# Absorption Spectrum of Water Vapor Between 4.5 and 13 Microns<sup>1</sup>

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The absorption spectrum of water vapor has been measured from 4.5 to 13 microns with a 3,600 line per inch replica echelette grating as the dispersing element. Various absorbing path lengths and concentrations of water vapor at atmospheric pressure were used up to 8 meters of steam near  $110^{\circ}$  C. Almost all of the previously unreported lines that have been found are also present in the solar spectrum. A rotational analysis shows that most of the lines can be assigned either to rotational transitions or to rotation-vibration transitions of the  $\nu_2$  fundamental of the water-vapor molecule. In addition, a few lines have also been assigned to the transitions  $(2\nu_2 - \nu_2)$ ,  $(\nu_1 - \nu_2)$ , and  $(\nu_3 - \nu_2)$ .

## 1. Introduction

Investigations of the infrared spectrum of sunlight between 7 and 13  $\mu$  have shown that essentially all the structure is due to absorption by polyatomic gases that are in the earth's atmosphere. In addition to the known bands of ozone, carbon dioxide, water vapor, nitrous oxide, and methane that have been identified in this region, there is also a large number of irregularly spaced absorption lines. Some of these lines are very strong. In solar spectra taken at Columbus, it has been observed that most of their intensities are dependent on the amount of water vapor in the atmosphere. One of us (W.S.B.) had previously found that many of these lines observed in the grating map of the solar spectrum published by Adel [1]<sup>5</sup> could be predicted from the known energy levels of the water-vapor molecule. To obtain a positive identification of these lines, it was thought desirable to observe them in a laboratory spectrum. If these lines are due to water vapor, most of them must originate from the higher rotational energy levels of the molecule, which have a low population at room temperature. Thus, in order to observe them in the laboratory, an absorbing layer giving an optical path length approaching that of the water vapor in the atmosphere is required. In this work a long path length was obtained by using a multiple-reflection cell similar to that described by J. U. White [2]. From this laboratory spectrum it has been possible to identify many lines in the solar spectrum with watervapor lines and also to evaluate rotational energy levels of the  $\nu_2$  band and of the pure-rotation band, which are in good agreement with levels derived from other observations.

# 2. Apparatus and Experimental Procedure

The absorption cell used in this investigation is shown schematically in figure 1. It consists of a large cast-aluminum tank wound with heating coils

and covered with a layer of asbestos for insulation. One end is held in place by bolts and can be completely removed for adjusting the optical system. Three spherical mirrors, each of 100-cm radius of curvature, are used in the optical system. Mirror  $M_3$  is fixed, but an external control is provided for varying the inclination of the optic axes of mirrors  $M_1$  and  $M_2$ . By means of this adjustment the path length of radiation passing through the cell can be changed in steps of 4 m. Figure 1 shows the arrangement for a path length of 8 m.

Rocksalt windows could not be used in this investigation, and it was found that silver-chloride windows reduced the energy very considerably. Satisfactory results were obtained by placing the Nernst filament, used as a source, inside the cell near mirror  $M_3$  and leaving the exit window open. Radiation was reflected out of the cell and through this window by a small plane mirror.

Before introducing the steam, the cell was heated to 110° C to prevent condensation on the mirrors. The cell was filled with steam by displacement. When running spectra, a continuous stream of vapor from a gently boiling flask of distilled water entered at one end and escaped through the exit window.

For mapping regions of high absorption, air was allowed to enter the cell until a suitable concentration of water vapor was obtained. Near the center of the  $\nu_2$  fundamental the atmospheric water vapor in the spectrometer alone was sufficient to produce very intense absorptions. The region from 5.7 to  $6.7 \mu$  has been previously mapped by H. H. Nielsen [3] with very high resolution and with small amounts of water vapor. The present workers were unable to improve upon these results, and this region has not been remeasured. In the solar spectrum no energy is observed between 5.5 and 7.5  $\mu$  at Columbus during most of the year.

The spectrometer with which these measurements were made has already been described [4]. For this work a 3,600 line/in replica echelette grating (width 5.5 in., height 3.5 in.) was used, together with a Perkin-Elmer thermocouple (rated sensitivity 8 v/w) detector and their 13 c/s chopping system and amplifier. All spectra were recorded in the first order of the grating. Portions of the spectrum at short wavelengths were also scanned in the second

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FIGURE 1. Multiple reflection cell.

order, but no improvement in the resolution was found.

After the spectrum of the 8-m path of water vapor had been obtained, the entire region was scanned by using the atmospheric path in the spectrometer alone. In this background spectrum, a line was found at 791 cm<sup>-1</sup>, which agrees in position with the Q-branch of a band of carbon dioxide. It is very weak in the laboratory spectrum and is probably caused by the carbon dioxide in the air path of the spectrometer.

To obtain the spectrum, small portions, corresponding to the rotation through 1 deg of the grating circle, were scanned. The rocksalt foreprism was adjusted for maximum energy in the middle of each degree, which was then run at least twice in each direction. Fiduciary marks were made on the chart paper by a mechanical trigger device operated by an observer viewing coincidences of the grating circle markings with a graduated scale in the eyepiece of a microscope. The "central image" corresponding to the position of the grating for zero-order diffraction was measured during each day's work, in both directions. It was found to vary a few seconds of arc from day to day.

Frequencies were calculated by using the formula

#### $\nu = nK/\sin \theta$ ,

where *n* is the spectral order,  $\nu$  the frequency, *K* is a constant, and  $\theta$  is the angular displacement from the central image. *K* was determined from measurements of the 0.546073-, 0.576960-, and 0.579066- $\mu$ mercury lines in the orders from 10 to 16 using a low-pressure (H-2) mercury arc as a source and a 931A photomultiplier tube as detector. These measurements gave an average value for *K* of 722.506  $\pm$  0.025 cm<sup>-1</sup>.

The four or more frequency measurements obtained for each line, including scannings in both directions, were averaged. These frequencies were then corrected to vacuum by assuming the laboratory air had an average refractive index of 1.0002565, corresponding to a barometric pressure of about 740 mm Hg and a temperature of  $25^{\circ}$  C.

#### 3. Results

A map of the water-vapor spectrum from 4.5 to 13  $\mu$  is shown in figures 2 to 6, with the exception of

a short interval from 5.7 to 6.6  $\mu$ , where water-vapor absorption is very intense. To reach the longest wavelengths, the grating had to be used at extreme angles to the incident radiation, and very little energy was obtained even with the widest slits permissible. This accounts for the small deflection shown in figure 2 beyond 12  $\mu$ . As short portions of the spectrum were run individually, the total energy varied slightly, giving the segments shown in the figures. The approximate amounts of water vapor in the path for each region are given in the captions of the figures, as well as the temperature of the cell or spectrometer. It should be noted that when a temperature of 110° C is indicated, this refers to the temperature of the absorption cell. There was always some water vapor in the spectrograph at about 25° C, also contributing to the absorption. This contribution is quite negligible when the cell contains steam at atmospheric pressure. The regions near the center of the 6.3- $\mu$  band were run with only the air path in the spectrometer. Although spectra were recorded on a cold, dry day in February, many lines were still more intense than was desired.

The effective slit widths used varied from 0.22  $cm^{-1}$  at 13  $\mu$  to 0.83  $cm^{-1}$  at 5  $\mu$ . These slit widths were approximately the same as those used in the investigation of the solar spectrum. Under these conditions the lines in the laboratory spectrum were wider than corresponding lines in the solar spectrum. Consequently, some lines clearly separated in the solar spectrum are blended in the laboratory spec-The increased line width observed in the trum. spectrum of steam is caused both by the higher temperature and the much higher concentration of water vapor, giving a marked self pressure broadening effect. Approximately the same deflections were obtained in windows between absorption lines whether the cell was full of steam or air, indicating that there was little continuous absorption by water vapor in these windows. Thus the wings of watervapor lines lying outside the region investigated played an insignificant part in the absorption observed.

The lines have been numbered in order of increasing frequency in the figures as an aid to their identification in table 1. The line number is given in the first column of this table and the observed frequency, corrected to vacuum, in the second.

It is believed that the absolute values for the frequencies of well-defined lines are correct to better than  $0.2 \text{ cm}^{-1}$  at short wavelengths and to  $0.1 \text{ cm}^{-1}$  elsewhere. Most line frequencies have been given to  $0.01 \text{ cm}^{-1}$ , but some of the very weak lines that are difficult to measure accurately are listed to the nearest  $0.1 \text{ cm}^{-1}$ . A number of absorptions can be seen to consist of more than one line, and here the frequency of the position of the maximum absorption is listed to the nearest  $0.1 \text{ cm}^{-1}$ . Scattered radiation did not amount to more than 10 percent of the total deflection anywhere in the region measured and would not give rise to any spurious lines.

When the solar spectrum is compared with the laboratory spectrum, it is found that the relative intensities of some of the lines are very different.

This is to be expected because the amount of water vapor in the atmospheric path traversed by the solar radiation varies between 1,000 to 10,000 atm-cm and is at an average temperature of about 14° C, whereas the laboratory path was only 800 atm-cm but at a temperature of 110° C. The higher temperature of the laboratory sample gave intensities comparable with those in the solar spectrum to lines that originate from higher energy levels. However, lines arising from low-energy levels are little affected by the change in temperature from 14° to 110° C, and these lines were very weak or completely absent from the laboratory spectrum, although they may be quite prominent in the solar spectrum. This is illustrated by figures 7 and 8, which show small portions of the solar and laboratory spectra for comparison, together with the calculated spectra for the corresponding temperatures. The calculated spectra are schematically shown by drawing each line as a triangle, the width corresponding to geometric slit width used in the observed spectra, and the altitude proportional to the logarithm of the calculated intensity. The solar spectrum in figure 8 was observed during a cold winter day when the water-vapor content of the atmosphere was unusually low; so the calculated intensities for 14° C of table 1 were arbitrarily divided by 5 and plotted on the logarithmic scale indicated in the figure. Figure 7 also illustrates the greater line width of the laboratory steam spectrum than that of the solar spectrum.

In a number of cases where laboratory lines are very weak, or where several lines are blended, it is believed that more accurate frequencies can be obtained from the solar spectrum. Lines that appear in both spectra, but for which the frequencies from the solar spectrum have been used because they seem more reliable, are indicated by a dagger (†) in table 1. A few lines which appear only in the solar spectrum, but which can be fairly positively identified with water vapor because their intensities change with the amount of water vapor in the atmosphere or because they are theoretically predicted, have also been included in the table. Such lines have a dash in place of a line number.

The observed intensities of the lines are given in the next four columns. The first column gives the intensities of lines observed in the solar spectrum. A question mark in this column indicates that the intensity is uncertain because of absorption by other gases in the atmosphere. The values given are for "average" conditions of humidity and altitude of the sun. The intensity scale is very approximate, but runs from 0 to 100, corresponding roughly to percentage absorption at the maximum. The next three columns give intensities as observed under various laboratory conditions; first, when water in the spectrometer alone was sufficient; next, when steam was diluted to a small fraction of its maximum value; and finally, when the full path of 800 atm-cm was used. Figures 2 to 6 may be consulted for details. The steam was diluted over a wide range of concentrations so that the intensities for the diluted water vapor should be compared only for lines in the same neighborhood. In general, the absolute intensities of lines in the solar spectrum and the laboratory spectrum of the long path length of steam are roughly the same, although the intensities of the atmospheric lines vary considerably, depending on the altitude of the sun, the humidity, and the temperature of the atmosphere.

The sixth column of table 1 gives the identification of the line; or, if several transitions overlap to give an unresolved blend, the strongest components are listed. The identification consists of a letter symbol for the vibrational transition involved, the rotational quantum numbers of the upper state,  $J'_{\tau}$ , and those of the lower state,  $J''_{\tau}$ . The letter symbol *a* refers to pure-rotation transitions within the ground vibrational level. All identified lines with frequencies below  $890 \text{ cm}^{-1}$  are of this type, and they predominate up to  $1,000 \text{ cm}^{-1}$ . Near  $1,000 \text{ cm}^{-1}$  is a region of minimum absorption by water vapor. At lower frequencies the pure-rotation lines become stronger, and at higher frequencies those of the  $\nu_2$  band rapidly increase in intensity. The symbol b identifies transitions in the  $\nu_2$  fundamental. The remaining strong lines in the laboratory spectrum belong to this band. In addition, a few weak lines have been assigned to bands in which  $\nu_2$  is the lower state: c refers to the transition  $2\nu_2 - \nu_2$ , d to  $\nu_1 - \nu_2$ , and e to  $v_3 - v_2$ .

The seventh column gives the calculated frequencies derived from the term values of the corresponding rotational states. These are listed for the three lowest levels in table 2, which has been compiled partly from the present research, and partly from an extensive reinvestigation of the complete watervapor spectrum. The term values of the upper states of bands d and e have previously been given by Benedict and Plyler [5]. A calculated frequency in table 1 is the difference between two energy levels,  $J'_{\tau}$  and  $J''_{\tau}$ , involved in the transition. The calculated frequencies of a few lines have been placed in parentheses because in each case one of the levels concerned has been located by means of the observed line. These assignments depend on approximate predictions of the levels from extrapolations of regularities among the reliably known levels and on the agreement between calculated and observed intensities. For the remaining lines, however, other observed transitions connect the levels in question, and the agreement between observed and calculated frequencies is a true measure of the consistency of the measurements and the interpretation of the  $H_2O$  spectrum.

The final columns of table 1 list relative values of the calculated intensities, for temperatures of 14° and 110° C. These temperatures approximately correspond to the average temperature in the solar and laboratory spectra, respectively. The tabulated numbers are, for 14° C,  $v(LS)ge^{-E''/kT}$ .10<sup>6</sup>, where (LS) is the line strength, as defined and tabulated by Cross, Hainer, and King [6], g is the statistical weight, and E'' is the energy of the lower state (from table 2). v is a numerical factor for the intensity of the vibrational transitions relative to the pure-rotation band, which will be discussed presently. The factor 10<sup>6</sup> was arbitrarily chosen so that medium-strength lines in the solar spectrum have intensities of the order of unity. For  $110^{\circ}$  C the factor was reduced to  $10^{5}$  to compensate for the increased Boltzmann factor.

