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U. S. DEPARTMENT OF COMMERCE NATIONAL BUREAU OF STANDARDS WASHINGTON

The Ionosphere and Radio Transmission Conditions, with Special Reference to the Observing and Reporting Service of the National Bureau of Standards

December 9, 1939.

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(December 9, 1939) (Replaces Letter Circular 499)

THE IONOSPHERE AND RADIO TRANSMISSION CONDITIONS,

WITH SPECIAL REFERENCE TO THE OBSERVING AND REPORTING

SERVICE OF THE NATIONAL BUREAU OF STANDARDS.

Long-distance radio transmission is made possible by reflection of the radio waves from ionized (electrically conducting) layers in the ionosphere, the upper region of the earth's atmosphere. How far the radio waves go, and what frequencies may be used, are determined by the heights and the degree of ionization of the ionized layers, which are located between about 50 kilometers (30 miles) and 400 kilometers (250 miles) above the earth's surface. Because of variations in the ionized layers, the conditions of radio wave transmission vary with time of day, with season, and from year to year. For the efficient maintenance of radio services it is important to have a constant supply of information on these changes. To this end an ionosphere observing and reporting service is provided by the National Bureau of Standards. It is analogous to the weather reporting service, though quite independent of it; the factors producing or affecting weather exist in much lower levels of the atmosphere than the ionosphere.

The results of the Bureau's ionosphere work are made available to the public by a radio broadcast each week and a publication each month as well as through special papers (see Bibliograp) at end hereof) published from time to time on particular phases of the subject. The weekly broadcast, from the Bureau's station WWV each Wednesday, reports ionosphere and radio conditions for noon of the day of the broadcast and for the preceding midnight, and also special ionospheric events of the preceding week. The time schedule and frequencies of the broadcast are given on page 7; other details are given on pages 6, 9, and 18 hereof. The monthly publication, in Proceedings of the Institute of Radio Engineers (address: 330 W. 42d St., New York, N.Y.), gives the average ionosphere and radio conditions for the second month before, and a prediction of radio conditions for the month following, the month of publication.

THE IONOSPHERE

In the high atmosphere, above about 50 kilometers (30 miles), the air particles are separated so far that collisions between them are far less frequent than in the lower atmosphere. and when an air particle is ionized by ultraviolet radiation from the sun it remains lonized for a considerable time. Therefore at any given time a large proportion of the air particles are in an ionized condition. This does not occur much below about 50 kilometers (30 miles), because the ionizing radiations from the sun are largely absorbed in the higher regions of the atmosphere. Likewise there is not very great ionization density above about 400 kilometers (250 miles), because the air is so rare at such heights that there are not enough atoms to provide for great ionization density. The region in which the ionization is great, great enough in fact to affect radio wave transmission, is thus between 50 and 400 kilometers (30 and 250 miles) above the earth's surface, and this region is called the ionosphere.

The ionization in the ionosphere is not uniformly distributed with altitude but is stratified, and there are certain definite layers in which the ionization density is such as to reflect radio waves. These layers do not remain/at the same height but vary diurnally, seasonally, and otherwise, in both height and ionization density. There may be a considerable number of such layers at a given time. There are two principal ones, called the E and F layers. The E layer is at a height of 90 to 140 kilometers at different times, usually about 110 kilometers. The term F-layer is ordinarily reserved for the other layer as it exists at night; in the daytime during most of the year it divides into two layers which are called the F_1 and F_2 . The night F layer is at a height of about 180 to 400 kilometers. The F_7 layer exists in the daytime, at a height of about 140 to 250 kilometers. The Fo layer exists in the daytime, at a height of about 250 to 350 or more kilometers in the summer, and about 150 kilometers in the winter day. (The "virtual" heights, defined later, are somewhat higher than these values). The fourth layer, which is semi-permanent, is the D layer; it exists only in: the daytime, and its height is of the order of 50 to 90 kilometers. Little has been done on the determination of the quantitative characteristics of the D layer, its effects being largely inferred rather than directly observed. Existing knowledge covers mainly the E, F, F1, and F2 layers.

The structure of the ionosphere may be visualized in an elementary way from Fig. 1, which is for a typical summer daytime condition, the E, F_1 , and F_2 layers all being present. This diagram is drawn to scale, so the angles of reflection of



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radio waves from the layers may be estimated correctly. The three layers are shown as mere thin lines, for simplicity. The layers have in fact a certain thickness, and the density of ionization varies somewhat in this thickness. At the right of the diagram is a rough illustration of a possible distribution of ionization density with height.

