130 1658 1917	10	<u>На</u> <u>ВО</u> , НА, МС <u>НА</u> <u>НА</u> <u>ВО</u> , МС <u>RE</u>
25 25 25 26 29			1 2 1 1 1	1 5 4 5 5	1444 1609 1625 1830 2012		1451 1617 1635 2330 2015		MC BO, MC, RE BO, MC BO, HA, MC, SP,(Noise Storm) BO, HA
29 29			1 1	5 1	2031 2103		2035 2105		<u>BO</u> , HA <u>HA</u>

SOLAR RADIO EMISSION

OUTSTANDING OCCURRENCES

Ottawa

JANUARY 1960

2800 Mc

Jan.	Туре	Start UT	Duration	Maxim	III).	Remarks
1960			Hrs:Mins	Time UT	Peak	
					Flux	
3	2 Simple 2	1725	2	1725.7	12	
8	3 Simple 3	1805	2 30	1900	15	
10	1 Simple 1	1617.5	2	1618	7	
11	2 Simple 2	2056	>35	2108	220	In sunset oscillations
12	6 Complex f	1647.3	9	1649	80	
1	-					
13	6 Complex f	1446	7	1449.5	18	
	4 Post Increase		15		6	
13	2 Simple 2	1847	4	1847.8	30	(
15	6 Complex f	b1340*	>1 40	1357	700	*In interference
1	4 Post Increase A		2 30		25	1
	2 Simple 2 f	1730.5	8	1732	300	
1			1			
16	3 Simple 3	b1543	>2 17	1620	10	1
17	3 Simple 3 A	b1600	>3 30	indet,	12	
1	3 Simple 3	1616	-40	1619	8	1
19	3 Simple 3 A	1925	>2 10	2007	20	
	6 Complex	1936.5	25	1945	65	
22	2 Simple 2	1649	1.5	1649.3	10	1
23	1 Simple 1	1519.5	2	1520.2	6	
24	1 Simple 1	1633	1.5	1633.7	6	
24	l Simple 1	1715.3	1	1715.5	7	
25	2 Simple 2	1712.5	3.5	1714.2	90	
	4 Post Increase		1 15		6	
30	3 Simple 3 f	2015	12	2021	10	
L						l

NINEACE - STANDARDS - BOULDER

SOLAR RADIO EMISSION OUTSTANDING OCCURRENCES JANUARY 1960

Jan. 1960	Туре	Start UT	Time of Maximum UT	Duration Minutes	Intensity
3 3 3 3 3	3 3 3 3 3	1607.1 1923.4 2133.0 2225.4 2258.3	1607.2 1923.4 2133.0 2225.6 2258.8	1.2 0.1 0.2 0.3 0.7	1 1 2 1**
3 4 4 4	3 6 3 7	2320.5 2325.8 1423 E 1806.9 1906	2320.5 2325.8 1806.9	0.1 0.2 165 D 0.1 97	1** 1** 2 1 2
5 5 6 7	3 2 2 3	1916.5 1919.6 1500.9 2319 2224.0	1916.5 1919.6 1502.8 2320 2224.0	0.1 0.2 4.1 12 0.1	2 2 1* 1** 1
7 7 8 9	3 3 6 6	2302.3 2305.4 2307.6 1423 E 1422 E	2302.3 2305.4 2308.3	0.1 0.2 1.2 554 D 458 D	1** 2** 2** 1 1
10 10 10 10 10	3 3 3 3 7	1836.0 1842.3 2021.5 2143.7 2229	1837.1 1842.3 2021.5 2143.7	2.0 0.3 0.1 0.2 71 D	2 2 1 2 2
11 11 12 12 13	6 9 6 8 6	1421 E 2056.0 1422 E 1648.9 1421 E	1935 U 1651.2 1957 U	562 D 167 D 561 D 12 562 D	2 3** 2 3 1
14 15 16 16 17	6 6 8 6	1421 E 1418 E 1422 E 2247 1421 E	1756 U 2250.0	564 D 567 D 565 D 6 566 D	2 1 2 3 1
17 18 18 18 18	2 2 3 3	1910.0 1705.0 1739 1757.0 1912.8	1911.0 1705.6 1740 1757.0 1912.8	1.6 1.0 7 0.2 0.3	2 2 2 2 2
18 19 19 19 19	3 3 3 3 3	2330.2 1431.6 1433.8 1455.8 1640.0	2330.2 1431.7 1433.8 1455.8 1640.0	0.2 0.8 0.1 0.3 0.1	2** 2* 2* 1

