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An Alinement Interferometer

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This paper describes a diverging-beam type of alinement interferometer that permits the use of large apertures, provides ruggedness, high sensitivity, and is relatively compact. Since this interferometer is fully compensated, white light can be used.

The narrowness of the zero order fringe, relative to separation of fringes, in such a well compensated system permits settings to better than one twentieth of a fringe width. This corresponds to a lateral displacement of the light source of less than 1.5 mm at a distance of one mile for a one-inch aperture instrument and to less than 0.15 mm for a 10-inch aperture instrument.

The increase in dmands for higher accuracy in metrology includes that of pointing. The alinement interferometer described here is a precision device that is rugged and relatively easy to apply. Its performance, because of its high sensitivity, is severally affected by the homogeneity of the atmosphere or medium through which the light passes. However, this effect is also greatly reduced by the proximity of any two interfering rays of light.

There appears to be a growing need for an accurate method of alining objects over distances of several miles. The diffraction methods, described by Van Heel¹ and Dyson², appear to be inadequate for such distances. Other methods such as the axicon³ and the telescope also have limited accuracy. The alinement interferometer described here is believed to be practical and more sensitive than previously described methods.

An arrangement of optical elements that form a precision alimement interferometer is shown in figure 1. It consists of two double-image prisms ⁴, P_1 and P_2 , a beam divider, B, a lens, L, a source, S, and two pairs of quarter-wave plates, C_1 and C_2 . Figure 2A shows the course of the light from B to the observer located either at F'_1 or F''_1 for measurements in the plane of figure 1, and figure 2B shows the same for measurement in the plane perpendicular to that of figure 1.

The alinement problem may consist of determining the position of the source, S, located a distance, d, from the optic axis of L, the precise adjustment of S

¹ Van Heel, A.C.S., Progress in optics **1**, 291 (1961). ² Dyson, J., Optics in metrology, p. 169 (Pergamon Press, Pol. Mollet Ed., 1660 on the axis of L, or the alimement of several sources along the axis of L. Obviously, if the problem is to determine the position of S relative to one plane only (the plane of fig. 1, for instance), then P_2 and B may be eliminated.

The optical sensitivity of this interferometer is proportional to the aperture. If the aperture of the prism is sufficient, the lens may be eliminated and the entrance tace of the prism may be plane or spherical. A plane face prism, used without the aperture-increasing lens L, requires a collector lens as shown in figure 3. The 2d prism and beam divider are not shown in figure 3. They are, however, necessary for alinement in the plane perpendicular to this figure.

This prism is difficult to make in large sizes. Small prisms are relatively easy to construct and may be used with a lens to increase the aperture of the interferometer. Aperture sizes are limited only by the difficulty of producing large objectives (lenses or mirrors). A mirror objective is as easily used as a refractor and is free from chromatic aberration.

The equation of sensitivity is obtained from figure 4, in which the aperture is represented by the aperture of a lens, L. Consider two rays of light that emerge from the lens at points A_1 and A_2 that are equally





The dividing plane of the double-image prism is a semi-reflecting optical film. It serves as the beam combiner in this interferometer.



FIGURE 2. Figure 2A represents the path of the light in the plane of figure 1. Figure 2B represents the same in a plane perpendicular to the plane of figure 1.

The intersection of these two planes represent the straight line of interest.



FIGURE 3. Optics of an alinement interferometer.



FIGURE 4. Rays of light, ρ_1 and ρ_2 , are combined in the prism. Consequently, points A_1 and A_2 appear to coincide at the reference points, indicated in figures 5 and 6.

distant (y and -y) from A_0 , which lie in the extension of the dividing plane of prism P_1 . The distance X(approximating the distance from lens to source), shown in figure 4, is measured from A_0 . The plane of figure 4 is assumed to include the optic axis of L. Let ρ_1 and ρ_2 equal the optical paths from the source to the observer (at F'_1 or F''_1 , fig. 2) through A_1 and A_2 , respectively. The optical path from A_1 to the observer is identical to that from A_2 , because of symmetry in the prism. The observer sees images of A_1 and A_2 superimposed, as shown in figure 5. The observed order of interference, N, at this pair of points, multiplied by the wavelength of the light, λ , is equal to $(\rho_1 - \rho_2)$. Since X is assumed quite large relative to y and d, we can ignore the refractive index of the lens and obtain,

$$\rho_1^2 = X^2 + (y - d)^2, \tag{1}$$

$$\rho_2^2 = X^2 + (y+d)^2.$$
 (2)