Dotted lines indicate two of many possible paths of radio waves from a transmitter to a receiver as transmitted by reflection from the ionosphere layers. This picture, simple as it is, does in fact represent the basic mechanism of radio wave transmission over long distances. When we consider the variations of ionization and height of the layers with time, and the effects of the ionization upon the received intensity and the limits of transmissible frequency at any particular time, the picture loses its simplicity. However, most of the phenomena of long-distance radio transmission are completely explainable in terms of the ionosphere.

Ionosphere Characteristics .- The principal ionosphere characteristics which control or determine long-distance radio transmission are the height and the ionization density of each of the ionosphere lavers. Since each layer has a certain thickness it is necessary to define the sense in which the term, height, is used. When a ray or train of waves is reflected by a layer, it is slowed down as soon as it starts to penetrate into the layer. The process of reflection thus goes on from the point at which the waves enter the layer until they have been fully turned down and leave the layer. This is true whether the waves travel vertically or obliquely to the iono-It is illustrated for the oblique case in Fig. 2. sphere. The waves follow a curved path in the layer until they emerge at the same vertical angle at which they entered. The time of transmission along the actual path BCD in the ionized layer is the same as would be required for transmission along the path BED if there were no ionized particles present. The height h from the ground to E, the intersection of the two projected straight parts of the path, is called the virtual height of the layer. This is the important quantity in all measurements and applications.

The virtual height of a layer is measured by transmitting a radio signal from A, and receiving at F both the signal transmitted along the ground and the echo, or signal reflected by the ionosphere, and measuring the difference in time of arrival of the two. The signal is a special, very short pulse in order that the two may be separated in an oscillograph, as the time differences are mere thousandths of a second. The difference between the distance (AE +EF) and AF is found by multiplying the measured time difference by the velocity of light. From this and the known distance AF, the virtual height h is cal-

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culated. It is usual to make AF zero, i.e., to transmit the signal vertically upward and receive it at the same place (and it is for this case that the term "virtual height" rigorously apolies). The virtual height varies somewhat with frequency.

The effectiveness of the ions in reflecting the waves back to earth depends on the number of ions present in a unit of volume, i.e., the ionization density. The higher the frequency, the greater is the density of ionization required to reflect the waves back to earth. It has been shown that, for electron ionization, the relation (for the ordinary ray, explained below) is $N = 0.0124 f^2$, where N is the number of electrons per cubic centimeter and f is the highest frequency, in kilocycles per second, at which waves sent vertically upward are reflected back to earth. Waves of all frequencies higher than this pass on through the ionized layer and are not reflected back to earth, while waves of all lower frequencies are reflected. This frequency is called the critical frequency, and measurement of it is, with the equation just given, a means of measuring the maximum ionization density in an ionized layer. (Waves of higher frequencies than the critical are sometimes reflected, by another mechanism see discussion of "Sporadic E", below).

Measurements of critical frequency are usually made by means of vertical or nearly vertical transmission (i.e., with the transmitter and receiver not far apart). The process is to measure the virtual height, by the method described above, repeating the determination at successively increasing frequencies until the waves are no longer received back from the layer. The highest frequency at which waves sent vertically upward are received back from the layer is the critical frequency of that layer. Typical results of such measurements are illustrated in Figs. 3, 4, and 5, for different times of year, day and night. They show critical frequencies as sharp increases in virtual height.

In Fig. 3, starting at a frequency below 2000 kc/s, the virtual height is found (in this example) to be about 110 kilometers, and remains at about this height until about 3300 kc/s. The critical frequency of the E layer at the time of this measurement is thus 3300 kc/s, i.e., this is the highest frequency at which vertically incident waves are reflected back to earth; all such waves of higher frequency penetrate through the E layer and go on up to a higher layer, the F_2 . The F_2 layer has a greater ionization density and so it reflects back waves of frequency greater than 3300 kc/s. It is not until frequencies greater than 11,500 kc/s are used that the F_2 layer fails to reflect them, in the case illustrated.

Near the critical frequency the waves are excessively retarded in the ionized layer, which accounts for the rise







of the curve at the critical frequency. At the right of the curve appear two critical frequencies for the F_2 layer. This is an indication of double refraction of the waves due to the earth's magnetic field, two components of different polarization being produced. One is called the ordinary wave and the other the extraordinary wave. The symbols o and x, respectively, are used for these components. The critical frequency of a layer n is represented by the symbol f_n , and to such symbol the o or x is added as a superscript. Thus the critical frequency waves are indicated by the respective symbols, f_F^0 and f_X^X .

In the case of the E layer, the ordinary wave usually predominates and the extraordinary wave is so weak it does not affect radio reception. At Washington the critical frequency for the extraordinary wave is about 750 kc/s higher than for the ordinary wave (for frequencies of 4000 kc/s or higher). The difference in frequency is proportional to the intensity of the earth's magnetic field at the place of reflection, and is therefore different at different places on the earth. In reporting results of measurements of critical frequency it is now customary to give the values for the ordinary wave; practice varied in the past.