For the pure-rotation band, it is possible to calculate intensities on an absolute basis, since all quantities in the theoretical formula

$$\int k_{\nu}d\nu = \frac{8\pi^3 N}{3hcQ} p^2 \nu.(LS)ge^{-E^{\nu}/kT}$$

are known.  $\int k_{\nu} d\nu$  is the integrated intensity of a line, N, h, c, and k are universal constants, Q is the state-sum,  $\Sigma g e^{-E'/kT}$ , and p is the static dipole moment, 1.86 Debye units. In order to convert our tabulated values for the pure-rotation band to the integrated intensity in units of cm<sup>-2</sup>atm<sup>-1</sup>, one must multiply by  $2.19 \times 10^{-7} \times \nu (\text{cm}^{-1})$ , at 287° K; at 383° K the factor is  $1.07 \times 10^{-6} \times \nu (\text{cm}^{-1})$ .

Not all of the transitions found in the course of this study are listed in the tables of reference [6]. The extension to  $J \ge 12$  requires a moderate extrapolation. Transitions in the  $R_{5,5}$  branch do not appear in their tables; according to a communication from Professor Cross, the line strength in nearly all these lines is less than 0.0001. For such lines the intensities have been estimated, and are placed in parentheses in table 1; the relative intensities at the two temperatures are significant, but the values may be in error by an order of magnitude.

The observed intensity of relatively weak lines in the pure rotation band, in both the solar and laboratory spectra, appear to be in fair agreement with the calculated absolute intensity. For example, the calculated intensity of line number one  $(16_{-8} - 15_{-10})$  $\int I_{\nu} d\nu = 0.45 \times 2.19 \times 10^{-7} \times 760 = 7.3 \times 10^{-7}$ is  $10^{-5}$  cm<sup>-2</sup> atm<sup>-1</sup>. If the atmospheric path is  $5 \times 10^{3}$ atm-cm, the equivalent width should be  $0.36 \text{ cm}^{-1}$ . which is of the same order as given by the observed maximum absorption of 45 percent, assuming that the line has a half-width  $0.10 \text{ cm}^{-1}$  and is observed with a slit width of  $0.22 \text{ cm}^{-1}$ . For the steam path of 800 atm-cm the calculated equivalent width of the same line becomes  $1.6 \times 1.07 \times 10^{-6} \times 760 \times 800 =$ 1.04 cm<sup>-1</sup>, again in order of magnitude agreement with the observed maximum absorption of 68 percent for a line of half-width  $0.4 \text{ cm}^{-1}$ .

For lines in  $\nu_2$ , and in the upper-state bands, the absolute value of the theoretical intensity cannot be calculated, as it depends on the empirical factor v. The following values of v were chosen by which to multiply the pure-rotation intensities, in order to obtain fair over-all agreement with the observed intensities:

Band	$b(v_2)$	$c(2  u_2 -  u_2)$	$d(\mathbf{v}_1 - \mathbf{v}_2)$	$e(\nu_3-\nu_2)$
Factor	0.005	0.01	0. 0002	0.001

These result in lines of the observed order of magnitude and are not unreasonable. The  $\nu_2$  intensities, calculated in this way, appear to be somewhat

weaker than the observed intensities, in the region 900 to 1,050 cm<sup>-1</sup>, and somewhat stronger than the observed intensities in the region 2,000 to 2,200 cm<sup>-1</sup>, by a factor of 2 or 3 in each case. These effects, which may be in part due to the greater effective slit width in the short-wave region, are similar to those encountered in other bands of H<sub>2</sub>O, and even more strikingly in H<sub>2</sub>S and other asymmetric-top molecules. An explanation of these effects has recently been proposed by Nielsen [7]. The anomaly in  $\nu_2$  of H<sub>2</sub>O does not appear outstandingly large in the portions of the *P* and *R* branches near the origin, but is confined to the transitions of high *J* and those in which *K* changes by more than 1.

## 4. Discussion

The agreement between the observed and calculated spectra as given in table 1, and illustrated in figures 7 and 8, is quite complete, and argues well both for the accuracy of the measurements and the correctness of the interpretation. The lines not accounted for constitute less than 1 percent of the total absorption intensity. A few of the assignments to transitions of highest J, in both the purerotation spectrum and in  $\nu_2$ , are rather uncertain. It would be desirable to obtain further accurate measurements of steam in the region 550 to 770  $\rm cm^{-1}$ and of superheated steam or  $H_2$ - $O_2$  emission spectra throughout the rest of the vibration-rotation region, for confirmation. However, all the lower energy levels, up to J=10, are well accounted for, in that all predicted lines appear with approximately the correct intensity and frequency, and that no strong lines remain unidentified.

An additional demonstration that the spectral analysis is quite complete and essentially correct is afforded by the arrangement of the lines into series. Table 3 presents the lines of the four prominent Pbranches of the  $\nu_2$  vibration-rotation band, grouped in such a way as to display the regular decrease in frequency with increasing J. A similar array may be made for the lines in the R branches, as well as for the Q branch lines, most of which, however, lie in the central region not measured in this study. Frequencies in parentheses are those of weak lines overlapped by a stronger component. It is clear, by a comparison of tables 3 and 1, that the identification of many lines of higher J cannot be expected without an increase of the path length, or, preferably, by increasing the temperature. As mentioned before, there is a distinct difference in temperature coefficient between the weaker lines of the  $P_{1,\overline{1}}$ ,  $P_{\overline{1},\overline{1}}$ , and  $P_{\overline{1},1}$  branches, which have relatively high transition probabilities but originate from high-energy levels, and those of the  $P_{\overline{3},1}$  branch (as well as the few identified lines of the  $P_{\overline{3},3}$  branch, not included in table 3), which have lower line strengths but originate from lower energy levels. This difference manifests itself in the changed relative intensities in the laboratory and solar paths.

Table 4 is a similar presentation of the lines of the low-probability branches of the pure-rotation band.

For completeness, there are included lower lines of the series, as observed by other investigators [12, 13] in the farther infrared, as well as frequencies, in parentheses, of lines that have not yet been measured with high accuracy. A comparison with table 1 again shows that the suggested array accounts for most of the observed lines from 760 to  $1,000 \text{ cm}^{-1}$ , with the lines that are most enhanced in the atmosphere belonging, for the most part, to the more highly forbidden  $R_{5,\overline{5}}$  and  $R_{5,\overline{5}}$  series. The energy levels of  $\nu_2$ , which are now known with

completeness and good precision up to J=9, permit an improved determination of the rotational con-stants for that vibrational level. The method is described elsewhere. Figure 9 presents the effective rotational constant  $A'_J$ ,  $B''_J$ , and  $C'_J$  for each J, derived from the levels of table 2, plotted against J(J+1). The points are seen to fall nicely on smooth curves. The intercepts give the rotational constants A, B, and C, and the limiting slopes give the centrifugal stretching constants  $D_A$ ,  $D_B$ , and  $D_C$ . These are listed in table 5 and compared with the corresponding constants of the ground state. There is an appreciable curvature in the plot of  $A'_J$ , showing that centrifugal stretching terms of higher order are important at relatively low J.

The few weak lines that are not assigned may be transitions of higher J in  $\nu_2$ , upper-state transitions, or lines in the isotopic molecules  $H_2O^{18}$  or HDO. The calculated vibrational shift for  $\nu_2$  of  $H_2O^{18}$  is -6.5 cm<sup>-1</sup>; this shift would be increased for lines in the R branch and diminished in the P branch. Tentative assignments of some of the observed lines can be made on this basis; for example, line 285  $(1414.99 \text{ cm}^{-1})$ , which lies  $4.3 \text{ cm}^{-1}$  below the strong line  $3_3-4_3$ , and line 353 (1764.22 cm<sup>-1</sup>), which lies 8.4 cm<sup>-1</sup> below the strong line  $3_3-2_1$ , have the proper intensity ratio (1:500) and positions to be their analogues in  $H_2O^{18}$ . Other such possible lines are suggested in the table. It should be noted that, on the basis of the expected relative intensities, it would be very unlikely that many absorption lines due to HDO would appear. Although absorp-tion due to the  $\nu_2$  band of this isotope has been postulated by Adel [8] in the 7.2  $\mu$  region of the solar spectrum, and the  $\nu_1$  fundamental has been definitely observed, resolved, and analyzed in the solar spectrum at 3.7  $\mu$  [9, 10], the intensity of the lines of the  $\nu_2$  fundamental of HDO relative to those of  $H_2O$  on our spectra would be very small. The frequencies of the  $\nu_2$  fundamental of HDO published by Barker and Sleator [11] do not show any definite coincidences with our data. There should, however, be a number of fairly strong lines at lower frequencies than their work reached, which would lie in the region of weaker H<sub>2</sub>O absorption and might thus appear as weak lines in our spectra. For example, on the basis of the known structure of HDO, we calculate that the line  $3_3-4_3$  should fall near 1,250 cm<sup>-1</sup>. Line 183, otherwise unaccounted for, has about the proper intensity in both laboratory and solar spectra for this transition.

It may also be of interest to point out that the completeness of the assignment, and the fact that the intensity differences in the spectra between atmospheric water vapor and nearly saturated steam can be explained in detail on the basis of temperature, is strong evidence, although of a negative sort, against the existence of  $(H_2O)_2$  molecules in the saturated vapor. The association of  $H_2O$  in the liquid phase is abundantly proved by the marked lowering of frequency in the  $\nu_3$  fundamental and a somewhat smaller shift in  $\nu_2$ . The existence of dimers in the vapor has often been postulated to account for changes in density, dielectric polarization, etc., near saturation. If such dimers exist as molecules with quantized vibrationrotation states, the evidence of the present study would be that their concentration, relative to H<sub>2</sub>O monomer, must be less than 1 percent, since in the region of 1,400 cm<sup>-1</sup>, where  $(H_2O)_2$  should be most strongly absorbing, the unaccounted-for absorption is very small. We have mentioned previously that the steam spectrum shows a somewhat greater line width than the atmospheric spectrum, but there is no appreciable frequency shift, and the amount of self-broadening is by no means abnormally large for a polar molecule.

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FIGURE 2. Absorption spectrum from 11.9 to 13.1µ of an eight meter path of water vapor at 110°C and one atmosphere pressure.



FIGURE 3. Absorption spectrum from 10.2 to  $11.9\mu$  of an eight meter path of water vapor at  $110^{\circ}C$  and one atmosphere pressure.



FIGURE 4. Absorption spectrum from 10.2 to  $11.9\mu$  of an eight meter path of water vapor at  $110^{\circ}C$  and one atmosphere pressure and with small sections also at lower concentration.



FIGURE 5. Absorption spectrum from 5.5 to  $5.7 \mu$  and from 6.6 to  $8.0 \mu$  of water vapor.

Solid line refers to water vapor in the air of the spectrometer at approximately  $25^{\circ}$ C, dotted line to an eight meter path at  $110^{\circ}$ C and various concentrations of water vapor in air.



FIGURE 7. Solar, laboratory, and calculated spectra of the region from 850 to 880 cm<sup>-1</sup>.

(A) Solar spectrum taken in August, 1950; (B) calculated spectrum at  $14^{\circ}$ C; (C) laboratory spectrum of an eight meter path of steam at one atmosphere and  $110^{\circ}$ C; (D) calculated spectrum at  $110^{\circ}$ C.



FIGURE 8. Solar, laboratory and calculated spectra of the region from 1860 to 1925 cm<sup>-1</sup>.

(A) Solar spectrum taken in February 1951, with an air temperature of  $-15^{\circ}$ C; (B) calculated spectrum at 14°C; (C) laboratory spectrum of a four meter path of steam at 10°C and nearly one atmosphere; (D) calculated spectrum at 110°C.



FIGURE 9. Curves from which may be derived the inertial and centrifugal stretching constants associated with rotation about the three axes of inertia: (A) The least axis; (B) the intermediate axis; (C) the greater axis.

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	OB	SERVED	DATA					*	CALCULA	TED DAT	Δ
mber	Wave		Intensit	у			Identific	ation	Wave	Intens	ity
Nu L	cm <sup>-</sup> ' (vac)	Solar	Lo	iborator	110° C	-	J' <sub>Z</sub> —	J <del>″</del>	cm-' (vac)	14° C	110° C
			only	<8 Atmo-M	8 Atmo-I	•				I	
1	767.37	45	0	-	68	8.	16_8	15-10	(67.4)	0.44	1.6
2	769.1	50?	0	-	13	8.	75	<sup>6</sup> -5	69.15	(0.7?)	(0.12?
3	770.17	60?	0		75	8	12 <sub>2</sub>	11_4	70.17	3.6	3.5
4	775.63	30?	0	- ,	73	8	12 -2	11-8	75.63	6.4	4.6
5	777.07	60?	0	68-	94	a	105	<sup>9</sup> 1	77.09	11	7.5
6	779.36	46	0	-	72	a	106	9 <sub>0</sub>	79.41	3.7	2.5
7	784.54	60?	0	8	73	<b>{</b> a a	$11_{-\bar{3}}_{16_{-7}}$	10_9 15_11	84.62 (84.54)	18 1.2	7.2 4.3
88	791.6	-	5	-	5						
9	794.01	53?	0	-	89	a	<sup>13</sup> -1	12_7	94.01	6.4	8.0
10	796.01	85	0	-	88	\$ 8.	101	9_7	96.08	20	8.1
11	797.65	50	0	2005	5 <b>6</b>	8	11_4	100	97.66	2.3	1.9
12	798 <b>.7</b> 5	99	0	73	95	8.	15_11	14-13	(98.75)	6.8	9.0
13	799.11	75	0	48	70	${a \\ a}$	10-1 15-10	<sup>9</sup> -9 14-14	99.16 (99.11)	25 2.2	7.7 3.0
14	803.09+	85	0	-	8 <b>8</b>	8.	115	10_1	03.10	6.7	5.5
15	803.60+	93	0	-	88	a	131	12_5	03.61	3.1	5.2
16	806.06	30	0	-	54	a	16 <sub>-10</sub>	<sup>15</sup> -12	(06.06)	0.3	0.75
17	806.75	10	0	-	36						
18 {	808.14 <sup>†</sup> 808.33 <sup>†</sup>	<sup>80</sup> 90 }	0	-	82	{a a	16_9 85	$15 \\ 7 \\ -3 \\ -3 \\ -3 \\ -3 \\ -3 \\ -3 \\ -3 $	(08.14) 08.29	0.9	2.2
19	813.97	92	0	-	62	a	99	<sup>8</sup> 3	13.97	2.5	2.2
20	814.61	95	0	-	77	a	123	11-1	14.60	3.4	4.1
See	footnotes	at a	nd of	tohl							

TABLE 1. Absorption Spectrum of  $H_2O$  between 4.5 $\mu$  and 13 $\mu$ .

TABLE 1-Continued.