				2	167 MC
Jan. 1960	Type	Sta rt UT	Time of Maximum UT	Duration Minutes	Intensity
19 19 19 20 20	7 9A 9B 2 3	1849 1941.0 1953 1432.0 1802.0	1946.2 1956.5 1433.1 1803.0	300 D 12 8 2.2 1.2	1 2 1 2* 2
20 21 21 21 21 21	3 3 3 3 3	1813.0 1458.0 1500.2 1602.0 1739.9	1813.0 1458.0 1500.2 1602.0 1739.9	0.1 0.1 0.2 0.1	1 1* 1* 1
21 22 22 22 24	3 3 3 3 3	2337.7 1423.0 1550.5 1805.4 2003.5	2337.7 1423.0 1550.5 1805.4 2003.5	0.2 0.1 0.2 0.2 0.3	1** 1* 2 2 2
26 26 26 26 27	3 3 3 2	1422.0 1425.0 1507.2 1643.9 1417.8	1422.2 1425.0 1507.2 1643.9 1417.8	0.1 0.2 0.4 0.6 3.2	1* 1* 2 3*
27 27 27 27 27 27	3 3 3 3 3	1428.0 1504.7 1531.3 1543.0 1805.5	1428.0 1504.7 1531.3 1543.0 1806.0	0.2 0.5 0.4 0.8 0.7	2* 2 1 1 2
27 27 27 27 27 28	3 3 3 3 3	1809.0 1838.5 1852.0 2348.8 1432.5	1809.0 1839.0 1852.5 2348.9 1432.5	0.2 1.2 1.0 0.8 0.5	2 1 2** 2*
28 29 30 30 30	3 3 3 2	2334.5 1823.2 1546.0 1640.0 1714	2334.5 1823.2 1546.0 1640.3 1714.0	0.3 0.2 1.0 0.4 7	2** 2 2 2 2
30 30 30 30 30	3 3 3 3 3	1723.0 1836.0 2116.6 2201.0 2348.6	1723.0 1836.0 2116.6 2201.0 2348.6	0.5 0.9 0.4 0.2 0.3	3 1 2 3 1**
31 31 31 31 31	3 3 3 3 3	1423.0 1451.2 2002.0 2350.9 2354.1	1423.0 1451.2 2002.0 2351.2 2355.2	0.2 0.1 0.8 1.6 2.8	1* 2* 2** 2**

* On sunrise pattern ** On sunset pattern

BOULDER

TIMES OF OBSERVATION

Jan. 1960	U.T.	Jan. 1960		
1 2 3 4 5 6 7 8 9 10 11 12 13 14 15	$\begin{array}{c} 1422-2329\\ 1424-2333\\ 1424-2334\\ 1423-2335\\ 1422-2335\\ 1422-2337\\ 1423-2337\\ 1423-2337\\ 1422-2200\\ 1730-2340\\ 1421-2343\\ 1422-2047\\ 2103-2343\\ 1421-2343\\ 1421-2343\\ 1421-2345\\ 1418-2345\\ \end{array}$	16 17 18 19 20 21 22 23 24 25 26 27 28 29 30 31	$\begin{array}{c} 1422-2347\\ 1421-2347\\ 1645-2348\\ 1419-2349\\ 1419-2350\\ 1417-2352\\ 1415-2354\\ 1416-2356\\ 1416-2356\\ 1416-2356\\ 1415-2357\\ 1415-0000\\ 1415-0000\\ 1415-0000\\ 1415-0002\\ 1410-0004\\ 1420-0005\\ 1410-0006\\ \end{array}$	I after 1830 I 1915-2030 I throughout day I throughout day I 1830-0004 I 1725-1745