Besides the virtual heights and critical frequencies, the absorption of the energy of radio waves by the ionosphere is an important factor in limiting radio transmission. This absorption exists because the ions set in motion by the radio waves collide with air molecules and dissipate as heat the energy they have taken from the radio waves. Consequently the energy thus absorbed from the radio waves is greater, the greater the distance of penetration of the waves into the ionized layer and the greater the density of ions and air molecules in the layer, and hence the greater the number of collisions between ions and air molecules. Absorption is especially great in the daytime, and it occurs chiefly in the low ionosphere, in the D or E layers. It also occurs in the high ionosphere, near critical frequencies. Much of the lowlayer absorption disappears with the decrease of low-layer ionization at night. Higher frequencies are less affected by absorption than are lower frequencies, for waves passing through the same ionized layers.

Regular Variations of Ionosphere Characteristics. - There are three principal types of variation of critical frequencies which are fairly regular with time. These are diurnal variations, seasonal variations, and year-to-year variations with the sunspot cycle.

The diurnal and seasonal variations of the critical freduencies of the normal E layer are particularly regular. The critical frequencies vary with the altitude of the sun, being highest when the sun is most nearly overhead. Thus the diurnal maximum of the E critical frequency (f_E) is at local noon, and the seasonal maximum is in midsummer. At night this layer usually does not reflect waves of frequencies higher than about 1000 ke/s.

The diurnal and seasonal variations of the critical frequencies of the F_2 layer at Washington are ouite different from those of the E layer. The winter F_2 critical frequencies exceed any regular critical frequency found during the summer. In the winter a broad diurnal maximum occurs in the daytime, centered around 1:00 P.M. local time. In the summer a broader diurnal maximum centers about sunset. During the night the winter critical frequencies are usually lower than the corresponding summer values. Thus, the highest F_2 -layer critical frequencies occur during the winter night; the summer day and night values are between.

The F₂ virtual heights are much lower during a winter day than during a summer day. The F virtual heights at night are about the same in winter as in summer.

The seasonal effects in the ionosphere synchronize with the sun's measonal position, not lagging a month or two as do the seasons of weather. Winter conditions in the F_2 layer obtain during a period of several months from about October to March, and summer conditions for a period of several months from about April to September. Around the equinoxes, there is a transition period of a month or two in which the change occurs between winter and summer conditions; in these transition periods the ionosphere characteristics (except for the effects of sporadic E - see below) fluctuate between winter and summer conditions and are thus more erratic than during the rest of the year.

There are important changes in ionosphere characteristics in the ll-year subspot cycle. From the subspot minimum in 1933 to the subspot maximum in 1937 the F- and F₂-layer critical frequencies doubled (for most hours of the day), and the E-layer critical frequencies became 1.25 times as great. A consequence is that the best radio frequencies for long-distance transmission were approximately twice as great in 1937 as in 1933 (except for summer daytime, when they were about 1.5 times as great). In about 1944 they will return to minimum values.

The condition of the ionosphere varies somewhat with latitude. For all latitudes of continental United States the differences from the Washington values appear to be negligible, but the values in Alaska and in the Canal Zone are somewhat different.

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The Bureau's monthly report on the ionosphere in the Proceedings of the Institute of Radio Engineers includes average data for the month on the virtual heights and critical frequencies of the ionosphere layers, and data on their variations. The data are given graphically, showing values for all hours of the day and night. Examples of these monthly graphs are given in Fig. 6. These examples show summer and winter conditions at minimum and maximum phases of the sunspot cycle. The values shown for the minimum are those observed in 1933, which are expected to be approximately the same in about 1944.

The Bureau's weekly broadcasts of ionosphere and radio data are given each Wednesday (except nationally legal holidays) from its station WWV. The broadcast is by voice on each of three radio carrier frequencies, as follows:

1:30 to 1:35 P.M., EST, 10 megacycles per second 1:40 to 1:45 P.M., EST, 5 megacycles per second 1:50 to 1:55 P.M., EST, 20 megacycles per second.

The broadcast includes statements of the virtual heights (minimum values) and critical frequencies of the ionosphere layers, limits of variation of the critical frequency throughout the day, and maximum usable frequencies for radio transmission; values are given for noon of the day of the broadcast and for the preceding mignight. In addition, data are given on ionosphere disturbances occurring throughout the preceding week. (Information on other services broadcast from the same station is given in the Bureau's Letter Circular, "Standard frequencies and other services broadcast by National Bureau of Standards"). Details of the ionosphere broadcast are shown in the convenient recording form for this broadcast attached hereto as Appendix 1.