	OBSERVED DATA								CALCULATED DATA		
ine	Wave		Intens	ity			Identifico	ation	Wave	Intens	sity
Ž	cm <sup>-</sup> ' (vac)	Solar	Air only	Laborator 110° C (8 Atmo-	110° C M 8 Atmo-		J' <sub>7</sub> —	J <del>"</del>	cm-' (vac)	14° C	110° C
21	816.55	10	0	-	20		*		-	1	
22	825.24	85	0	-	5	a	84	<sup>7</sup> -6	25.16	1.0?	0.4?
23	827.21	50	0	-	18	a	132	<sup>12</sup> -2	27.31	0.42	0.73
24	827.80	95	0	-	42	{a a	12 <sub>4</sub> 9 <sub>4</sub>	<sup>11</sup> -2 <sup>8</sup> -4	27.86 27.67	0.45 1.4	0.57 0.47
25	835.64	32	0	-	36	{a a	141 17_9	13 16-3 16-11	35.64 (35.64)	0.30 0.10	0.78 0.55
26	840.01	90	0	-	66	{a a	10 107 108	9 <sub>3</sub> 92	40.01 40.15	1.6 0.53	1.7 0.56
27	840.36	45	0	-	46	a	140	13_6	(40.36)	0.24	0.50
28	841.16	5	0	-	13	a	150	14-4	(41.16)	0.016	0.065
-	841.97	26	0	-	0	a	77	6_3	41.93	(0.2)	(0.04)
29	845.83	3	0	-	9						
30	849.69	78	0		56	a	110	10_8	49.79	2.9	1.6
31	852.62	99	0	50	92	{a a	10 133 3	9-5 12-3	52.51 52.60	4.3 0.88	2.0 1.5
32	852.91	86	0	59	96	a	16-11	<sup>15</sup> -15	(52.92)	1.7	3.4
33	853.41	40	0	12	40	a	14_2	13 <u>-</u> 8	53.41	0.59	1.0
34	854.66	92	0	40	75	8	13 <u>-</u> 3	12_9	54.66	3•5	3.3
35	858.65	27	0	-	49	a	17 -11	<sup>16</sup> -13	(58.65)	0.15	0.64
36	859.78	9	0	-	29	a	17-10	<sup>16</sup> -14	(59.78)	0.05	0.21
37	861.0	-	Ó	-	4						
38	864.0	-	0	-	3						
39	865.02 <sup>T</sup>	32	0	-	37	a	116	102	65.01	0.46	0.49
40	865.51	58	0	-	54	а	<sup>11</sup> 7	101	65.49	1.4	1.5

TABLE 1-Co	ontinued.
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	OBS	SERVED	DATA						CALCULA	TED DAT	Δ
ine mber	Wave	×	Intens	ity			Identific	ation	Wave	Intens	ity
- NZ	cm-' (vac)	Solar	Air only	Laboratory	IIO* C 8 Atmo-M		J'	J <u>″</u>	cm-' (vac)	14° C	110° C
41	871.32	78	0	-	<b>3</b> 6	8.	11_2	10_10	71.49	2.3	0.92
42	878.61+	68	0	-	<b>3</b> 6 *	{ a a	$12_{-4}$ $14_{2}$	11-10 13-4	78 <b>.7</b> 0 78.9	1.3 0.08	0.65 0.20
43	881.15	23	0	-	2	8.	86	7_4	81.20	(0.3)	(0.08)
44	883.18	24	0	-	32	a	15_1	14_7	(83.18)	0.13	0.18
45	883.89	59	0	-	25	a	112	10_6	84.01	0.95	0.57
46	887.33	72	0	-	22	a	93	8_7	87.44	(1.1)	(0.27)
47	888.71	<b>4</b> 6	0	-	41	8	125	111	88.71	0.63	0.90
48	890.14	26	0	-	16	a	126	110	90.16	0.21	0.30
49	891.33	41	0	-	17	8.	109	95	91.33	0.28	0.26
50	892.1	6	0	-	4						
51	896.57	10	0	-	4	Ъ	<sup>12</sup> -12	13_8	96.53	0.006	0.009
52	897.77	22	0	-	14	Ъ	12-11	13_9	97.77	0.017	0.027
-	905.50 <sup>†</sup>	10	0	-	0	Ъ	10_7	11_1	05.34	0.012	0.014
53	906.32	56	0	-	6 <b>9</b>	8.	17-13	<sup>16</sup> -15	(06.32)	0.3	0.8
54	906.8	13	0	-	30	e.	151	14_5	(06.6)	0.06	0.09
55	908.09†	15	0	-	12						
56	909.02	83	0	-	6 <b>7</b>	a	12-1	11_9	09.12	3.0	2.0
5 <b>7</b>	910.17	20	0	-	21	a	134	120	10.4	0.08	0.17
58	910.77	12	0	-	27	a	<sup>18</sup> -11	<sup>17</sup> -15	(10.8)	0.02	0.12
59	914.06	34	0	-	33	8	135	12-1	14.06	0.24	0.50
60	918.52	37	0	-	23	a	11,9	103	18.52	0.25	0.31

TABLE 1-Continued.

	OBS	•					CALCULATED DATA				
mber	Wave		Intens	ity			Identifica	ition	Wave	Intensity	
2 Z	cm-' (vac)	Solar	Air	Laboratory	110° C 8 Atmo-M		J <sub>7</sub> — .	J <del>″</del>	cm-' (vac)	14° C	110° C
61	921.48	35	0	-	31	a	15 2	14_0	(21.48)	0.29	0.70
62	022.10+	89	0		52	A	12.	11 -	22.41	1.5	1.2
(02		80	Õ		2	ũ	1	<b>**-</b> 7	05 10		(0, 2)
63	925.05	00	0	-	15	a	95	° <b>-</b> 5	23.12	(1.0)	(0.3)
64	929.00	22	0	-	18	a	<sup>14</sup> 3	<sup>13</sup> -1	29.00	0.063	0.20
65	929.4	0	0	-	6						
66	932.32	1	0	-	5						
67	937.40	5	0	- 1	6	a	144	13_2	37.2	0.020	0.06
68	938.2	2	0	-	.4						
69	941.12	20	0	-	8	a	14-4	13_10	(41.1)	0.29	0.36
70	944•95	13	0	-	11	{b a	$\frac{9}{127}$	10 113	44•90 44•96	0.022	0.029 0.35
-	946.73 <sup>†</sup>	6	0		0	a	<sup>8</sup> 7	7_1	46.73	0.13	0.04
71	948.35	66	0	-	27	a	<sup>12</sup> -3	11_11	48.40	1.9	1.0
72	953•49	30	0	-	25	ъ	11_11	12 <sub>-7</sub>	53•49	0.069	0.078
73	<b>954.1</b> 7	7	0	-	6	b	8-1	95	53•99	0.014	0.013
-	955.33+	24	0	-	0	a	102	9_8	55.50	(0.3)	(0.09)
74	955.71	9	0	-	13	ъ	11_10	12_8	55.71	0.023	0.026
<b>7</b> 5	959.33	10	0	-	36	${a \\ a}$	18-13 96	17 8_17 8_2	(59•33) 59•79	0.05 (0.05)	0.21 (0.02)
76	{960.5 <sup>†</sup> 961.1 <sup>†</sup>	5 4	0	-	12	{a b	153 80	14-3 94	(60.5) 60.91	0.003	0.014
77	962.05	-	0	-	5	a	1111	105	62.1	0.025	0.04
78	963.0	- 1	0	-	3						
79	966.01	11	0	-	7	ъ	9_6	100	65.93	0.019	0,016
80	967.00	15	0	-	16	a	130	12_8	(67.00)	0.20	0,22

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	OBS	SERVED	DATA						CALCULA	TED DATA	7
mber	Wave		Intensit	у			Identificat	ion	Wave	Intens	ity
Z L	cm <sup>-</sup> ' (vac)	Solar	Air only	110° C (8 Atmo-M	110° C 8 Atmo-M		J <mark>7</mark> — J	C I	cm-' (vac)	14* C	110* C
81	970.66	10	0	-	8	{a a	13 <sub>6</sub> 13 <sub>7</sub>	12 <sub>2</sub> 12 <sub>1</sub>	70.52 70.75	0.016 0.047	0.04
-	971.43	18	0	-	0	а	105	9 <b>-</b> 3	71.53	0.20	0.10
82	971.75	12	0	-	20						
83	973-59	15	0	-	19	а	13_2	12-10	73.65	0.33	0.30
84	974.04	16	0	-	7	a	104	9_6	74.04	0.5	0.19
85	976.07	40	0	-	24	a	13_5	12_11	76.07	0.7	0.5
86	977.54	15	0	-	8	Ъ	8-7	9_1	77.28	0.084	0.045
87	984.2	5	0	-	5	a	114	10-4	84.15	0.14	0.10
88	990.3	6	0	-	3	a	129	115	(90.3)	0.007	0.015
89	994.4	10	0	-	4						
90	998.87	30?	0	-	10	b	<sup>8</sup> -3	93	98.70	0.080	0.026
91	1000.35	25?	0	-	6	а	123	<sup>11</sup> -5	00.47	0.20	0.19
92	1003.8	10?	0	-	3	а	<sup>9</sup> 7	<sup>8</sup> -3	03.85	(0.40)	(0.14)
93	1007.3	15?	0	-	4	р	11_5	12 -1	07.08	0.017	0.035
94	1010.12	15?	0	-	16	b	<sup>10</sup> -10	11_6	09.96	0.11	0.085
95	1010.86	15?	0	- ,	22	b	11_9	12 <u>-</u> 5	11.06	0.10	0.143
96	1011.64	10?	0	-	7					•	
97	1014.56	50?	0	-	38	ъ	109	11_7	14.38	0.318	0.265
98	1017.6	20?	0	-	15	ъ	8-5	91	17.79	0.202	0.128
99	1017.9	30?	0	-	17	a	14-1	13_9	(17.9)	0.16	0.22
100	1019.5	5?	0	-	2	b	10_4	110	19.52	0.013	0.018

-						-					
	OB	SERVED	DATA						CALCULA	TED DATA	
ine mber	Wave		Intensit	у			Identific	ation	Wave	Intensi	ty
Nn	cm-' (vac)	Solar	Air only	ilo* c {8 Atmo-M	IIO* C 8 Atmo-N	•	J <sub>7</sub> —	J <sub>C</sub>	cm-' (vac)	14* C	110° C
101	1022.0	5?	0	-	4	a	132	12_6	21.55	0.053	0.078
102	1028.47	30?	0	-	13	{ b a	$113^{7-2}$	<sup>8</sup> 4 10_7	28.47 28.45	0.043 0.5	0.024 0.25
103	1029.69	20?	0	-	17	${b \\ a}$	<sup>9</sup> -3 11 <sub>1</sub>	101 10_9	29.58 29.87	0.084 1.14	0.082 0.45
104	1030.58	10?	0	-	13	b	11_8	12-6	30.0	0.033	0.048
105	1032.76	20?	0	-	5						
106	1039.53	?	0	-	30	{ b b	8-2 11-7	92 12-3	<b>39.31</b> <b>39.</b> 35	0.043 0.075	0.034 0.128
107	1042.57	15?	0	-	17	a	14_3	13-11	(42.5)	0.22	0.28
108	1044.4	?	0	-	3						
<b>1</b> 09	1049.39	?	0	-	7	b	7-6	<sup>8</sup> 0	49.39	0.13	0.052
110	1050.22	?	0	-	6						
111	1051.28	?	0	-	11	b	7 <sub>-1</sub>	83	51.28	0.176	0.096
112	1055.56	20?	0	-	8	Ъ	6_1	75	55.21	0.164	0.075
113	1058.69	?	0	-	13	Ъ	12_7	13_5	58.57	0.028	0.065
114	1060.14	?	0	-	23	Ъ	10-8	11_4	60.06	0.146	0.146
115	1062.61	?	0	-	8	Ъ	7-4	<sup>8</sup> 2	62.50	0.154	0.075
116	1063.6	?	0	-	3						
117	1066.20	85?	0	-	70	{ъ ъ	9_9 60	<sup>10</sup> -5 74	66 <b>.14</b> 66 <b>.</b> 17	1.17 0.057	0.71 0.025
118	1072.69	?	0	-	15	Ъ	10_6	11-2	72.52	0.084	0.10
119	1074.46	70?	0	-	41	Ъ	<sup>9</sup> -8	10_6	74.31	0.415	0.25
120	1085.4	?	0	-	3	Ъ	<sup>5</sup> 1	6 <sub>5</sub>	85.37	0.096	0.035

TABLE 1-Continued.

	OBS	SERVED	DATA						CALCULA	TED DATA	
ine mber	Wave		Intensity	y			Identific	ation	Wave	Intensit	у
Nn Nn	cm-' (vac)	Solar	La Air only	boratory 110° C (8 Atmo-M	110° C 8 Atmo-	м	J' <sub>7</sub> —	J <sup>"</sup>	cm-' (vac)	14• C	110° C
121	1088.1	?	0	-	2						
122	1091.24	82	0	-	55	Ъ	10_7	11_5	91.21	0.575	0.54
123	1099.74	72	0	-	54	ზ	9_5	10_1	99.04	0.665	0.55
124	1101.47	76	0	-	34	Ъ	<sup>6</sup> -3	73	01.42	0.85	0.32
125	1105.3	5?	0	-	8						
126	1106.76	83	0	-	78	Ъ	<sup>9</sup> -7	10_3	06.73	1.60	1.15
127	1111.59	82	0	-	41	Ъ	6_5	71	11.50	1.21	0.41
128	1114.8	?	0	-	4	c	7_5	<sup>8</sup> -1	14.30	0.010	0.03
129	1117.71	23?	0	-	19	Ъ	11_6	12_4	17.60	0.063	0.098
130	1120.9	84	0	14	76	Ъ	<sup>8</sup> -4	9 <sub>0</sub>	20.76	0.45	0.29
131	1121.24	89	0	21)	10	Ъ	8-8	<sup>9</sup> -4	21.24	1.37	0.62
132	1135.80	98	0	44	86	ъ	8_7	<sup>9</sup> -5	35.80	1.78	0.80
133	1137.46	85	0	16	65	Ъ	7-3	<sup>8</sup> 1	37.19	1.63	0.78
134	1141.64	22 ?	0	-	23	Ъ	12_5	13_3	41.59	0.034	0.088
135	1149.48	77	0	22	78	ъ	8-6	<sup>9</sup> -2	49.39	1.82	0.98
136	1151.59	70	0	5	47	Ъ	<sup>6</sup> -2	7 <sub>2</sub>	51.55	0.66	0.25
137	1152.44	77	0	19	72	Ъ	9-6	10_4	52.42	0.91	0.63
138	1165.27	87	0	29	59	Ъ	5-4	<sup>6</sup> 2	65.10	0.86	0.22
139	1165.4 <sup>†</sup> 1165.9	83 73 }	0	15	35	{b b	<sup>5</sup> -1 4-1	63 55	65.40 65.84	1.56	0.48
140	1167.04	?	0	5	16	$\left\{ {}^{\circ}_{\circ} \right\}$	8-5 5_3	9-3 61	66.8 67.1	0.009	0.032

TABLE 1-Continued.