JOHNERCE - STANDARDS - BOULDER

GEOMAGNETIC ACTIVITY INDICES

DECEMBER 1959

Dec. 1959	С	Values Kp Three hour Gr. interval 1 2 3 4 5 6 7 8	Sum	Ар	Final Selected Days
1 2 3 4 5	1.1 1.2 1.5 0.7 1.7	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	31- 33+ 40+ 210 36-	28 30 50 12 68	Five Quiet 7 10
6 7 8 9 10	0.9 0.3 0.3 0.3 0.2	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	24+ 120 14- 16- 12-	16 6 7 8 6	11 21 22
11 12 13 14 15	0.2 1.0 1.1 1.4 1.1	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	120 240 230 370 310	6 16 18 40 26	Five Disturbed 3 5
16 17 18 19 20	1.0 0.3 0.5 0.8 0.3	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	250 15+ 18+ 22 - 140	17 8 10 14 6	14 27 28
21 22 23 24 25	0.1 0.4 1.3 0.9 0.4	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	8- 110 29+ 22- 17-	4 6 28 15 8	Ten Quiet 7 8 9
26 27 28 29 30 31	1.2 1.4 1.4 0.9 0.8 0.4	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	31- 36+ 370 250 22- 14-	26 38 40 18 14 7	10 11 17 20 21 22 31
Mean:	0.81		Mean:	19	

DAYS IN SOLAR ROTATION INTERVAL 16 17 18 19 20 21 22 23 24 25 26 27 14 15 ROT.= 1959 1717 Jon De 1718 Fet Jan 1719 Feh Mch 1720 Mch AD. 1721 Apr Apr 1722 Ap May 1723 Mai 1724 Ju 1725 Aug Jul 1726 Aug Sar 1727 Sep 1728 De Nou 1729 Nov D/ 1730 1101 ------Dec Dec 28 1731 **1**.11, P ÷Ĥ 1960 , t -----Ť, Dec 2.9 31 Jan Jan KEY PLANETARY THREE - HOUR - RANGE INDICES Kp 🛦 = sudden FOR THE INTERNATIONAL GEOPHYSICAL COOPERATION commencement 1958 December 16 to 1959 December 31 (and preliminary indices to 1960 January 20) 1 + - 2 + - 0 + - 0 + - 0 + - 0 + - 0 + - 0 + - 0 + - 0 0 + -

CRPL RADIO PROPAGATION QUALITY FIGURES AND FORECASTS NORTH ATLANTIC

DECEMBER 1959

Dec. 1959	North Atlantic 6-hourly quality figures	Short-term forecasts issued about one hour in advance of:	Whole day index	Advance forecasts (J-reports) for whole day; issued in advance by:	Geomag- netic K _{Fr}
	00 06 12 18 to to to to 06 12 18 24	00 06 12 18		1-7 1-7 1-7 1-7 days days days days Final Js SDW J	Half Day (1) (2)
1 2 3 4 5	50 5- 6+ 50 5+ 50 6- 40 5- 4+ 6- 4+ 5- 50 70 6+ 6+ 6- 60 5-	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	5+ 5- 5- 6- 6-	4 4 5 5 4 4 5 5 5 5 5 5	(4) 3 (4) (4) (5) (4) 3 1 2 (5)
6 7 8 9 10	5+ 50 70 6+ 6- 60 70 7- 6+ 6+ 7+ 6- 5+ 6+ 70 70 7- 60 70 6+	4 4 5 6 5 5 7 7 6 6 7 7 6 6 7 6 6 6 7 7	60 6+ 6+ 6+ 6+	6 6 7 7 7 7 7 7 6 6	$\begin{array}{ccc} 3 & 2 \\ 1 & 1 \\ 1 & 2 \\ 2 & 2 \\ 1 & 1 \end{array}$
11 12 13 14 15	6+ 6- 7- 7- 60 6+ 70 60 5- 60 7+ 5+ 4+ 5- 60 5+ 5+ 4+ 6+ 4+	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	6+ 6+ 50 50	6 6 7 7 7 7 7 7 7 7 6 6	$ \begin{array}{rrrrr} 1 & 1 \\ 3 & 3 \\ 3 & 3 \\ 3 & (4) \\ (4) & 3 \end{array} $
16 17 18 19 20	50 5+ 60 4+ 50 6- 7+ 5- 50 6- 70 7- 5+ 5+ 7- 6+ 50 5+ 7+ 6+	4 4 6 5 5 6 7 6 5 5 7 6 6 5 7 7 5 5 7 7	50 6- 60 6- 60	6 6 6 6 6 6 6 6 5 5	3 2 2 1 2 2 3 1 2 1
21 22 23 24 25	6- 60 70 7- 6+ 60 70 6- 5- 50 70 4+ 5- 50 7- 60 50 6- 7- 6+	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	6+ 6+ 5+ 6- 60	5 5 6 6 6 6 4 4 5 5	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$
26 27 28 29 30 31	6- 6- 7- 50 50 50 6+ 4+ 5- 4+ 6- 5- 50 50 6+ 6- 50 6- 6- 50 5- 5+ 7- 6-	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	6- 50 5- 6- 5+ 6-	6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6	$\begin{array}{cccc} (4) & 3\\ (4) & (4)\\ (4) & (4)\\ (4) & 2\\ 3 & 2\\ 2 & 1\\ \end{array}$
Score	: Quiet Periods	P 13 14 27 9 S 17 14 3 16 U 0 0 1 0 F 0 0 0 0		$\begin{array}{cccc} 11 & 11 \\ 18 & 18 \\ 1 & 1 \\ 1 & 1 \\ 1 & 1 \\ \end{array}$	
D	isturbed Periods	P 0 1 0 0 S 1 2 0 5 U 0 0 0 0 F 0 0 0 1		0 0 0 0 0 0 0 0 0 0	