APPLICATIONS TO RADIO TRANSMISSION.

From the vertical-incidence critical frequencies and virtual heights of the ionosphere layers, at any given time, it is possible to calculate the upper limit of radio frequency that can be transmitted over any distance. The calculated values of maximum usable frequency are found to agree with direct observation of radio transmission over such distances. The maximum usable frequencies are nearly the best frequencies to use see "Optimum Frequencies", below.

When radio waves are transmitted along the earth over any distance, they strike the ionosphere obliquely (Fig. 1). Such obliquely incident waves can be reflected back down with lower ionization densities than can vertically incident waves. It results that the larger the angle of incidence (angle of wave path with the vertical), i.e., the greater the transmission distance, the higher is the upper limit of frequency of waves that can be reflected from a layer of given ionization density or critical frequency. This upper limit of frequency, for transmission via an ionosphere layer, for a particular time and transmission distance, is called the maximum usable frequency. It may be calculated roughly, to a first approximation, by multiplying the critical frequency of the layer by the secant of the angle of incidence.

The accurate calculation of maximum usable frequencies from vertical-incidence critical frequencies is complicated (see references 20, 22, 24, in Appendix 2). For convenience, typical values for the conversion are given in Table 1.

Table 1.

Typical Average Ratios of Maximum Usable Frequency

to Critical Frequency		
(for One-Hop Transmission)	,	
	-	

	Distance, km					
	500	1000	1500	2500	3500	
Winter Midnight F Noon F ₂	1.2 1.2	1.5. 1.6	1.8 2.1	2.8 2.9	2.9 3.4	
Summer Midnight F Noon F ₂ Noon F ₁ Noon E	1.2 1.2 1.3 2.0	1.4 1.5 2.0 3.4	1.7 1.8 2.7 4.4	2.4 2.5 3.6	2.8 2.9 -	
Sporadic E*	2.5	4.2	5.1		~	

*Sporadic E transmission has no critical frequency. The values given are ratios of maximum usable frequency to the approximate upper limit of frequency of the stronger sporadic-E reflections at vertical incidence.

To obtain the maximum usable frequency for transmission over a given distance by way of a given layer, multiply the critical frequency by the ratio given in the table. Where blanks appear in the table, and for distances over 3500 kilometers, the distance is too great for one-hop transmission, i.e., transmission over such distances requires multiple reflection from the ionosphere with intervening reflection from the ground. Information is given below on such multihop transmission, for which calculations must be made separately for each hop. SUNSPOT MINIMUM

SUNSPOT MAXIMUM

SUMMER

WINTER



FIG. 6

The distance at which a given frequency is the maximum usable frequency is also the minimum distance over which that frequency is receivable. This minimum distance for any frequency is called the skip distance; at any less distance it is impossible to receive on that or higher frequencies (except for sporadic or scattered reflections - see below).

The Bureau's monthly reports and weekly broadcasts include the maximum usable frequencies for various distances, in addition to the ionosphere data (critical frequencies and virtual heights). The nature of the weekly radio broadcast is described above and is indicated also in Appendix 1. Examples of the data on maximum usable frequencies, as published each month (in Proceedings of the Institute of Radio Engineers), are given in Fig. 7. Each point on the graphs gives not only the upper limit of frequency usable over the distance, but also conversely gives the skip distance, i.e., the lower limit of distance over which the frequency gives satisfactory sky-wave transmission.

Each graph in Fig. 7 shows, for each hour of the day and night, the average values for the month. The examples in Fig. 7 show summer and winter conditions at minimum and maximum phases of the sunspot cycle (i.e., for the same months and years as Fig. 6).

For a given month, the shape of the graphs changes little from year to year, but the absolute values change materially. The general level of absolute values changes gradually, and it is therefore possible to predict specific values of maximum usable frequencies months in advance. This is regularly done by the National Bureau of Standards. In its published monthly report the Bureau includes a predicted set of graphs of maximum usable frequencies for the month following the month in which the publication appears. These predictions are well verified by the measurements made subsequently.

The highest maximum usable frequencies (in north temperate latitudes) occur during the winter day and the lowest during the winter night. The summer values for both night and day lie between these two extremes except as modified by sporadic-E or scattered reflections (see below).

Maximum Usable Frequencies over Long Paths.- Since the local time of day, and hence the ionosphere characteristics, may vary a large amount throughout a long transmission path, it is necessary to consider what part of the path determines the conditions of transmission. For single-hop transmission, i.e., for transmission by a single reflection from the ionosphere, it is the region half-way between the transmitter and

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receiver whose conditions determine the transmission, because it is there that the reflection from the ionosphere takes place. In the case of multi-hop transmission, i.e., when the radio waves are reflected from the ionosphere, then from the ground, then back to the ionosphere, etc., the determining conditions are in the middle of each hop.