ſ	OBS	ERVED I	DATA						CALC	CULATED DAT	A
ine mber	Wave	Ir	ntensity	1			Identific	cation	Wave	Inten	sity
Nr L	cm-' (vac)	Solar	Air only (	HO° C I B Atmo-M B	IO* C Atmo-M		J <sub>ζ</sub> –	- J <sup>"</sup>	cm-' (vac)	14° C	110° C
141	1169.4	?	0	-	4						
142	1171.4	?	0	-	11	ъ	<sup>9</sup> -8	9 <sub>-2</sub>	71.47	0.043	0.021
143	1173.76	<b>)</b>	0	33	99	Ъ	<sup>10</sup> -5	<sup>11</sup> -3	73•74	1.11	1.09
144	1174-54	<b>\$</b> <i>""</i>	0	68		ъ	7-7	<sup>8</sup> -3	74.54	12.6	4.4
145	1178.55	15?	0	-	15						
146	1180.75	55?	0	-	9	b	40	54	80.94	0.29	0.06
147	1182.2	?	0	-	3						
148	1184.14	?	0	6	44	b	<sup>12</sup> -3	<sup>13</sup> -1	83 <b>•9</b> 5	0.068	0.206
149	1187.00	99	0	70	98	{b b	11 7_5	12 8-2 -1	86.65 87.11	0.105 11.4	0.189 4.7
150	1191.16	35 <b>?</b>	0	-	9						
151	1193.67	28?	0	-	9						
152	1195.41	?	0	6	27	ъ	<sup>12</sup> -1	131	95:28	0.046	0.196
153	1198.22	98	0	48	95	ъ	7_6	8-4	98.18	5.2	1.78
154	1200.80	?	0	2	17						
155	1201.55	?	0	4	30	{b b	12 <sub>1</sub> 12 <sub>2</sub>	13 <sub>3</sub> 132	01.58 01.8	0.021	0.099
156	{1206.07 1206.37	t 52? t 30?	} 。	-	13						į
157	1207.35	?	0	-	13	c	5-5	<sup>6</sup> -1	07.5	0.048	0.082
158	1209.79	?	.0	3	46	{° b	6 12 <del>5</del> 3	7-3 135-3	09.6 09.7	0.036 0.007	0.072 0.047
159	1211.29	99	0	23)	00	ъ	4-3	53	11.29	3.4	0.73
160	1212.28	99	0	60	77	b	8-5	9-3	12.23	12.8	6.4

TABLE 1- Continued.

	OBS	SERVED	DATA					CALCULA	TED DATA	
ine mber	Wave		Intensit	y		Identific	ation	Wave	Intensit	у
S L	cm-' (vac)	Solar	Air only	boratory 110° 0 110° C (8 Atmo-M 8 Atmo-	M	J' <sub>Z</sub> —	J	cm-' (vac)	14* C	110° C
161	1213.0	?	0	19 99	b	11 _2	120	13.00	0.10	0.22
162	1214.91	?	0	3 25	b	<sup>8</sup> -7	8-1	15.06	0.36	0.15
163	1218.63	98	0	40	Ъ	6_4	70	18.63	6.70	2.07
164	1219.1	?	0	9	Ъ	110	122	19.2	0.060	0.16
165	1220.43	?	0	18 70	Ъ	111	121	20.60	0.18	0.47
166	1225.08	99	0	38	<pre>{b b b b</pre>	11 6-6 9-4	$^{12}_{7-2}_{10-2}$	24.76 25.08 25.08	0.10 11.0 2.05	0.34 3.08 1.56
167	1225.5	99	0	25 99	b	<sup>10</sup> -3	111	25.46	1.97	2.36
168	1226.1	78?	0	9)	b	11_1	12_1	26.06	0.355	0.74
169	1228.53	?	0	- 13						
170	1229.43	?	0	- 14	b	9_9	9-3	29.35	0.80	0.40
171	1232.1	?	0	- 3					3	
172	1233.31	?	0	- 47	Ъ	115	125	32.88	0.041	0.18
173	1235.23	?	0	- 19						
174	1237.20	?	0	- 19	{b b	<sup>11</sup> _6 <sup>-9</sup>	<sup>11</sup> -3 6_6	36.7 37.41	0.056 0.021	0.061 0.008
175	1239.25	?	0	25 98	b	10_1	111	39.19	1.50	2.17
176	1240.6	?	0	- 35						
177	1242.90	95 <b>?</b>	0	17 98	{b b	10 101 2	11 113 112	42.75 43.1	1.17 0.39	2.11 0.70
178	1244.18	00	0	40	Ъ	5 <u>-</u> 3	6 <sub>1</sub>	44.04	24.1	6.27
179	1244.77	77	0	11 599	þ	100	110	44.66	0.52	0.99
180	1246.63	?	-	0 7	ъ	7_1	75	46.48	0.14	0.06

	OBSERVED DATA								CALCULA	TED DATA	
Line	Wave		Intensit	у			Identific	ation	Wave	Intensi	ly
Ž	cm-' (vac)	Solar	Air only	110° C (8 Atmo-M 8	HO* C Atmo-M		J <sub>7</sub> —	J <del>Ľ</del>	cm-' (vac)	14* C	110° C
181	1248.52	?	-	11	86	b	103	115	48.49	0.46	1.06
182	1251.43	?	-	-	13	Ъ	3_2	44	51.46	0.74	0.14
183	<b>12</b> 52•44	?	-	-	17						
184	1253.67	?	-	0	5	b	7_6	70	53.51	0.43	0.14
185	1255.95	?	-	3	40	b	<sup>16</sup> -15	17 <sub>-17</sub>	(56.0)	0.095	0.40
186	1257.07	?	-	6	67	{b b	10 <sub>5</sub> 11 <sup>5</sup> -3	11 127 -3	57.07 57.10	0 <b>.1</b> 82 0.645	0.56 1.10
187	1258.63	95	-	21	91	b	9_2	100	58.63	3.1	2.7
188	1260.38	99	-	67	99	Ъ	6 <b>-</b> 5	7 <b>_</b> 3	60.32	53•7	14.5
189	1264.04	?	-	22	92	ъ	90	102	63.93	2.0	2.1
190	1265.42	?	-	12)		b	4-2	52	65.30	5.5	1.15
191	1266.11	?	-	41		b	91	101	66.31	6.2	6.5
<b>19</b> 2	1266.63	?	-	39		b	93	<sup>10</sup> 3	66.55	4.2	4.5
193	1268.40	99	-	49	00	b b b	$10^{-2}$ $10^{7}$ $7^{-4}$	11 - 2 11 - 2 89 - 2	67.83 67.93 68.28	0.90 0.06 19.2	1.02 0.26 7.1
194	1269.97	99	-	60	• 79	ъ	8 -3	9-1	69.72	31.1	17.0
195	1271.80	99	-	61		Ъ	5_5	<sup>6</sup> -1	71.73	73	16.3
196	1272.37	?	-	31		Ъ	95	10 <sub>5</sub>	72.22	1.80	2.90
197	1273.1	80?	-	6)		{b b	6 9 <b>-</b> 3 9 <b>-</b> 3	6 93 3	72.64 73.24	0.78 0.14	0.23
198	1276.63	?	-	6	65	Ъ	15-15	<sup>16</sup> -15	76.67	0.50	1.35
199	1280.09	99	-	48	99	b	9-1	10_1	80.09	10.4	8.88
200	1281.22	85?	-	20	89	ъ	97	107	80.99	1.06	2.30

	OBS	DATA						CALCULA	TED DATA		
ine	Wave		Intensit	у			Identific	ation	Wave	Intensi	ty
ž	cm-' (vac)	Solar	Lo Air only	IDOratory	110° C Atmo-M		J'	JĽ	cm-' (vac)	14° C	110° C
201	1282.9	?	-	3	7						
202	1284.37	?	-	10	63	b	3_1	43	84.14	5.9	1.09
203	1284.89	?	-	0	40					*	
204	1287.38		-	61 )		b	8-1	91	87.29	33•9	21.7
205	1288.28	99	-	55	99	ъ	81	9 <sub>3</sub>	88.18	21.4	16.5
206	1288.92	)	-	33)		ъ	82	9 <sub>2</sub>	88.82	7.15	5.50
207	1290.59	?	-	47	98	ъ	83	95	90.52	14.6	18.2
208	1292.40	?	-	6	75	{c b	63 99	7 <sub>5</sub> 109	92.0 92.30	0.10 0.25	0.05
209	1296.67	?	-	55	99	{b b	8 <sub>5</sub> 80	97 90	96•59 96•53	6.25 12.1	7.81 7.6
210	1300.9	?	-	0	2	c	65	<sup>7</sup> 7	01.0	0.04	0.19
211	1302.7	?	-	0	2						
212	1304.46	?	-	0	22						
213	1305.60	?	-	23	89	Ъ	87	99	05.49	2.48	3.09
214	1306.31	?	-	5	54						
215	1308.25	?	-	55	96	b	7_2	80	08.18	42.0	17.4
216	1312.61	?	-	75	-	b	73	83	12.71	88	52
217	1313.64	?	-	64	-	{b b	4-4 70	50 82	13.35 13.68	39 36•7	7.4 17.6
218	1314.82	?	-	59	-	b	75	85	14.73	43.8	31.9
219	1316.20	?	-	32	-	b	<sup>11</sup> -5	<sup>12</sup> -5	16.38	1.67	2.39
220	1317.04	?	-	75	-	ъ	71	81	16.92	113	54

	OB	SERVED	DATA						CALCULA	TED DATA	
Line Number	Wave Number cm-' (vac)	Solar	Intensity Lab	oratory	110° C		Identifico $J_{\mathcal{T}}^{\prime}-J_{\mathcal{T}}^{\prime}$	ution J <del>",</del>	Wave Number cm-' (vac)	Intensi	ty 110* c
221	1317.6		-	37	-	Ъ	<sup>13</sup> -13	14-13	17.65	8.15	10.9
222	1318.97		-	77	-	{ b b	6-3 10-4	7-1 11-4	18.95 19.21	211 2.0	61 2.0
223	1320.09		-	58	-	ъ	<sup>5</sup> -4	<sup>6</sup> -2	20.13	61.4	13.3
224	1320.90		-	44	-	Ъ	77	87	20.93	19.2	18.0
225	1323.31		-	51	-	b	<sup>9</sup> -3	<sup>10</sup> -3	23.29	19.5	14.2
226	1324.30		, <b>-</b> ,	10	-	c	51	<sup>6</sup> 1	24.2	0.53	1.07
227	1325.63		-	8	-						
228	1326.14	ption	-	14	-	{ b b	$13 - 11 \\ 13 - 9 - 9$	14-11 14-9	26.08 26.4	1.93 0.50	3.48 1.20
229	1327.72	bs or	, <b>-</b> 1	2 <b>2</b>	-	Ъ	<sup>9</sup> -7	<sup>9</sup> -1	27.70	2.50	1.37
230	1329.90	e P	-	52	-	b	<sup>8</sup> -2	<sup>9</sup> -2	29.72	19.3	10.4
231	1332.70	ıp <b>le</b> t	-	8	-	Ъ	7_3	73	32.49	0.78	0.30
232	1335.55	Con	-	14	-						
233	1336.64		-	88	-	b	<sup>6</sup> 3	75	36.56	239	110
234	1338.56		55	92		{ъ ъ	$12_{61}$	$13 - 13 - 13 7_3$	37.83 38.50	28.6 308	27.5 117
235	1339.23	2	40)	07	-	Ъ	<sup>6</sup> 5	77	39.39	117.5	68.1
236	1339.55		39	93	-	Ъ	<sup>6</sup> 2	72	39.59	103	39
237	1340.35		67	95		{b b	7_1 6_1	$\frac{8}{71}$	40.09 40.37	169 364	69 118
238	1344.04		-	22	-	Ъ	12_9	13 <sub>-11</sub>	44.0	7.4	9.5
239	1344.60		-	7	-						
240	1345.59	•	-	13	-	ъ	12_7	13_9	45.7	1.55	2.60

TABLE 1- Continued.

	OB	SERVED	DATA		Τ			CALCUL	ATED DAT	7
ine mber	Wave		Intensit	y		Identific	ation	Wave	Intens	ity
Ž	cm-' (vac)	Solar	Air only	110° 0 110° C (8 Atmo-M 8 Atmo-		J'	J <del>″</del>	cm-' (vac)	14* C	110* C
241	1345.97		-	18 -	c	41	53	46.0	1.22	2.04
242	1347.01		-	6 -	c	42	5 <sub>2</sub>	47.1	0.41	0.68
243	1349.0		-	34 -						
244	1349.39		-	64 -	Ъ	<sup>3</sup> -3	4 <sub>1</sub>	49.43	113	18.4
245	1351.67		-	15 -						
246	1352.41		-	21 -	ъ	11-7	<sup>12</sup> -7	52.4	5.65	6.61
247	1354.87		-	80 -	b	6 <sub>0</sub>	° 7 <sub>0</sub>	54.80	164	53
248	1356.0		<u> </u>	13 -	{c b	5-1 8-6	6 8 <sub>0</sub> -1	55•4 56•54	0.86 2.6	1.47 1.08
249	1357.21	ptio	-	13 -						
250	1358.06	bsor	-	55 -	b	<sup>11</sup> -11	<sup>12</sup> -11	58.02	91	65
251	1361.09	ete A	70		{b b	11_9 55	12_9 65	61.02 61.09	19.5 570	18.1 211
252	1362.70	- Compl	70		{b b	5-2 52	6 <sub>0</sub> 6 <u>4</u>	62.70 62.96	215 250	50 76
253	1363.17		80		Ъ	5 <sub>3</sub>	<sup>6</sup> 3	63.16	745	226
254	1365.9		2							
255	1368.60		57		. b	50	62	68.74	347	89
256	1369 <b>.</b> 78		6	80 -	{b b	12 12 <sup>11</sup> -12	12 12-10	69.89 70.02	3.8 1.3	3.5 1.2
257	<b>1370.9</b> 5		4	74 -	c	3 <sub>3</sub>	43	70.9	3.2	4.6
258	1372.28		17		Ъ	<sup>9</sup> -5	<sup>10</sup> -5	72.17	52.5	31.7
259	1373.76		76		ъ	5	61	73.76	1070	275
260	1375.09		70		þ	4-3	5-1	75.13	670	107

TABLE 1-Continued.