() represent disturbed values.

CRPL RADIO PROPAGATION QUALITY FIGURES AND FORECASTS NORTH ATLANTIC



COMMERCE - STANDARDS

BOULDER

DECEMBER 1959









CRPL RADIO PROPAGATION QUALITY FIGURES AND FORECASTS

NORTH PACIFIC

DECEMBER 1959

Dec. 1959	North Pacific 12-hourly quality figures		Short-t casts i	erm fore- ssued at	Whole day index	Adva (Jp whol in	nce f repor e day adva	oreca ts) ; is ince l	asts for sued by:	Geom net K _S	ag- ic i
	0700 to 1900	1900 to 0700	0600	1800		l - 7 days Final	1 - 7 days Jps	l-7 days SDW	l-7 days Jp	Half (1)	Day (2)
1 2 3 4 5	6 6 4 6 3	6 6 4 5 2	4 4 5 5 5	5 5 5 5 5	6 5 (4) 5 (2)	4 5 5 5 6			4 5 5 5 6	(5) (4) (5) 3 2	(4) (4) (6) 2 (8)
6 7 8 9 10	6 5 5 6 5	6 5 6 5 5	5 6 5 5 5	6 5 5 6 5	6 5 5 6 5	6 6 6 5			6 6 6 5	3 1 1 2 1	2 2 1 1 1
11 12 13 14 15	5 7 6 5 5	4 5 6 5 5	5 5 5 4	5 5 5 5 5	5 6 (4) 5	5 6 6 6 6			5 6 6 6	1 2 (5) (4)	1 2 (4) (4)
16 17 18 19 20	5 6 7 6 6	5 6 7 6 5	5 5 5 5	5 6 6 6	5 6 7 6 5	5 5 6 5			5 5 6 5	3 2 2 (4) 1	(4) 2 2 3 2
21 22 23 24 25	6 5 5 6 6	6 6 5 7	5 5 5 5 5	5 5 6 6	6 6 6 6	5 6 5 6			5 6 5 6	0 0 2 (4) 2	0 2 (4) 2 2
26 27 28 29 30 31	5 5 4 5 6 6	6 5 5 6 6	6 5 4 5 5 6	5 5 4 5 6 6	6 5 (4) 5 6 6	6 5 5 5 5 6			6 5 5 5 5 6	(4) (4) (4) (4) 2 1	(4) (4) (5) 3 3 2
Score: Quiet Periods		P 10 S 14 U 2 F 2	14 14 0 0		18 7 1 1						
Disturbed Periods		P 1 S 1 U 1 F 0	0 2 0 1		0 2 0 2						

() represent disturbed values.

COMMERCE - STANDARDS - BOULDER

Errata: In CRPL-F 185 Part B the score for December 1959 under "Whole day index" should have been under "1-7 days Final".

NORTH PACIFIC DECEMBER 1959

OUTCOME OF ADVANCED FORECASTS FINAL ESTIMATE



ALERT PERIODS AND SPECIAL WORLD INTERVALS

INTERNATIONAL WORLD DAY SERVICE JANUARY 1960

Issued Day/Time UT Jan 1960	Advance Geophysical Alert	No.	World-Wide Geophysical Alert	Special World Interval
11/1600		44	Magnetic Storm 10/0719Z	
12/0000	Burbank Solar Flare 11/2140Z			
14/1005	Fort Belvoir Magnetic Storm 13/19002			
14/1600		45	Magnetic Storm 13/1900Z	
18/1115	Fort Belvoir Magnetic Storm 17/12XXZ			
18/1600		46	Magnetic Storm 17/1200Z	
21/1600		47	Magnetic Storm 21/00XXZ	