The maximum possible distance of transmission by a single hop, i.e., reflection from any one ionosphere layer, is limited by the geometry of the earth's surface and the layer, and also by absorption or other limitation at the ground of those waves which are nearly tangential to the earth's surface. It is found in practice that the minimum angle with the ground of the radio waves transmitted or received (over land) averages about 3 1/2 degrees. From these considerations the geometry indicates that the maximum distance along the earth by a single hop is ordinarily about 3500 kilometers for the F₂ layer, and about 1700 kilometers for the E layer. Single-hop transmission may sometimes be possible at greater distances than these while at the same time multi-hop transmission over the same path may be more efficient.

The curves of maximum usable frequencies are drawn for single-hop transmission; they are nevertheless available also for solving problems of multi-hop transmission. Calculation of the maximum usable frequency for multi-hop transmission is necessarily somewhat complicated. In the first place, it is necessary to consider the time of day of the locality where each reflection from the ionosphere layer takes place. The maximum usable frequency is the lowest one of the several corresponding to the times of day at the localities where reflection takes place, i.e., at the midpoints of the successive hops. To a first approximation these several freduencies are obtained from the curve for that distance which is equal to the total transmission distance divided by the number of hops.

Because of the variation of ionosphere characteristics with longitude, different frequencies may be necessary for transmission in different directions from a given place. For example, around sunset in winter lower frequencies are used in transmitting eastward than in transmitting westward from the same location. This does not mean, however, that different frequencies would be necessary or desirable in opposite directions over the same path.

For very long paths in which widely different longitudes (i.e., times of day) are involved, it sometimes happens that the waves travel different parts of the way by different layers. For such cases, it is necessary to take account of the heights

SUNSPOT MINIMUM

SUNSPOT MAXIMUM



FIG. 7

SUMMER

WINTER

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of the different layers to determine the lengths of the several hops, and also to employ separate curves of maximum usable frequency for each layer.

The waves reaching a given point may be a combination of waves having traveled by different numbers of hops. For each of these it is necessary to take account of the time of day, and to consider which layer is effective, for the locality where each reflection from the ionosphere occurs.

Optimum Frequencies.- It is found that in general (especially in the daytime) the absorption is greater, i.e., received intensities are less, as the frequency is lowered below the maximum usable frequency. Thus it requires much greater power to get satisfactory communication on frequencies very much below the maximum usable. On the other hand, it is necessary in practice to use a frequency somewhat below the values indicated in the curves of monthly average, because of the variability from day to day, which is generally within 15% of the monthly average.

Fair efficiency of communication is usually provided in the daytime by frequencies down to about 50% of the maximum usable frequencies, and at night by frequencies down to somewhat less than 50% of the maximum usable frequencies. Definite limits cannot be set because there are large irregular variations of absorption with time. At frequencies near the maximum usable frequencies there is relatively little difference between night and day absorption. As the frequency is lowered, however, the daytime absorption increases relatively much more rapidly.

It is desirable therefore to use frequencies not much below the maximum usable. A satisfactory general rule is to use a frequency between 50% and 85% of the monthly average maximum usable frequency for the given distance and time. It is not ordinarily possible to keep changing frequency continuously, so some such range of choice is necessary. Below 50%, the received waves may be too weak for use, and above 85% communication will be impossible on some days because of the variability from day to day.

Calculation of Working Frequencies. - The procedure in calculating frequencies to use over a given transmission distance for one-hop transmission is as follows. From the graph for the given distance, frequencies are selected for certain times of day as outlined just above. Each frequency may be used with reasonable satisfaction for a certain number of hours before it becomes desirable to shift to another. The times for such shifts of frequency are determined from the graph. To these times half the difference between the local times at the transmitter and receiver should be added if the transmission path lies to the west, and subtracted if the path lies to the east, of the place for which the calculations are being made. This is because reflection of the waves takes place halfway between the transmitter and the receiver. The resulting times for the shifts of frequency are in terms of time at the place for which the calculations are made.

For multi-hop transmission the same calculation should be made for each hop, and the results for the several hops combined according to the rule that the maximum usable frequency over the whole path is the least of the maximum usable frequencies over the several hops.

Consider the case of single-hop transmission between Washington, D.C., and Chicago, Ill., a distance of 1000 kilometers (620 miles). Fig. 8 shows the path taken by the radio waves. The local time at Chicago is about 40 minutes earlier than at Washington. Therefore the local time at the reflection point (midpoint of the transmission path) is 20 minutes earlier than at Washington. From the graphs (Fig.7) for 1000 km it may be seen that a possible set of transmission frequencies, between 50 and 85% of the maximum usable frequency, are those given in table 2.