	OBS	SERVED	DATA						CALC	JLATED DAT	4
mber	Wave		Intensity				Identificat	lion	Wave	Intens	ity
Ñ L	cm-' (vac)	Solar	Lab Air only (8	oratory 110° C Atmo-M	110°C 8 Atmo-M		J' J	τ	cm-' (vac)	14° C	110° C
261	1376.3	+	3	-	-						
262	1378.04		35	-	-	{b b	10 10_9	11 1111	77.98 78.07	263 19•5	1 <b>3</b> 9 13.3
263	1378.5		6	-	-	Ъ	7 <u>-</u> 5	71	78.66	17.2	5.5
264	1 <b>37</b> 9.63		25	-	-	{b b	<sup>8</sup> -4 10-7	9-4 11-9	79.5 79.66	45•5 58•5	20.9 39.8
2 <b>6</b> 5	1382.11		7	85	-	b	10_5	11_7	82.10	18.0	14.9
266	1383.40		5	44	-						
267	1384.26		4	35	-	c	7 <b>_</b> 5	8 <u>-</u> 5	83.8	0.56	1.26
268	1386.51	tion	56	-	-	Ъ	<sup>7</sup> -3	<sup>8</sup> -3	86.24	311	110
269	1387•55	d ros	82	-	-	ъ	<sup>4</sup> 3	<sup>5</sup> 5	87.50	2160	548
270	1390.0	Ab	4	-	-	Ъ	5_3	5 <sub>3</sub>	90.46	22.7	4.9
271	1390.95	1plete	12	-	-	{b b	11_10 11_11	11_8 11_9	90.78 91.07	4•5 13•4	3.1 9.3
272	1392.38	- Con	5	-	-						
273	1394:50		88	-	-	{b b b	9-7 41 6-2	<sup>10</sup> -7 53 7_2	94•46 94•55 94•58	159 2305 233	77 500 65
274	1395.81		75	-	-	ъ	42	5 <sub>2</sub>	96.00	770	167
275	1397•74		55	-	-	{b b	9-6 9-9	10_8 10_9	97•51 97•78	54 685	30 280
276	1399.16		82	-	-	(b) C	4-1 7-7	51 8-7	99.15 00.8	2170 1.9	414 3.6
277	1402.0		0	5	-	c	7_6	<sup>8</sup> -8	01.9	0.63	1.2
278	1403.54		8	-	-	ъ	9_4	10_6	03.38	17	10
279	1404.98		80	-	-	Ъ	5_1	<sup>6</sup> -1	05.10	1432	326
280	1407.09		1	22	-						

TABLE 1- Continued.

	OBSERVED DATA								CALCULA	TED DATA	
mber	Wave		Intensity	1			Identifico	ation.	Wave	Intensi	y
N <sup>L</sup>	cm-' (vac)	Solar	Air	HO* C	IIO* C 8 Atmo-k		J'7 —	J <del>″</del>	cm-' (vac)	14° C	110° C
281	1408.51	4	3	46	-	c	21	33	08.3	5.3	5.8
282	1409.94		24	85	-	Ъ	<sup>8</sup> -6	9_6	09.91	135	43
283	1411.52		27	-	-	{b b	10-3 10-9	11-5 10-7	11.33 11.43	5.2 42	5.1 23
284	1411.91		26	-	-	{b b	5-5 10-10	51 10_8	11.91 11.98	73•5 14	14.0 7.7
285	1414.99		5	38	-	Por	ssibly	$H_20^{18}$	ъ <sup>3</sup> 3	43	
286	1416.11		41	-	-	Ъ	<sup>8</sup> -5	<sup>9</sup> -7	16.09	382	149
287	1417•39	on no	69	-	-	{b b	8-8 8-7	9_8 9_9	17 <i>.3</i> 9 17.64	402 1225	181 392
288 <sup>h</sup>	1419.3	sorpti	95	-	-	{b b	40 33	50 43	19.04 19.55	935 5805	177 1090
289	1423.90	te Ab	73	-	-	{b b	3-2 7-5	40 8_5	23.90 24.25	625 890	94 269
290	1425.83	mple	0	19	-	c	<sup>6</sup> -3	7_5	26.3	0.5	0.9
291	1426.60	8 	0	45	-	c	3_1	4-1	26.6	4.1	4.5
292	1428.31		21	-	-	b	<sup>8</sup> -3	9 <b>_</b> 5	28.24	128	58
293	1429.97		80	-	-	b	30	42	30.01	1580	258
294	1432.06		22	-	-	Ъ	9_8	9_6	31.99	41	16
<b>29</b> 5	1433.31		40	-	-	{b b	9-9 4-4	9-7 42	33.21 33.49	123 36	38 5•7
296	1435.77		57	-	-	ъ	7_4	8_6	35.51	292	88
297	1436.72		93	-	-		7-7 31 6-4 7_6	8-7 41 7-4 8-8	36.54 36.67 36.89 37.18	2540 5150 585 835	650 840 143 214
298	1440.64		2	28	-						
299	1441.42		2	45	-	ъ	11_8	11_6	41.42	3.8	3.2
300	1442.80	+	0	9	-						

		OBS						CALCU	LATED DA	ТА			
	ine mber	Wave		Intensity	1			Identific	ation		Wave	Inte	nsity
	Nu L	cm-' (vac)	Solar	Air only (	HO* C 8 Atmo-M	110° C 8 Atmo-M		J' <sub>7</sub> —	· J <del>″</del>		cm-' (va	c) 14• c	110° C
:	301	1445.09	t	7	60	-	{Ъ Ъ	119 9_2	11_7 10_4		45.05 45.12	11.8	9.8 9.3.4
	302	1446.49		13	65	-	ъ	3_3	33		46.52	60	8.2
3	303	1447.91		85	-	-	b	5-3	6_3		47.88	2955	594
20	304	1450.53		1	19	-	Po	ssibly	H20 <sup>18</sup>	b	6_5	7_7	
3	305	1452.01		58	-	-	b	8-7	<sup>8</sup> -5	:	52.20	324	98
3	306	1453.4		1	-	-							
3	307	1454.59		44	-	-	b	<sup>8</sup> -8	<sup>8</sup> -€	Ę	64.64	109	33
3	808	1455.26	 ¤	75	-	-	b	6_6	7-6	5	5.30	1600	310
3	809	1456.49	ptio	95	-	-	Ъ	6_5	7_7	ŧ	6.45	4785	995
3	310	1457.09	Absor	89	-	- <	(b b	<sup>2</sup> 1 <sup>6</sup> -3	3 <sub>3</sub> 7_5	5	7.11	8400 1650	1210 402
3	11	1458.24	ete	89	-	-	Ъ	22	<sup>3</sup> 2	5	8.39	2880	413
3	12	1459.26	ompl	84	-	-	b	4-2	5-2	5	9.30	1525	260
3	13	1460.76	0	2	68	-							<u></u>
3	14	1462.71		2	-								
3	15	1464.92		92	-	-	b	2-1	31	6	4.91	4800	635
3	16	1466.59		4	-	-	Pos	sibly	H20 <sup>18</sup>	ъ	<sup>3</sup> -1	41	
3	17	1467.61		6	-	-	Pos	sibly	H_0 <sup>18</sup>	ъ	5-5	6 <sub>-5</sub>	
3	18	1469.3 🔹		2	-	-							
3	19	1472.0		98	-	- {	b	<sup>7</sup> -6 3 <sub>-1</sub>	7-4 4-1	7	1.77	<b>4</b> 41 6580	107 960
3	20	1473.44	+	94	-	-	ъ	5_5	6_5	7	3.46	7980	1410

	OB	DATA						CALCUL	ATED DAT	А	
ine mber	Wave		Intensit	ý			Identifico	ation	Wave	Inten	sity
Nn L	cm <sup>-</sup> ' (vac)	Solar	Air only	aboratory 110°C 1 (8 Atmo-M 8	IO°C Atmo-M		J'	J <del>″</del>	cm-' (vac)	14• C	110• c
321	1476.29		94	-	-	{b b	5-4	6-6 7-5	76.09 76.48	2500 800	440 190
322	1478.1		2	-	-		-7				
323	1480.4		1	40	-						
324	1481.33		62	-	-	{b b	5-2 8-1	6-4 9-3	81.37 81.73	890 36	180 18
325	1482.5		0	40	-						
326	1484.3		0	18	-	b	10_1	<sup>11</sup> -3	84.44	1.	.5 1.7
327	1486.27		40	-	-	þ	<sup>9</sup> -7	<sup>9</sup> -5	86.22	140	63
328	1487.34	- noi:	91	-	-	ъ	20	30	(87.34)	2970	390
329	1489.23	bsorpt	85	-	-	{b b	8 6-5 -1	8-3 7-3	89.11 89.19	343 456	120 122
330	1489.81	e	88		-	Ъ	6 <b>-</b> 5	<sup>6</sup> -3	89.81	1685	340
331	1490.81	plet	88	-	-	ъ	4-4	5-4	90.76	3890	590
332	1496.23	- Com	94	-	-	Ъ	4-3	5-5	96.27	11560	1760
333	1498.79		92	-	-	{b b	10 6_6	<sup>2</sup> <sub>6</sub> <sup>2</sup> <sub>-4</sub>	(98.79) 98.86	3145 610	390 120
334	1500.51		10	-	-	{Point b	ssibly <sup>9</sup> -4	also 9_2	н <sub>2</sub> 0 <sup>18</sup> ь 00.54	3-3 29	4-3 <sub>15</sub>
3 <b>3</b> 5	1501.83		52	-	-	b	7_4	7_2	01.83	280	77
336 <sup>j</sup>	1506		99	-	-	{b b b	11 5-4 3_3	<sup>2</sup> 1 5-2 4-3	05•57 06•68 07•09	11460 1065 14670	1370 180 1960
337	1508.55		97	-	-	Ъ	4-1	5-3	08.49	3340	560
338	1509.79		85	-	-	<pre>b b b b b</pre>	8 6-1 53	81 63 55	09•56 09•72 09•73	111 365 540	54 111 137
339	1510.56	ł	43	-	-	{b b	75 62	77 64	10.58 10.56	20 122	12 37

TABLE 1- Continued.

	ØB	OBSERVED DATA							CALCULA	TED DAT	7
ine mber	Wave		Intensity	1			Identific	ation	Wave	Intens	itỷ
Nu L	cm-' (vac)	Solar	Air only	borator 110° c (8 Atmo-M	y 110° C 8 Atmo-M	-	J'	J <sup>"</sup>	Number cm-' (vac)	14• C	110° C
340	1512.37		90	-	-	{b b b	$717^{-2}6^{-2}_{-3}$	73 70 6-1	12.22 12.30 12.34	159 185 1735	60 59 396
341	1515.01		72	-	-	Ъ	<sup>6</sup> -1	<sup>6</sup> 1	14.84	980	250
342	1516.77		94	-	-	{ b b	5 40 1	$524_{3}$	16.32 16.73	475 1420	103 267
343	1517.50	on "	96	-	-	{b b	<sup>3</sup> -2 4 <sub>2</sub>	$4_{4_{4}}$	17.59 17.95	4680 <b>470</b>	620 88
344	1520.36	sorpti	93	-	-	{ b b	5 5 -2	53 50	20.18 20.24	1470 1330	320 253
345	1521.4	te Ab	95	-	-	{ Ъ Ъ	4-3 5-5	4-1 5-3	21.25 21.25	5230 3870	760 650
346	1522.67	mple	92	-	-	Ъ	<sup>2</sup> -2	<sup>3</sup> -2	22.66	5300	650
347	1525.1	ο Ω Ι	72	-	-	Ъ	6 <sub>0</sub>	6 <sub>2</sub>	24.86	347	88
<b>34</b> 8	1525.52		92	-	-	Ъ	4-1	4 <sub>1</sub>	25.45	4100	670
349	1527.38		79	-	-	ъ	7 -5	7_3	27.48	1254	335
350	1528.66		75	-	-	ď	3	<sup>3</sup> 2	28.62	1240	179
351	1531.71	¥	55	-	-	b	7-1	71	31.64	595	185

Note: Between lines 351 and 352 the absorption has not been measured, since our resolution is not improved over that published[3]. The calculated spectrum from our energy levels yields results agreeing to  $\pm 0.2$  cm<sup>-1</sup> with the earlier work.

	OBS	DATA						CALC	CULATED DA	ТА	
ine	Wave		Intensity				Identifica	tion	Way	e Inte	nsity
Ž Ž	cm-' (vac)	Solar	Air only (8	oratory 110° C 110° Atmo-M 8 Att	C mo-M		J <sub>Z</sub> — .	J <u>"</u>	cm-' (		110° C
352	1761.88	+	88	-	-	b	<sup>6</sup> -3	5-3	61.86	6388	1070
353	1764.22		7	-	-	Po	ssibly	H2018	ъ	3 <sub>3</sub> 2	21
354	1765.40		6	-	-	р	10_6	10_8	65 <b>.3</b> 8	29	16
355	1767.3		20	-	-	Ъ	<sup>9</sup> -2	9_4	67.06	54	25
356	1768.20		87	-	-	{b b b	9-7 9-9 9-8	9-9 8-7 8_8	68.06 68.23 68.41	275 2910 970	88 740 250
357	1771.38		91	, <b>-</b> ,	-	Ъ	32	22	71.45	6050	725
358	1772.64		93	-	-	Ъ	33	21	72.80	17850	2140
359	1775.64	oti on	77		-	b	7_4	6-4	75,61	1320	260
360	1778.6	Absorp	22	-	-	{b c	$^{10}_{41}$ -3	<sup>10</sup> -5 31	78.57 78.64	55 8	33 8
361	1779.14	lete	49	-	-	{b b	8-4 10_5	7_2 10_7	79 <b>.</b> 2 79 <b>.</b> 15	210 84	59 46
362	1780.70	- Comp	84	-	-	{b b	40 8_6	3-2 7-4	80.77 80.92	1485 715	181 174
363	1781.96		28	-	-	b	<sup>9</sup> -3	<sup>8</sup> -1	82.04	80	33
364	1783.95		23	- )	-	Ъ	42	4-4	84.08	80	11
365	1784.93		76	-	-	Ъ	<sup>10</sup> -9	9-9	85.03	1810	580
366	1788.45		18	-	-	{b b	<sup>8</sup> 10 <sup>2</sup> <sub>-8</sub>	80 10-10	88 <b>.</b> 38 88.79	73 34	30 14
367	1790.1		37	-	-	{b b	10-7 11-7	109 11_9	90.09 90.0	104 27	42 18
368	1791.02		88	-	-	ъ	<sup>8</sup> -5	7_5	91.05	2210	540
369	1792.63		92	-	-	b	41	31	92.73	12050	1590
370	1795.22	+	41	-	-	b	6,	62	95.28	235	60