On subtracting eq (1) from (2), and expanding, we readily obtain

$$\rho_2^2 - \rho_1^2 = (\rho_2 - \rho_1)(\rho_2 + \rho_1) = 4yd.$$

On putting $(\rho_2 + \rho_1) = 2X$ (which it approximates for large values of X) and solving for d we obtain

$$d = \frac{2X \cdot N\lambda}{4y}$$

If A equals the aperture of the interferometer and the reference point is chosen at or near the margin then $y = \frac{1}{2}A$ and

$$d = \frac{XN\lambda}{A} \tag{3}$$



FIGURE 5. The appearance of interference fringes when the point source is: A, above the line; B, on the line; and C, below the line of interest.

If the problem is to adjust S on the axis, and the prism is adjusted to place the zero order of interference on the reference point where d=0, then the adjustment consists in moving S perpendicular to the axis until N equal zero. If white light is used with a reflecting objective (a parabolic mirror), the zero order of interference appears to have excellent contrast and an experienced observer can adjust the fringes to an accuracy of less than 0.05 fringe spacing. The corresponding error in N is 0.05 and for $\lambda = 5.0$ $\times 10^{-5}$ cm (the approximate effective value for white light sources), the corresponding error in d is X/A $\times 2.5 \times 10^{-6}$ cm. This corresponds to an error in d/. (the angular position of S) of $\frac{2.5}{A} \times 10^{-6} (A \text{ ex-}$ pressed in cm). If A equals 6 in. and X equals 1 mile, the corresponding error in d is approximately 0.3 mm.

Two photomultiplier tubes placed in a differential arrangement can detect changes in the position of S to much less than this. Obviously, this principle can be used to guide an astronomical telescope when mounted on a satellite (above the atmosphere of the earth) to very high precision.

Visual observations are most easily made with a prism that has a "built-in-wedge" of convenient magnitude to produce interference fringes whose widths are most appropriate for precision of readings. Prisms may easily be adjusted to provide fringes that are neither too broad nor too narrow. When photoelectric recording is to be used for null settings the prism should be adjusted to have no built-in wedge. This permits the use of the entire beam for detection.

Figures 5A, 5B, and 5C respectively show the appearance of the fringes in a built-in wedge interferometer of 4 in. aperture when S is one second of arc above, on, and one second below, the dividing plane of the prism through which observation is made.

Because of differential polarization of light in the prism, unpolarized light cannot be used without properly oriented quarter-wave plates or a polarizer. Use of polarized light results in a loss of one-half the available light. The polarizer is used by placing it either between the prism and observer or near the source. If a polarizer is adjusted to transmit light vibrating parallel to the dividing plane of the prism, and if the observed fringe pattern is represented by figure 6A, then the pattern will change to that of figure 6B when the polarizer is rotated 90° (light vibrating perpendicular to the dividing plane of the prism). Therefore, unpolarized



FIGURE 6. The appearance of interference fringes when: A, the light is polarized parallel to the dividing plane of figure 1; B, the light is polarized perpendicular to the plane of figure 1; and C, when quarter-wave plates are used instead of a polarizer.



FIGURE 7. Arrangement of quarter-wave plates to give good contrast with unpolarized light.

light produces two sets of fringes with the maximum intensity areas of one set coinciding approximately, with the minimum intensity areas of the other set, thus eliminating contrast so that no fringes can be observed.

Dr. Charles J. Koester ⁵ suggested and demonstrated that by cutting a quarter-wave sheet ⁶ of mica at 45° to its optic axis and mounting it in front of the prism, in the manner shown in figure 7, the relative phase changes in the two polarized beams are practically equalized. The two sets of fringes produced by the two component beams (vibrating parallel and perpendicular to the dividing plane of the prism) will then coincide in phase. Good contrast is thereby obtained with little loss of light. The resultant fringe pattern (fig. 6C) indicates that the two sets of fringes, shown in figures 6A and 6B, are shifted in opposite directions by an amount that places them in coincidence.

Tests made in the laboratory showed that the sensitivity of the instrument, obtained from eq (3), agreed within the limits of observations. The theoretical accuracy claimed above requires that the medium through which the light travels be quite homogeneous. An evacuated light tube would probably be required to obtain the accuracy indicated by eq (3) when measurements are made over long distances.

⁵ Working at NBS under a National Research Council Postdoctoral Research Associateship and now with American Optical Co., Southbridge, Mass. ⁶ Subsequent tests showed that the relative retardation is not critical for obtaining good contrast.