Table 2

Month and year	ith and Frequency, year kc/s		Local time at midpoint of path		n rd	Layer
		from	to	from	to	
June 1933	300C 8000	2100 0700	0700 2100	2120 0720	0720 2120	one-hop F one-hop E
Dec. 1933	3000 5000 8000 5000	1900 0730 0930 1530	0730 0930 1530 1900	1920 0750 0950 155 0	0750 0950 1550 1920	one-hop F one-hop F ₂ one-hop F2 one-hop F
June 1937	6000 11000	2100 C700	C700 2100	2120 0720	0720 2120	one-hop F one+hop F ₂ one-hop E
D €0. 1937	450 0 10000 16000 10000	1900 0800 0930 1530	080C 093C 1530 19CC	1920 0820 0950 1550	0820 0950 1550 1920	one-hop F one-hop F ₂ one-hop F ₂ one-hop F

Frequencies for Dependable Transmission over a Distance of 1000 Kilometers.



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Consider another case: transmission Between Washington, D.C., and San Francisco, California, a distance of 4000 kilometers (2500 miles). Such transmission is not ordinarily possible by one hop. It is a case of two-hop or three-hop transmission. Table 3 gives data similar to table 2, calculated for this transmission path. At times (e.g., daytime in summer during the sunspot minimum) the maximum usable frequency over this path is determined by E-layer transmission. At other times it is determined by two-hop F- or F-layer transmission or by a combination of E- and Fo-layer trafismission. In determining the frequencies of Table 3, calculations are made for each hop separately in terms of local time at the midpoint of the hop and these local times are reduced to Eastern Standard Time. The frequency given in the table for any particular Eastern Standard Time is the lowest of the frequencies selected for each of the hops at the given time. The underlining of certain times indicates which hop determines the maximum usable frequency for the entire path.

Month and year	Freq., kc/s	Local time at midpoint of path		EST for eastern hop		EST for ES Western en hop		EST for EST for western entire hop path		for re 1	Layer
1		from	to	from	to	from	<u>to</u>	from	to		
June 1933	4500 10000	2100 0700	0700 2100	2145 0745	0745 2145	2315 0915	<u>0915</u> 2315	2145	0915 2145	two-hop F three-hop E	
Dec. 1933	4000 7000 13000 7000	1900 0730 0930 1530	0730 0930 1530 1900	<u>1945</u> 0815 1015 1615	0815 1015 1615 1945	2115 0945 1145 1745	0945 1145 1745 2115	1945 0945 1145 1615	0945 1145 1615 1945	two-hop F two-hop F2 two-hop F2 two-hop F	
June 1937	9000 15000	2100 0800	0800 2100	2145 0845	0845 2145	2315 1015	1015 2315	2145	1015 2145	two-hop F two-hop F2 or three- hop E	
Dec. 1937	6000 13000 27000 13000	1900 0800 0930 1530	0800 0930 1530 1900.	<u>1945</u> 0845 1015 1615	0845 1015 1615 1945	2115 1015 1145 1745	1015 1145 1745 2115	1945 1015 1145 1615	1015 1145 1615 1945	two-hop F two-hop F2 two-hop F2 two-hop F	

Table 3.

Frequencies for Dependable Transmission over a Distance of 4000 Kilometers.

The optimum frequencies were much higher in 1937, the time of sunspot maximum, than in 1933, the time of sunspot minimum. The next sunspot minimum will be about 1944, and it is likely that the optimum frequencies will decrease till then, reaching values somewhere near those of 1933.

EFFECTS OF IONOSPHERE IRFEGULARITIES.

The primary effects of the ionosphere on radio wave propagation are those already described, which are due to the normal or regular characteristics of the ionosphere. The modes of variation of those characteristics have been shown to be of a regular and fairly predictable nature. There are some other ionosphere phenomena which are irregular in their times of occurrence, and make radio phenomena in general much less predictable. Five types of such phenomena have been identified: sporadic-E-layer reflections, scattered reflections, sudden ionosphere disturbances, prolonged periods of lowlayer absorption, and ionosphere storms. While all five are irregular in time, the first two are primarily due to irregularities in space.

Detailed information on the third, fourth, and fifth of these are included in the Bureau's weekly broadcast and monthly report on the ionosphere. The monthly report includes also information on the prevalence of sporadic E in the months when it is most prevalent.