	OB	SERVED	DATA						CALCULA	TED DATA	
Line Number	Wave Number		Intensity Labor	atory			Identific	ation J#	Wave Number	Intensi	ty
	cm-' (vac)	Solar	Air IIC only (8)	D* C Atmo-M 8	110* C B Atmo-M		υ <sub>τ</sub>	52	cm-' (vac)	14* C	110° C
371	1796.03	ŧ	79	-		{b b b	5 63 72	53 61 70	95.90 96.19 96.39	790 710 165	170 180 53
372	1796.87		30	-	-	Ъ	81	8-1	96.98	217	89
373	1799.63		86	-	-	b	<sup>4</sup> 2	3 <sub>0</sub>	99.84	3860	506
374	1801.36		55	-	-	Ъ	1111	10_9	01.50	760	310
375	1802.43		65	-	-	Ъ	<sup>9</sup> -7	8-5	02.62	1065	320
<b>37</b> 6	1805.17		24	, , , ,	-	b	51	5_5	05.16	181	27
377	1807.75		52	Ŧ	-	Ъ	<sup>9</sup> -6	8-6	07.76	362	110
<b>37</b> 8	1808.62	tion —	3	-	-	{Ъ Ъ	11-9 11-8	11-11 11-10	08.6 09 <b>.4</b> 1	36 12	18 6
379	1810.63	sorp	81	-	-	ъ	5 <sub>0</sub>	<b>4</b> 0	10.71	2400	360
380	1812.22	Ab	54	-	-	b	<sup>9</sup> -5	8-3	12.26	342	106
381	1814.73	lete	4	-	-		2				
382	1815.6	Comp	3	-	-						
<b>3</b> 83	1817.47		<b>4</b> 0	-	-	b	<sup>12</sup> -11	<sup>11</sup> -11	17.52	290	154
384	1821.46		7	-	- ,	Ъ	10_4	<sup>9</sup> -2	21.50	13.1	6
<b>3</b> 85	1822.82		37	-	-	Ъ	10_8	9-6	22.87	157	61
386	1825.31		87	-	- •	Ъ	6-1 10-7	5-1 9-7	25.10 25.52	3940 478	690 186
387	1829.42		88	-	-	Ъ	<sup>5</sup> -1	4_3	29.24	1850	250
388	1830.26		91	-	-	ъ	<sup>5</sup> 1	4-1	30.14	6150	900
389	1833.39		22	-	-	ъ	<sup>13</sup> -13	<sup>12</sup> -11	33.39	100	72
390	1834.76	+	6	-	-	b	11,	11_1	34.52	4.2	5.1

\*

	OBS	SERVED	DATA						CALC	CULATED D	ата
ine	Wave		Intensi	ity			Identifico	ition	Way	e Inte	ensity
Ž	cm-' (vac)	Solar	Air	aboratory	IO° C Atmo-M		J' <sub>Z</sub> —	J <del>″</del>	cm-' (	vac) 14. c	110° C
391	1836.07	1	5	-	-	{b b	6 <sub>0</sub>	<sup>6</sup> -6	35.85	27	, 4.9
						(Po	ssibly	H20	b	43 3	3
392	1837.32		65	-	-	{b b	7-2 3 <sub>2</sub>	6_2 2_2	37•39 37•54	680 47	147 5
393	1839.29		4	20	-	c	43	33	39.1	11	12
394	1840.5		0	7	-	b	102	100	40.5	4.	5 3.9
395	1842.25		37	-	-	{b b	11_9 10_6	<sup>10</sup> -7 9_4	42 <b>.</b> 10 42 <b>.</b> 23	187 52	103 25
396	1844.24		92	-	-	{b b	11_8 43	10_8 33	43•44 44•20	62 16000	34 2300
397	1845.4	sorption	52	-	-	{b b b	65 75 83	63 73 81	45•57 45•54 46•09	260 210 87	79 80 42
398	1847.82	Ab:	67	-	-	Ъ	<sup>8</sup> -3	7_3	47.87	945	253
399	1848.80	ete	15	-	-	° p	14-13	<sup>13</sup> -13	48.80	31	30
400	1852.44	ompl	8	19	-	ъ	5 <sub>3</sub>	<sup>5</sup> -3	52.38	64	9.5
401	1854.1		2	4	-	ъ	44	4-2	54.12	9.7	1.3
402	1856.30		4	11	-	ъ	62	6_4	56.40	28.1	5.7
403	1858.52		32	70	-	ъ	9_4	8-4	58.48	145	50
404	1859.73		8	40	-	ъ	11 <b>_</b> 5	10-3	59.69	16.6	12
405	1860.96		10	39	-	ъ	12_10	11_8	60.96	22.1	15
406	1861.51		18	52	-	ъ	<sup>12</sup> -9	11_9	61.51	66.6	55
407	1863.1		0	4	-	c	53	41	63.6	4.5	5.7
408	1864.07	-	7	34	-	Ъ	15-15	14-13	64.06	8.8	12
409	1866.39		66	-	-	b	6 <sub>0</sub>	5_2	66.44	825	140
410	1867.94	ł	82	-	-	{b b	5 <sub>2</sub> 4 <sub>1</sub>	42 3_3	67.81 68.12	2400 388	390 47

TABLE 1- Continued.

	OB	SERVED DATA							CALCUL	ATED DATA		
ine mber	Wave		Intensity				Identifica	ation	Wave	Intensi	ty	
N <sup>n</sup> T	cm-' (vac)	Solar	Lab Air only (8	oratory 110° C Atmo-M E	110° C Atmo-M		J'	J <del>″</del>	cm-' (vac)	14• C	110° C	
411	1869.25	•	90	-	-	b	<sup>5</sup> 3	<b>4</b> 1	69.34	7220	1180	
412	1870.78		45	-	-	b	10 <sub>=5</sub>	<sup>9</sup> -5	70.91	179	80	
413	187667		5	11	-	b	7_1	7-7	76.59	25.0	5.2	?
414	1879.43	tion	12	50	-	Ъ	<sup>13</sup> -11	12_9	79.40	28.2	26	
415	1882.3	sorp	0	2	-							
416	1884.63	Ab	36	62	-	Ъ	6_2	<sup>5</sup> -4	84.64	243	37	
417	1885.28	lete	9	42	-	Ъ	11_6	10_6	85.36	22	13	
418	1889.59	Comp	87	95	-	Ъ	<sup>6</sup> 1	<sup>5</sup> 1	89.60	3715	709	
419	1892.63		2	17	-	Ъ	12_8	11_6	(92.63)	7.1	5.8	3
420	1893.87		7	25	_	ъ	77	75	93.63	52	23	
421	1895.27	Ļ	65	85	-	{Ъ Ъ Ъ	85 62 95	83 50 93	94.99 95.27 95.19	37 1230 18	21 230 14	
422	1897.57	<b>4</b> 0	5	32	•	Ъ	14_11	3-11	(97.57)	8.1	10.4	
423	1898.81	10	0	6	-							
424	1901.87	75	10	40	-	b	<sup>12</sup> -7	<sup>11</sup> -7	01.86	22	18	
<b>42</b> 5	1904.50	9 <b>9</b>	23	43	-	Ъ	5	4-4	04.40	159	21	
426	1908.12	]	53	82	-	b	7 <sub>0</sub>	6 <sub>0</sub>	08.12	555	128	
427	1910.09		68	87	-	Ъ	7_1	6_3	09.95	925	186	
428	1914.70	899	3	25	-	Po	ssibly	H2018	b 7 <sub>1</sub>	<sup>6</sup> -1		
429	1915.30		1	15	-	d	<sup>6</sup> -5	7-7	15.16	0.07	0.10	>
430	1918.08	J	82	_	-	ъ	55	<b>4</b> <sub>3</sub>	18.08	7660	1440	

TABLE 1 - Continued. OBSERVED DATA CALCULATED DATA Line Number Wave Intensity Identification Wave Intensity Number Number Laboratory  $J'_{T} - J''_{T}$ cm<sup>-'</sup> (vac) Solar cm-' (vac) 14. C 110° C 110° C (8 Atmo-M 8 Atmo-M Air 1919.3 4 9-7 9\_1 20.29 9.2 1920.7 3 b 99 8-1 665 1922.42 54 7-1 22.39 193 b 6\_1 340 1923.14 63 71 23.14 1490 b 1927.82 30 3 8\_2 9\_2 1933.09 97 33.23 86 14 b -1935-24 15 0 20 • 82 8-4 37.17 2.8 1937.87 40 0 40 ъ -9\_5 91 1939.2 15 0 9 b 39.51 4.7 87 1940.2 85 3 40.35 8.0 ъ 10-3 9-3 95 7-3 86 41.78 Ъ 41.82 1941.73 23 b 5.0 9.  $7^{7}_{3}$ 41.89 lb 9.5 63 64 42.61 {b b 3026 660 53 52 80 1942.49 99 42.86 1009 220 7-3 6\_5 1945.28 42 ъ 45.14 287 51 <sup>6</sup>-1 5-5 1946.25 46.24 364 55 39 ъ •

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5-5 11-5

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	OBS	SERVED	DATA					CALCUL	ATED DATA	
nber	Wave		Intensit	у		Identifico	ation	Wave	Intens	ty
Nur	cm <sup>-</sup> ' (vac)	Solar	Air	iboratory	alar a S	J'	J <del>″</del>	cm-' (vac	) 14• c	110° C
451	1961.19	96	26	95 -	ъ	8	7 .	61.25	108	25
452	1966.36	99	52	99 -	Ъ	-2 7 <sub>0</sub>	6-4	66.45	469	120
453	1967.49	99	66	99 -	ď	7	6,	67.54	1390	357
454	1970.0	10	0	52 -	Ъ	11_1	1 11_7	69.35	0.8	0.6
455	1975-43	5	2	32 -	{e e	50 8_8	6 9-1 9-9	72•75 75•84	0.2 0.1	0.3 0.3
456	1976.19	83	3	50 -	р	52	4-2	76.19	29	4
457	1977.6	14	0	- 8	Pc	ssibly	H2018	b é	5 5 <sub>5</sub>	
458	1981.32	40	2	50 -	e	6_4	7_5	80.83	0.25	0.45
459	1982.1	15	1	27 -						
460	1984.61	33	0	25 -	{b b	9 <sub>9</sub> 9-3	97 9-9	84.40 84.62	1.0	1.3 0.7
461	1988.53	99	56	<b>9</b> 9 -	{b b	44 81	$3_{-2}$ $7_{1}$	87•33 88•53	4 <b>.3</b> 570	0.5 184
462	1992.06	99	71	99 -	ď	65	55	92.14	2640	670
463	1992.6	99	66	99 -	р р р	8 9-1 7_2	70 8-3 6-6	92•50 93•31 93•35	211 183 65	68 65 11
-	-	-	0		e	7_6	<sup>8</sup> -7	95.08	0.26	0.5
464	1998.94	94	17	85 -	р	61	<sup>5</sup> -3	98.94	144	24
465	2000.90	10	0	20 -						
466	2003.20	8	0	10 -	$d_d$	<sup>2</sup> -1 1_1	3-3 2-1	02.96 03.29	0.18 0.20	0.16 0.17
-	-	0	-	0 -	е	5-2	6-3	04.14	0.31	0.4
467	2005.8	27	0	12 -	с	65	5 <sub>5</sub>	04.8	1.3	2.7
468	2007.66	95	26	92 -	р	90	80	07.77	67.5	28
469	2009.30	93	9	85 -	þ	8-4	7_6	09.47	37.8	7.9
470	2012.3	3	С	8 -	{Pc e	ussibly 40	H <sub>2</sub> 0 <sup>18</sup> 5-1	b 7 <sub>5</sub> 12.29	63.53	0.7

TABLE 1-Continued.

	OBS	SERVED	DATA						CALCULA	TED DATA	
mber	Wave		Intens	ity			Identifi	cation	Wave	Intensity	
Nr L	cm-' (vac)	Solar	Air	Laboratory	110* C 8 A1mo-M		J'_7 -	- J <u>′</u>	cm-' (vac)	14• C	110• C
471	2016.78)		73	-	-	b	75	63	16.76	1235	377
472	2018.30	99	50	-	-	b	<sup>9</sup> 1	8-1	18.77	202	83
473	2018.92		34	-	-	b	<sup>9</sup> -3	8-5	19.18	93	28
474	2022.95	86	30	95	-	b	10-1	9-1	23.09	62.5	34
475	2026.58	90	10	84	-	{b b	7 <sub>0</sub> 53	6_4 4-3	26.79 27.00	41.3 23.6	8.3 3.1
476	2030.03	16	0 *	10	-						
477	2034.03	52	0	4	68	{b b e	$11 \\ 81 \\ 5-4$	10-2 8-5 6-5	34.35 34.12 34.13	5.8 0.87 0.63	4.4 0.26 0.82
478	2037.54	75	5	8	80	ъ	10_2	9-4	37.54	15.3	7.0
479	2041.35	93	31	70	95	(b) (b) (b)	12-3 83 84	$^{11}_{73}_{72}$	41 • <u>35</u> 41 • 39 41 • 60	4.2 384 128	4.6 146 48
480	2043.97	93	9	18	86	b	<sup>8</sup> -3	7-7	44.00	102	21
481	2046.58	70	4	13	80	Ъ	100	9 <u>-</u> 2	46.64	20.8	11
482	2048.7	?	0	0	7	е	30	4-1	48.66	0.71	0.76
483	2050.8	?	0	0	11	b	84	<sup>8</sup> -2	51.08	0.37	0.12
484	2051.49	?	0	0	18	е	3_1	4-2	51.33	0.31	0.32
485	2053.06	?	0	0	12						
486	2053.96	?	0	0	16	{e e	4-4 4-3	5-5 5-4	53.84 54.47	0.96 0.31	1.06 0.34
487	2060.55	?	8	16	80	Ъ	<sup>8</sup> -1	7_5	60.55	92	22.2
488	2064.90	?	29	50	-	{b b	7 <sub>7</sub> 9 <sub>2</sub>	65 82	64.87 65.02	690 46.7	255 22.4
489	2065.83	?	13	30	-	Ъ	<sup>9</sup> 3	81	65.91	141	67.5
490	2068.74	?	0	3	8*	{e co	<sup>2</sup> 2 <sub>P(</sub>	3 <sub>1</sub> 18)	67.98 68.85	0.60	0.58

TABLE 1-Continued.

	OBS	ERVED	DATA						CALCULATED DATA		
ine mber	Wave		Intensit	у			Identific	ation	Wave	Intensi	ty
Su L	cm-' (vac)	Solar	Air	no• c	110° C		J' <sub>Z</sub> —	J <del>″</del>	cm-' (vac)	14• C	110° C
491	2072.9	?	0	<u>(6 Aimo-M</u>	12*		66 62 P(1	60 5-4	72.70 72.68 73.27	0.17 6.1	0.04 0.92
492	2074.22	?	0	10	.58	{e b	3-2 9_5	4-3 8-7	73.80 74.26	1.24 45.1	1.21 11.5
493 <sup>8</sup>	2077•3	?	0	8	14*	{e Co	2 <sub>0</sub> P(	3-1 (16)	76.83 77.65	1.12	1.02
494	2078.61	?	0	11	60	þ	11_1	10-3	78.67	*15.6	11.1
495	2081.98	?	0	10	34*	{co b	P( 10_4	15) 9 <b>-</b> 6	82.00 82.02	8.95	3.40
496	2086.52	?	0	8	17*	{ <sup>CO</sup> b	Р( <sup>11</sup> -3	(14) <sup>10</sup> -5	86.32 86.52	10.6	6.4
497	2087-54	?	5	13	61	b	101	91	87.54	45.4	28.7
498	2090.20	?	16	38	86	{b b	85 102	7 <sub>5</sub> 90	90.19 90.2	286 15•1	130 9•5
499	2093.36	?	-	-	17	{e e	2-2 2-1	3-3 3-2	93.02 93.73	1.64 0.53	1.40 0.46
500	2095.29	?	-	-	23*	{b e	83 44	8-3 43	95.14 95.28	0.90 0.72	0.31 1.01
501	2097.43	48	-	-	48	Ъ	9_4	8-8	97.48	14.1	3.53
502	2100.46	50	-	-	61	Ъ	<sup>9</sup> -2	<sup>8</sup> -6	00.46	13.9	4.20
503	2103.31	?	-	-	17*	{e co	1 <sub>1 P(</sub>	2 <sub>0</sub>	03.13 03.27	0.31	0.25
-	2105.93+	20	-	-	-	Ъ	5 <sub>5</sub>	4-1	05.86	5.35	0.77
504	2106.45+	40	-	-	29	р	6 <sub>3</sub>	5-1	06.45	14.4	2.52
505	2107.47	?	-	10	55	Ъ	110	100	07.9	4.25	3.61
506	2111.59	?	-	-	13*	{ª co	<sup>2</sup> 1 <sub>P(8</sub>	<sup>2</sup> -1	11.51 11,54	0.12	0.10
507	2114.86	80?	-	44	99	d e b b d	$3_{-3}$ $1_0$ $11_1$ 95 $2_{-1}$	2 - 1 2 - 1 10 - 1 $8_3$ 1 - 1	14.28 14.45 14.50 15.18 16.47	0.24 1.01 12.6 107.5 0.29	0.20 0.82 10.7 62.5 0.22

TABLE 1-Continu	ed.
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	OBS	SERVED	DATA						CALCULA	TED DATA	
ine mber	Wave		Intensit	у		]	Identifica	ition	Wave	Intensit	у
N <sup>n</sup> L	cm-' (vac)	Solar	Air only	110° C (8 Atmo-M	110° C 8 Atmo-M		J' <sub>Z</sub> — .	Jζ	Number cm-' (vac)	14* C	110° C
80 ئ	2119.48	?	-	-	7*	{e (co	4 <sub>1 P(</sub>	6) <sup>4</sup> 2	19.07 19.68	0.18	0.22
509	2121.54	20	-	-	28	{b e	7 <sub>2</sub> 4 <sub>2</sub> .	6-2 41	21.48 21.73	6.5 0.53	1.37 0.48
510	2122.58	2	-	-	25	{e e e	32 33 3_2	33 32 3_1	22.31 22.67 23.13	1.14 0.38 0.25	1.23 0.41 0.22
511	2124.27	30?	-	-	62	ъ	<sup>12</sup> -1	111	24.39	3.15	3.81
-	2125.00+	10	-	-	-	Ъ	71	<sup>6</sup> -5	24.87	13.7	2.36
512	2127.57	?	-	-	5*	co	P( .	4)	27.68		
513	2129.54	39	-	-	5				*		
514	2131.58	?	-	-	10*	{ª co	<sup>3</sup> -2 <sub>P(</sub>	3) <sup>2</sup> -2	31.58 31.63	0.10	0.08
515	2136.06	63?	9	31	96	Ъ	<sup>8</sup> 7	<sup>7</sup> 7	36.20	135.5	76.8
516	2137.12	55	0	-	75	{b b	64 84	5-2 7-3	36.86 37.35	3.5 18.3	0.60 4.83
-	2138.29+	15	2	0	?	b	10_6	9_8	<b>3</b> 8.38	5.6	1.88
517	2139.30	?	4	30	95	ъ	103	93	39.37	25.7	19.8
518	2141.3	?	0	-	2						
519	2142.4	δ	0	-	2	е	2 <sub>1</sub>	22	41.94	0.52	0.47
520	2145.41	52	0	12	77	{e b	<sup>2</sup> 10 <sup>2</sup> -3	<sup>2</sup> 1 9_7	44.63 45.64	1.57 20.3	1.40 7.8
521	2147.38	?	0	9	55	{ d b	4-3 11-5	3-3 10-7	45.68 47.42	0.27 7.8	0.24 3.9
522	2151.24	?	0	-	20	$\Big\{^{d}_{e}$	21 10	1 11	48.10 51.04	0.08 0.39	0.06 0.31
523	2152.58	45	0	-	50	b	<sup>10</sup> -5	9_9	52.75	16.3	5.15
524	2154.68	?	0	-	12	* {a {co	5-5 <sub>F</sub>	4-3	54.60 54.60	0.25	0.25

TABLE 1-Continued.

								and the second se		
	OBS	SERVED	UATA					CALCULA	TED DATA	
Ine	Wave		Intensity			Identific	ation	Wave	Intensi	y
Ž	cm-' (vac)	Solar	Air I only (8	HO* C HO* C Atmo-M 8 Atmo-	м	J'7 —	J <u>′′</u>	cm-' (vac)	14• C	110• C
525	2156.60	22	0	- 15	ъ	90	8-4	56.56	5.3	1.8
526	2158.26	?	0	- 6*	{e co	4 <sub>0</sub> R(	3)4-1	58.22 58.30	0.33	0.36
527	2161.75	?	0	- 84	b	97	85	61.88	51	36
528	{2162.90 2163.46	$\frac{10}{15}$	0	- 69	{b b	11 <sub>2</sub> 11 <sub>3</sub>	10 <sub>2</sub> 10 <sub>1</sub>	62 <b>.</b> 90 63 <b>.</b> 46	2.48 7.44	2•58 7•73
529	2165.50	?	0	- 8*	{co b	R( <sup>12</sup> -3	5) 12-9	65.60 65.68	0.05	0.05
530	2167.35	?	0	- 20	b	13 <sub>-1</sub>	<sup>12</sup> -3	(67.35)	0.60	1.06
531	2169.17	?	0	- 6*	Co	R(	6)	69.20		
532	2171.31	15?	0	- 15	р	73	6 <b>-</b> 3	71.38	10.7	2.12
533	2172.43	?	0	- 16*	$\begin{cases} e \\ d \\ CO \end{cases}$	20 6-6 -8(	2-1 5-4 7)	72.06 72.35 72.76	0.60 0.07	0.49 0.08
534	2175.09	10?	0	- 11	d	<sup>6</sup> -5	5 <b>-</b> 5	75.14	0.20	0.22
535	2176.45	8	0	- 7*	CO	R(	8)	76.28		
536	2178.99	?	0	- 8	đ	4-1	3-1	78.91	0.17	0.16
537	2179.87	?	0	- 10*	CO	R(	9)	79.77		
538	2181.41	?	0	- 40	ъ	<sup>10</sup> -1	<sup>9</sup> -5	81.61	8.84	3.98
539	2183.33	?	0	- 6	CO	R(	10)	83.22		
540	2185.41	?	0	- 38	{b b	8 12 <sup>0</sup> 1	7-6 111	85.24 85.37	2.0 1.9	0.42
541	2187.05	?	0	- 28*	{b b	12 102 5	11 9 <sup>0</sup> 5	87.0 87.04	0.65 16.6	0.92 15.8
542	2189.69	3	0	- 11	đ	7-7	6-5	89.52	0.13	0.16
543	2190.4	?	0	- 8	${co \atop a}$	R(: 7_6	12) 6-6	90.02 90.7	0.04	0.05
544	2191.83	?	0	- 2	đ	32	22	91.34	0.086	0.07

TABLE 1-Continued.

	OB	SERVED	ΠΔΤΔ	10000						CAL CUI	ATED DA	ΤΔ
	Wave		Intensi	tv		Identification				Wave		
Lin	Number		Γ ι	aboratory			J'	J <del>"</del>		Number	inter	
	cm-' (vac)	Solar	Air only	110° C (8 Atmo-M 8	IIO* C B Atmo-M		-2	-1		cm-' (vac)	14° C	110° C
545	2193.36	?	0	-	9*	{d Co	33 <sub>R(</sub>	13) <sup>2</sup> 1		92.66	0.25	0.22
546	2194.50	?	0	-	13	ъ	11-4	10_8	4	94•37	2.25	1.11
547	2196.69	?	0	-	5*	CO	R(	14)	4	96.66		
548	2198.63	?	0	-	11	е	<sup>2</sup> -1 <sup>1</sup> 0		\$	98.42	0.40	0.31
549	2200.29	?	С	-	25*	ъ	<sup>11</sup> -7	10_9	1	00.4	5.5	2.23
550	2201.2	?	0	-	4							
551	2202.0	?	0	-	2							
552	2203.23	?	0	-	4*	CO	R(	16)	(	03.16		
553	2205.19	?	0	-	60	р	9 <sub>9</sub>	87	(	05.30	19.3	18.1
554	2206.48	?	0	-	25	e	2_2	1 <sub>-1</sub>	(	06.53	1.74	1.31
555	2208.83	?	0	-	16	{e b	2 <sub>0</sub> 116	1 10-10	0	08.65 08.96	1.12 1.88	0.86
556	2209.9	?	0	-	5	ъ	131	12_1	(0	9.9)	0.42	0.87
557	2210.4	?	0	-	5	ъ	82	7_4	:	10.76	2.38	0.58
558	2211.50	?	0	-	40	{b b	11 <sub>5</sub> 11 <sub>-2</sub>	10 <sub>3</sub> 10 <u>-</u> 6		11.50 12.65	4.83 1.30	6.27 0.79
559	2214.0	?	0		8	е	30	21		14.09	0.80	0.70
560	2214.3	?	0	-	8							
561	2215.80	?	0	-	5							
562	2218:48	?	0	-	26	е	3_2	2_1	]	18.36	2.00	1.73
563	2221.87	?	0	-	4	đ	9_9	8-7		21.96	0.10	0.18
564	2223.61	?	0	-	4							

TABLE 1-Continued.

	OB	SERVED	DATA						CALCULA	TED DAT	Д
ne nber	Wave		Intensit	у			Identifi	cation	Wave	Intens	sity
LI	Number cm-' (vac)	Solar	Air only	110° C (8 Atmo-1	ту 110°С В Атто-М		J' <sub>7</sub> –	- J <del>″</del>	Number cm-' (vac)	1 <b>4</b> * C	110°C -
565	2225.84	?	0	-	8	b e	6 3_3_3	<sup>5</sup> 2 <sup>1</sup> -2	25.45 25.87	1.55	0.29 0.57
566	2227.30	?	0	-	16	d d b	4 43 63	<sup>3</sup> 2 33 5-5	27.09 27.45 27.59	0.07 0.22 0.06	0.08 0.24 0.07
5 <b>67</b>	2231.01	?	0	- 1	40	Ъ	107	<sup>9</sup> 7	31.01	7.10	9.0
568	2232.93	?	0	-	10	е	<sup>3</sup> -1	20	33.13	0.55	0.45
569	2235.62	?	0	-	25	b	123	113	35.62	1.30	2.30
570	2237.76	?	0	-	10	е	4-3	3-2	37.64	0.66	0.58
571	2239.30	?	0	-	8						

†, Frequency obtained from solar spectrum; g, probably due to  $\infty_2$ ; h, this line looks like an unresolved doublet; j, obviously consists of several unresolved components; \*, intensity variable, due to  $\infty$ contaminant.

$J_{\tau}$		V <sub>1</sub> V <sub>2</sub> V <sub>3</sub>		J <sub>T</sub>		<sup>V</sup> l <sup>V</sup> 2 <sup>V</sup> 3	
c	000	010	020		000	010	020
0 1_1 10 11 2_2 2_1 20 21 22	0.00 23.79 37.13 42.37 70.06 79.48 95.17 134.91 136.15	1594.59 $1618.41$ $1634.94$ $1640.48$ $1664.93$ $1677.07$ $1693.62$ $1742.51$ $1743.64$	3151.53 3175.4 3196.2 3202.0 3222.06 3237.77 3255.26 3316.0 3317.0	$\begin{array}{c} 4_{-4} \\ 4_{-3} \\ 4_{-2} \\ 4_{-1} \\ 4_{0} \\ 4_{1} \\ 4_{2} \\ 4_{3} \\ 4_{4} \end{array}$	222.04 224.83 275.48 300.38 315.73 382.49 383.86 488.16 488.17	1817.35 1821.63 1875.42 1907.94 1923.04 2004.89 2006.12 2129.60 2129.60	3375.3 3381.9 3438.6 3482.0 3495.85 3597.80 3598.8 3746.8 3746.8
$     3_{-3} \\     3_{-2} \\     3_{-1} \\     3_{0} \\     3_{1} \\     3_{2} \\     3_{3} \\     3_{3}     3_{3}     $	136.77 142.27 173.36 206.28 212.16 285.25 285.40	1731.92 1739.63 1772.30 1813.87 1819.16 1907.60 1907.71	3289.27 3300.1 3334.55 3387.8 3392.90 3500.4 3500.5	$5_{-5}$ $5_{-4}$ $5_{-3}$ $5_{-2}$ $5_{-1}$ $5_{0}$ $5_{1}$ $5_{2}$ $5_{3}$ $5_{4}$ $5_{5}$	325.36 326.59 399.45 416.12 446.50 504.00 508.79 610.12 610.34 742.10	1920.70 1922.80 2000.80 2024.24 2054.07 2126.44 2130.52 2251.67 2251.83 2406.24	3479.1 3482.5 3565.5 3598.6 3627.0 3719.2 3722.55 3868.3 3868.5 4052.8

TABLE 2. Energy Levels of  $H_20$  (cm<sup>-1</sup>).