Sporadic E .- It sometimes happens that waves are reflected by the E layer on frequencies higher than that at which the E-layer waves normally disappear and the reflection of waves by higher layers begins. Thus, in the example shown in Fig. 3 waves may sometimes be reflected at the E-layer height of 110 kilometers on frequencies of 4 or 5 or more megacycles. These reflections are due to a process different from the normal reflection in the ionized layer; the process is probably one of reflection from a sharp boundary of stratified ionization. The existence of these "sporadic E" reflections necessitates a redefinition of the term "critical frequency", previously defined as the highest frequency at which waves sent vertically upward are received back from the layer. When sporadic E reflections occur they may be received simultaneously with reflections from higher layers; thus, e.g. in the case shown in Fig. 3, vertical-incidence reflections might be received at 8 Mc/s from both the E and the Fo layers. The E-layer critical frequency, more precisely defined, is the value (3300 kc/s in the example shown in Fig. 3) at which the observed virtual height shows a sudden rise to large values as the frequency is increased. Except for the occasional occurrence of sporadic E reflections, all waves of higher frequency pass through the E layer and are not reflected by it.

Sporadic E leads to interesting results in long-distance radio transmission. It accounts for long-distance transmission up to higher frequencies than by any other means. Strong vertical-incidence reflections by sporadic E sometimes occur at frequencies up to about 12 Mc/s. By reason of the large angles of incidence possible with the E layer, this makes occasional long-distance communication possible on frequencies as high as 60 Mc/s. Such communication is generally for only a short time and for restricted localities. Sporadic E is thus patchy or sporadic in both geographical distribution and time.

Sporadic E occurs most commonly in the summer, May to August, particularly in the morning and evening but may occur any time of day or night. It occurs occasionally at all seasons, particularly in the evening. It occurs more at high latitudes than in equatorial regions.

Scattered Reflections .- An irregular type of reflection from the ionosphere occurs at all seasons and is prevalent both day and night. These reflections are most noticeable within the skip zone, or at frequencies higher than those nominally receivable from the regular layers. Like sporadic E, they occur at frequencies which may exceed the Fo critical frequencies, but are unlike sporadic E in that they are complex and jumpy thus causing signal distortion; they occur and disappear fitfully, and are almost useless for communication purposes. Some types are of very weak intensity. The scattered reflections are characterized by very great virtual heights, usually somewhere from 400 to 1500 kilometers. Their occurrence was for a time thought to indicate the existence of another layer above the F₂ layer which might be called the G layer. It is now however thought that they are of several types, and that some of them are due to complex reflections from small, epheneral, scattered patches or "clouds" of ionization in or between the normal ionosphere layers, and thence to one or more layers and ground by single or multiple reflection.

Sudden Ionosphere Disturbances.- The most startling of all the irregularities of the ionosphere and of radio wave transmission is the sudden type of disturbance manifested by a radio fadeout. This phenomenon is the result of a burst of ionizing radiation from a bright chromospheric eruption on the sun, causing a sudden abnormal increase in the ionization in the D layer (below the E layer), frequently with resultant disturbances in terrestrial magnetism and earth currents as well as radio transmission. The radio effect is the sudden cessation of radio transmission on frequencies over about 1500 kc/s. The drop of the radio signals to zero usually occurs within a minute. The effects occur simultaneously throughout the hemisphere illuminated by the sun, and do not occur at night. The effects last from about ten minutes to an hour or more, the occurrences of greater intensity in general producing effects of longer duration. The effects are more intense, and last longer, the lower the frequency in the highfrequency range (i.e., from about 1500 kc/s up). It is consequently sometimes possible to continue communication during a radio fadeout by raising the working frequency.

The radio, magnetic, and other effects are markedly different from other types of changes in these quantities. The effects are most intense in that region of the earth where the sun's radiation is perpendicular, i.e., greater at noon than at other times of day and greater in equatorial than in higher latitudes.

Taking due account of the variation of the effects with frequency and distance, varying effects in differing directions can be explained. Reception in the United States from stations in the southern hemisphere usually exhibits greater effects than reception from other directions (because of passing the equatorial regions). Similarly, when the disturbance occurs at a time when it is morning at the receiving point the effects are usually greater in reception from the east than from the west, and vice versa for the afternoon (because of passing the region where it is noon). A radio fadeout sometimes occurs when it is night at the receiving point, but only when the path of the waves is somewhere in daylight.

<u>Prolonged Periods of Low-Layer Absorption.</u> This phenomenon is similar to the sudden ionosphere disturbance in its effects and characteristics except that its beginning as well as recovery is gradual and it has a longer time duration, commonly several hours. The intensity diminution is in general not as severe as in the more intense fadeouts, but sometimes the intensities fall to zero.