TABLE 2 - Continued.

$J_{\tau}$		V <sub>1</sub> V <sub>2</sub> V <sub>3</sub>		JT		VlV2V3	
L	000	010	020		000	010	020
6_6 6_5 6_4 6_3 6_2 6_1 60 61 62 63 64 65,6	000 446.71 447.24 542.87 552.92 602.67 648.97 661.54 756.76 757.70 888.67 888.71 1045.15 586.28	010 2041.73 2042.73 2146.39 2161.31 2211.23 2271.60 2282.56 2398.39 2399.27 2552.95 2552.98 2734.24 2180.68	020 3600.4 3602.0 3713.1 3736.2 3784.3 3864.8 3873.55 4015.1 4015.8 4197.4 4197.4 4197.4 3738.6	9-3 9-2 9-1 90 91 92 93 94,5 96,7 98,9 10-10 10_9	000 1283.02 1341.03 1360.56 1475.14 1477.46 1631.44 1631.58 1810.76 2009.99 2225.56 1114.59 1114.59	010 2904.82 2983.43 2999.45 3139.65 3141.55 3321.0 3321.10 3526.77 3752.58 3994.39 2705.20 2705.22	4260.7
7-6 7-5 7-4 7-3 7-2 7-1 70 71 72 73 74,5 76,7	586.43 586.43 704.20 709.50 782.41 816.65 842.36 927.76 931.23 1059.68 1059.68 1059.89 1216.39 1394.85	2180.08 2181.27 2309.89 2318.48 2392.38 2440.06 2462.87 2569.66 2572.11 2724.15 2724.15 2724.30 2905.43 3110.02	3739.5 3879.05 3894.5 3967.4 4038.7 4053.0 74187.8 4190.8	$10^{-9}_{10}$ $10^{-7}_{10-6}$ $10^{-5}_{10-4}$ $10^{-3}_{10-2}$ $10^{-1}_{100}$ $10^{1}_{102}$ $10^{3}_{3}, 4$ $10^{5}_{5}, 6$ $10^{5}_{10}, 6$	1293.22 1293.80 1438.19 1446.23 1538.31 1581.53 1616.49 1719.36 1724.80 1875.24 1875.72 2054.55 2254.55 2471.59	2903.38 2904.68 3058.60 3072.95 3162.53 3224.80 3253.91 3383.65 3387.67 3565.00 3565.3 3770.95 3997.80 4241.0	
8-7 8-6 8-5 8-5 8-4 8-5 8-5 8-5 8-5 8-5 8-5 8-5 8-5 8-5 8-5	744.14 882.97 885.64 983.09 1006.14 1050.20 1122.78 1131.88 1255.19 1255.98 1411.59 1590.70 1789.09 920.22 920.22 1079.16 1080.51 1202.04 1216.37	2337.84 2490.42 2495.25 2595.9 2630.28 2670.75 2764.75 2771.67 2919.76 2920.26 3101.28 3306.58 3531.05 2512.37 2512.50 2688.26 2690.73 2818.40 2841.57	3894.8 4063.9 4071.6 4224.8 4383.9 4068.8 4266.7	109,10 11_11,-10 11_9 11_8 11_7 11_6 11_5 11_4 11_3 11_2 11_1 11_2 11_3 11_4,5 11_6,7 11_8,9 11_0,11	2702.09 1327.25 1525.02 1525.31 1690.85 1695.24 1813.47 1843.32 1899.21 1986.08 1999.34 2143.01 2144.46 2322.20 2322.25 2522.46 2740.73 2973.07 3216.6	2916.09 3135.9 3136.66 3315.0 3323.55 3441.22 3487.59 3532.75 3650.84 73660.20 3832.7 3833.86 4038.62 4038.70 4266.05	4470.1

TABLE 2 - Continued.

J <sub>T</sub>		<sup>V</sup> 1 <sup>V</sup> 2 <sup>V</sup> 3		J <sub>T</sub>		V <sub>1</sub> V <sub>2</sub> V <sub>3</sub>	
· ·	000	010	020		000	010	020
12_12,-11 12_10 12_9 12_8 12_7 12_6 12_5	1558.07 1774.75 1774.88 1960.38 1962.60 2106.7 2124.84	3144.77 3386.27 3386.53 3587.87 3592.71 3771.13		$ \begin{array}{r}     14_2 \\     14_1 \\     14_0 \\     14_1 \\     14_2 \\     14_{3,4} \end{array} $	3101.65 3264.2 3266.36 3465.18 3465.4 3485.6		
12_4 12_3 12_2 12_1 120 121 122 123,4 125,6 127,8 129,10 1211,12	2205.95 2275.65 2300.94 2434.14 2437.84 2613.26 2613.49 2813.94 3033.17 3267.2 3512.8 3767.1	3940.56 4123.73 4329.83 4330.0 4557.87		15-15,-1415-13,-1215-1115-1015-915-815-715-615-515-515-415-315-2	2358.58 2631.6 2872.56 2872.9 3081.2 3084.2 3252.0 3277.0 3365.0 3446.0 3473.0 3624.0 7629.7	3937.87	
13_13,-12 13_11 13_10 13_9 13_8 13_7 13_6 13_5 13_4 13_3 13_2 13_2 13_2 13_1	1806.94 2042.5 2042.5 2247.0 2248.24 2415.95 2426.0 2534.14 2586.5 2629.54 2748.4 2756.61	3391.46 3654.28 3877.9 4443.0		$15_{-1}$ $15_{0}$ $15_{1}$ $15_{2,3}$ $16_{-16,-15}$ $16_{-14,-13}$ $16_{-12,-11}$ $16_{-10}$ $16_{-9}$ $16_{-8}$ $16_{-7}$	3628.7 3824.8 3826.1 4045.8 2661.2 2953.0 3211.5 3437.7 3439.7 3640.3 3657.1	4237.5	
$130^{1}$ $131^{1}$ $132,3^{1}$ $134,5^{1}$ $136,7^{1}$	2927.38 2928.45 3128.25 3348.2 3584.0	4644.0		17-17, -1617-15, -1417-13, -1217-1117-10	2931.5 3291.0 3567.5 3011.7 3812.8	4554.6	
$14_{-14}, -13\\14_{-12}, -11\\14_{-10}\\14_{-9}\\14_{-8}\\14_{-7}\\14_{-6}\\14_{-5}\\14_{-4}\\14_{-3}$	2073.81 2328.2 2551.0 2551.5 2740.5 2745.5 2883.5 2919.5 2983.6 3085.0	3655.74 3940.1		17_9 17_8 18_18,-17 18_16,-15 18_14,-13 18_12,-11	4047.1 4057.3 3319.4 3648.0 3940.8 4201.8		

TABLE 3. Series regularities in the  $P_{1,\overline{1}}$ ,  $P_{\overline{1},\overline{1}}$ ,  $P_{\overline{1},1}$ , and  $P_{\overline{3}1}$  branches of  $\nu_2$ 

3,1						4	Э	6	7	8	9	10	11	12	13	14	15	16
3,1			- 253					(1000 0)		(1000 = 20)	(1000 00)		1005.0					
	6	- *					1085.4	(1066. 2)	1051, 28	(1039.53)	(1029, 69)	1019.5	1007.3					
3,1	4				1004.05	1180.75	1165.4	1151, 59	1137, 46	1120.9	1099, 74	1072,69	1039.53	0.01 0.0				
3,1	2			(1050 0)	1284.37	1265.42	1244, 18	1218, 63	1187.00	1149, 48	1106.76	1060.14	1010.86	961.09				
3,1	0			(1379, 6)	1349, 39	(1313, 64)	1271.80	1225. 08	1174. 54	1121.24	1066.20	1010, 12	953.49	896.65				
$\frac{1}{7},\frac{1}{7}$	1		1498.79	1464, 92	(1423.9)	1375.09	1320.09	1260.38	1198. 22	1135, 80	1074.46	1014, 56	955.71	897.77				
1,1	0	(1557.5)	1538, 8	1522.67	1507	1490, 81	1473, 44	1455, 26	(1436, 7)	1417.39	1397.74	1378.04	1358.06	(1338.5)	1317.6	(1296.7)	1276, 63	1255.95
1,1	1		1564, 7	1540.1	1517.50	1496.23	1476.29	1456.49	(1436.7)	1417.39	1397.74	1378.04	1358.06	(1338, 5)	1317.6	(1296.7)	1276.63	1255.95
1,1	2		1505.8	1487.34	1472.0	1459.26	1447.91	(1436.7)	1423, 90	1409.94	(1394, 5)	(1378, 0)	(1361, 1)		1326.14			
1,1	3			(1569, 4)	(1538.8)	1503.55	1481, 33	(1457.0)	1435. 77	1416.11	(1397.7)	1379.63	(1362, 7)	1344.04				
1,1	3			1457.09	1429.97	1399, 16	1362.70	1318, 97	1268.40	1212.28	1152.44	1091, 24	1030.58	(970.66)				
1,1	4			1458.24	1436.72	(1419.3)	1404.98	(1394, 5)	1386.51	1379, 63	1372, 28	(1363, 2)	1352, 41	(1340)	(1326.14)			
1,1	5				1591.9	1558, 5	(1522, 7)	1489, 23	(1457.1)	1428.31	1403.54	1382, 11	(1362.7)	1345. 51				
1,1	5				1419.3	1394.50	1368.60	1340.35	1308.25	1269.97	(1225.0)	1173.76	1117.71	1058.69				
1,1	6				1419.3	1395, 81	1373, 76	1354.87	(1340, 3)	1329.90	1323, 31	(1319, 0)	1316, 20					
1,1	7					1387.55	1363, 17	1338.56	1313.64	1287.38	1258.63	1225.5	(1187.0)	1141.64				
1,1	8					1387.55	1363.17	1339.55	1317.04	1296.67	1280.09	(1268, 4)	1257.07					
1,1	9						1361.09	1336.64	1312.61	1288, 28	1264.04	1239, 25	1213.0	1184.14				
1,1	10						1361.09	1336.64	1312.61	1288, 92	1266.11	1244.77	1226.1					
1,1	11							1339. 23	1314, 82	1290.59	1266.63	1242.90	1219, 1	1195.41				
1,1	12							1339. 23	1314, 82	1290.59	1266.63	1242.90	1220, 43					
1,1	13	]		1. S. S. S.					1000 00	(1000 07)	1050 05	1010 10	(100* 00)	1001 55				
1,1	14	<u></u>	-,						1320, 90	(1296, 67)	1272, 37	1248, 52	(1225, 08)	1201, 55				
1,1	15	1							1.2.2.3		1001 00	(1055 1)	1000 01	1000 -				
1.1	16	<u></u>								1305, 60	1281, 22	(1257.1)	1233, 31	1209.79		1.1		
1.1	17	1						Sheers.						1.1		2141 1		
11	18										1292, 40	(1268, 4)						

## TABLE 4. Lines in the $R_{3,\overline{1}}$ , $R_{3,\overline{3}}$ , $R_{5,3}$ , and $R_{5,\overline{5}}$ branches, pure rotation

Branch	$(J+\tau)^{\prime\prime}$	J'' = 6	7	8	9	10	11	12	13	14	15	16	17
3,1	0	370.16	419, 98	472. 54	526,08	580, 7	635, 3	(690, 3)	(744, 1)	799.11	852, 91	906.32	959.33
3,3	1	484,05	545. 55	(616.4)	(696.7)	784.54	878.61	976.07			1.1.2		
3,1	2	385.06	418.57	(457.9)	502.31	550.18	599.82	(650.5)	(703.0)	(756)	808.14	859.78	910.77
3,3	3	506.96	546.50	592.0	(644.7)	(705.5)	775.63	854.66	941, 12	1999		Star Sec.	
3,1	4	(457.0)	(472.5)	492, 08	517.00	547.89	574.74	(625.3)	(672.5)	(726)	(784, 54)	840.36	
3,3	5	566.9	594.5	625. 25	(659, 2)	(696.6)	(742.6)	794.01	853, 41	921.48			
3,1	6	554, 82	569.2	(580.7)	(592.0)	605.0	620.7	(641.4)	(669, 6)	(707)			
3,3	7	(638.1)	(663.0)	(688, 0)	(713.5)	(740.7)	770.17	803.60	841.16	883, 18	0.000-000		
3,1	8	(637.2)	(659.4)	(678.9)	(694.0)	(705.2)	(713.0)	(721.5)	(730.8)	(741)			
3,3	9		(729.4)	(754.8)	779.36	803, 09	827, 80	(852.62)	878.9	906.8		2971331	
3,1	10		(729.2)	(754.0)	777.07	797.65	814.61	827. 21	835.64	(841.1)		-01201 B	
3,3	11			813.97	840.01	865.51	890.14	914.06	937.40	960.77			
3,1	12			813.97	840.01	865,02	888.71	910.17	929,00	~1.1.1.1.2.0.3			
3,3	13	l			801 33	018 59	044 05	070 66		1.1.1.1.1.1.1.1		1.72	
3,1	14	]			001.00	010, 02	544, 50	310.00	1.1.1.1.1.1.1	A Constraint of the	Star South A		
3,3	15					062 05	000 3						
3,1	16	}				302.00	330. 5		1.1.1.1				
5,3	0	(613.0)	(668.9)	(731.0)	798.75	871.36	948.35	(1028.5)					
5,3	2	(673.5)	(707.4)	(748.5)	796.01	849.69	909.02	973. 59	1042, 57				
5,3	4	792.2	808.33	827.80	852.62	883. 89	922.19	967.00	1017.9			1000	
5,3	6		946.73	959.78	971, 43	984.2	1000.35	1022.0					
5.5	1	769.1	825, 24	887, 33	955, 32	1029.69			1.1				
5,5	3	841.97	881, 15	925.05	974.04	1028, 47	1088,1	28-10 C - 10		1.1.1.2.2.2.2			

TABLE 5. Rotational constants for  $\nu_2$  of  $H_2O$ 

	Constant									
<i>v</i> <sub>1</sub> <i>v</i> <sub>2</sub> <i>v</i> <sub>3</sub>	A	В	С	$D_A$	$D_B$	Dc				
010	$cm^{-1}$ 31. 12 27. 877	$cm^{-1}$ 14.66 14.512	$cm^{-1}$ 9. 15 9. 285	$10^{-4}cm^{-1}\ 56.7\ 33.7$	$10^{-4}cm^{-1}$ 2.6 2.39	$10^{-4}cm^{-1}$ 0.4 .25				

WASHINGTON, November 9, 1951.