The low-layer absorption effect appears to be due to increased ionization in the D layer (below the E layer), exactly as for the sudden ionosphere disturbances. The increased ionization is caused by an abnormally great outpouring of ultraviolet light from the sun, but in this case it is not so sudden as in the eruptions which cause the sudden ionosphere disturbances. The variation of the effects with frequency, and other characteristics, are the same as for the sudden ionosphere disturbances. Both phenomena occur at all seasons, but the prolonged periods of low-layer absorption have been found to occur in a group of several weeks duration at periods of high sunspot activity, the groups being separated by more or less quiet periods of several months. They frequently but not always occur during periods when sudden ionosphere disturbances are numerous.

Ionosphere Storms .- An ionosphere storm is a period of disturbance in the ionosphere in which there are great anomalies of critical frequencies, virtual heights, and absorption. Radio transmission is poor (except for the low frequencies, below 500 kc/s, which are sometimes improved). An ionosphere storm usually lasts one or two days, and occurs both day and night. It is usually accompanied by a magnetic storm (i.e., a period of unusual fluctuation of terrestrial magnetic intensity). During the first few hours of very severe ionosphere storms the ionosphere is turbulent, stratification is destroyed, and radio wave propagation er-During the later stages of very severe storms and ratic. during the whole of more moderate storms, the upper part of the ionosphere, principally the Fo layer, is expanded and diffused. The critical frequencies are much lower than normal and the virtual heights much greater, and therefore the maximum usable frequencies are much lower than normal. It is often necessary to lower the working frequency in order to maintain communication during one of these storms. There is also increased absorption of radio waves during an ionosphere storm. Ionosphere storms are most severe in auroral latitudes and decrease in intensity as the equator is approached. Ionosphere storms occur approximately simultaneously over wide geographical areas. The condition of the ionosphere is much less uniform from point to point than on undisturbed days.

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APPENDIX 1

Recording Form for Weekly Broadcast from Station WWV							
of lonosphere .	Data of Nation	al Bureau of t	tandar	Nashing	ton, D.C.		
			Date		19		
1.	A. <u>NOON</u> Critical f	TODAY requency (kc)		Virtu (kile	al height ometers)		
E layer							
F _l layer		and a second					
F ₂ layer	6						
Sporadic E reflections	ap to	• • • • • • • .					
2. Limits of variation of	critical freq	uencies from r	loon va	lues:			
E layer:% below at	about	A.M. and	P•M•	local time	e		
F ₂ layer:% below at	A.M.			P.M., 1	ocal time.		
3. Maximum usable frequen	cies (and)	layer	transmiss	ion.		
Distance E layer F _l lay (km) (kc) (kc	er F ₂ layer) (kc)	Distance E (lim)	layor (kc)	F _l layer (kc)	F ₂ layer (hc)		
200	Ballined Production of the advant	1600					
400		2000					
700		2500					
1000		35°0					
1300							
4. F. MIDNIGHT LAST NIGHT							
	Critical f	requency (kc)		Virtua (kilon	l height noters)		
F layer	b						
Sporadic E reflections up to							
5. Limits of variation of critical frequency from midnight value:							
F layer:% above atP.M. to% below atA.M. local time.							

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6. Maximum usable frequencies, F layer transmission.

Distance Flayer (lon) (kc)	Distance (lum)	F layer (lcc)	Distanco (lzm)	F layor (lzc)			
400	1300		2500				
700	1600	0	3500				
1000	2000						
7. C. <u>101</u>	NOSPHERE IRREG	ULARITIES D	URING PAST W	EEK .			
ionosphere storm	ns wo re o bserv	red					
Time, EST Date	e Time	, EST D	ate	Intensity			
From	to			4 0-100-0-100-00-00-00-00-00-00-00-00-00-0			
From	to						
From	to			See align an international management of a state of			
sudden fanösphere disturbances were observed							
Date Time, GMT	Date	Time, GMT	Date	Time, GMT			
at	at		at	₩~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~			
at	atat						
at	atat						
Remarks:							

Notes: (Not broadcast):-

(a) All critical frequencies are given in terms of the "ordinary wave", not the "extraordinary wave".

(b) Maximum usable frequencies are not given for E layer transmission when they are less, for all distances, than those for F₂ layer transmission.

(c) Each distance given in the tables of maximum usable frequencies is the ship distance for the corresponding frequency, if the frequency is given for only one layer, or for the higher corresponding frequency, if frequencies are given for two layers.

(d) For ionospherically quiet days, the critical frequencies, heights, and maximum usable frequencies do not in general vary greatly from day to day. On ionospherically disturbed days the values may differ materially.

(e) The data as broadcast are sometimes revised or corrected after further study and comparison of records. Corrected data are given in the monthly ionosphere reports in Proc.I.R.E.

(f) For further information on these broadcasts and the uses which may be made of the data, see the Bureau's Letter Circular on this subject, obtainable by mail upon request.

APPENDIX 2

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